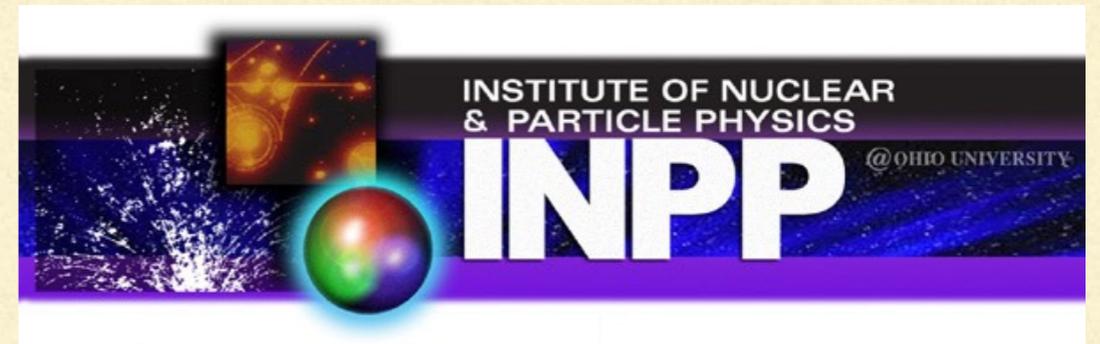


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# Knowing What You Don't Know: Nuclear Reactions, Effective Field Theory & Uncertainty Quantification

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Daniel Phillips  
Ohio University  
TU Darmstadt  
EMMI

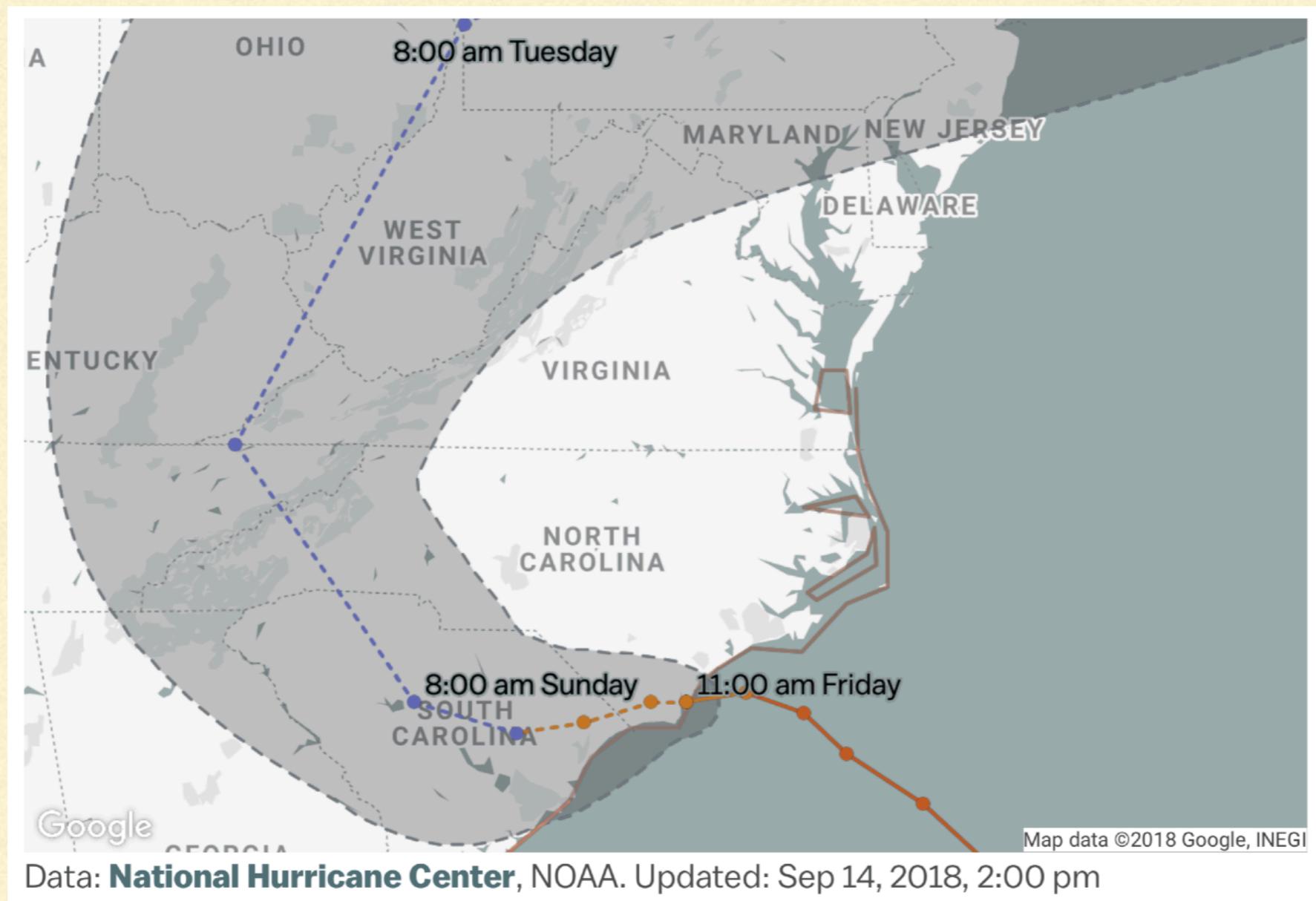


OHIO  
UNIVERSITY

RESEARCH SUPPORTED BY THE US DOE AND BY EMMI

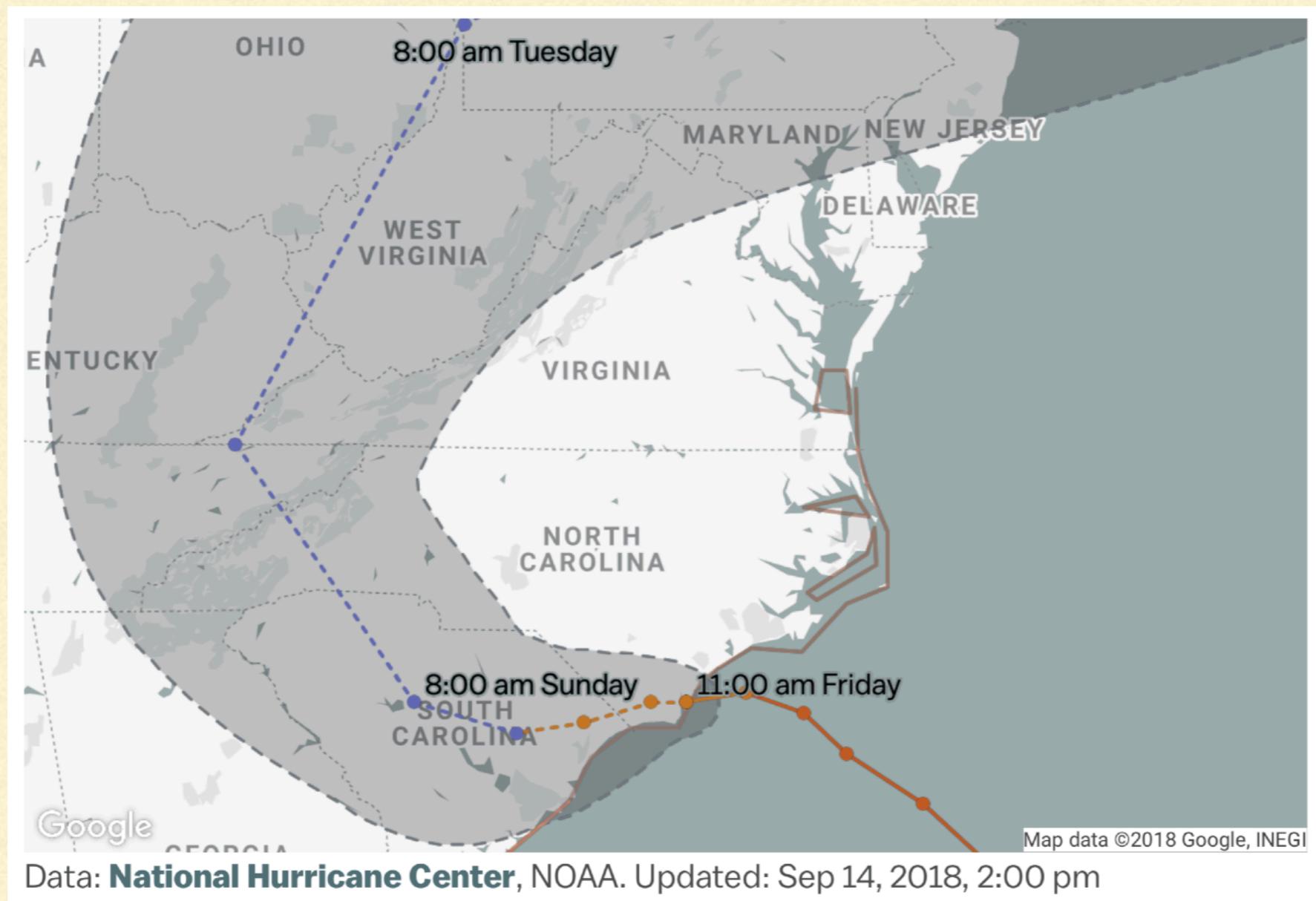
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# Hurricane forecasting



<http://www.vox.com>

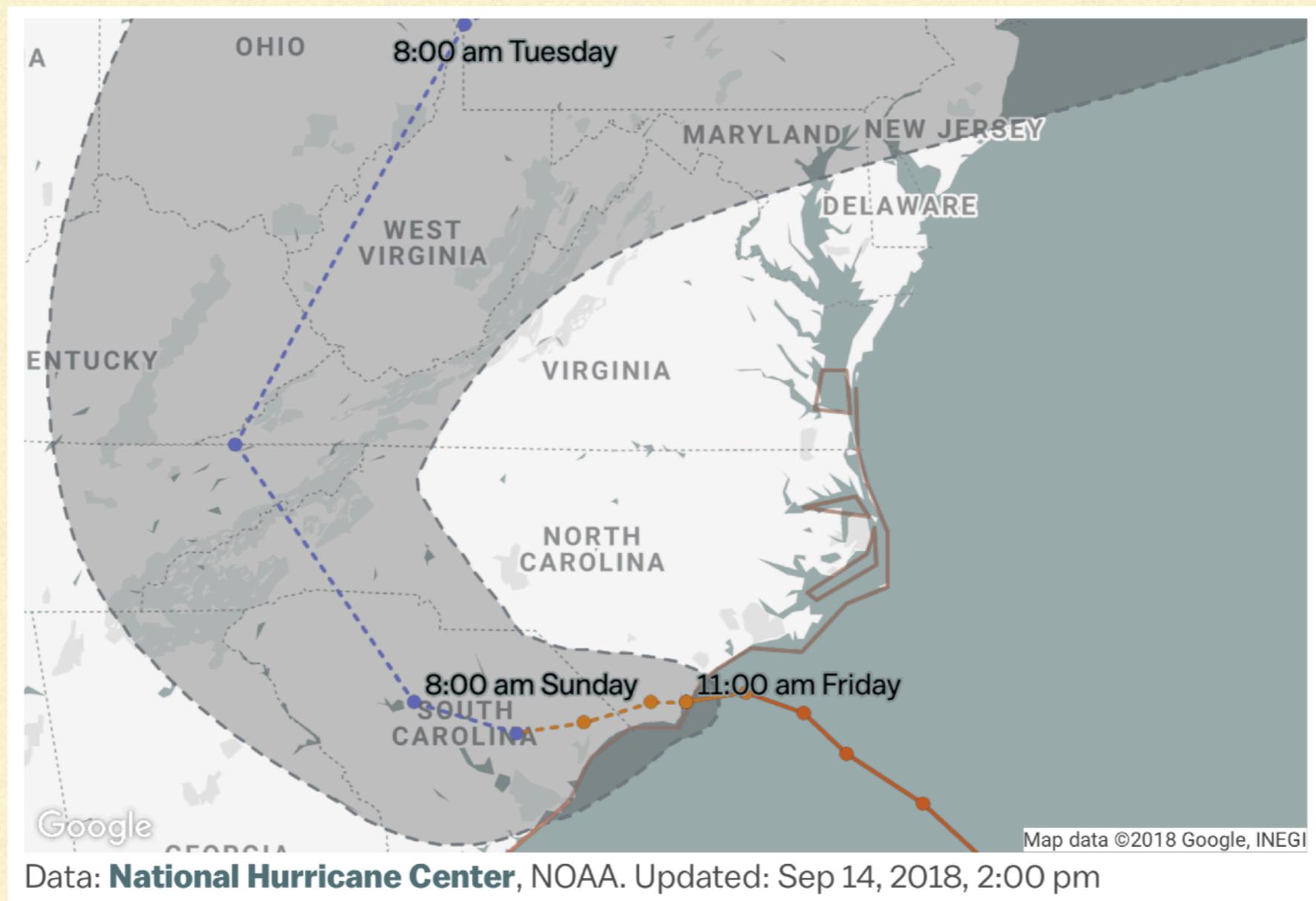
# Hurricane forecasting



<http://www.vox.com>

- Forces, e.g., Coriolis
- Conservation laws
- Parameterizations

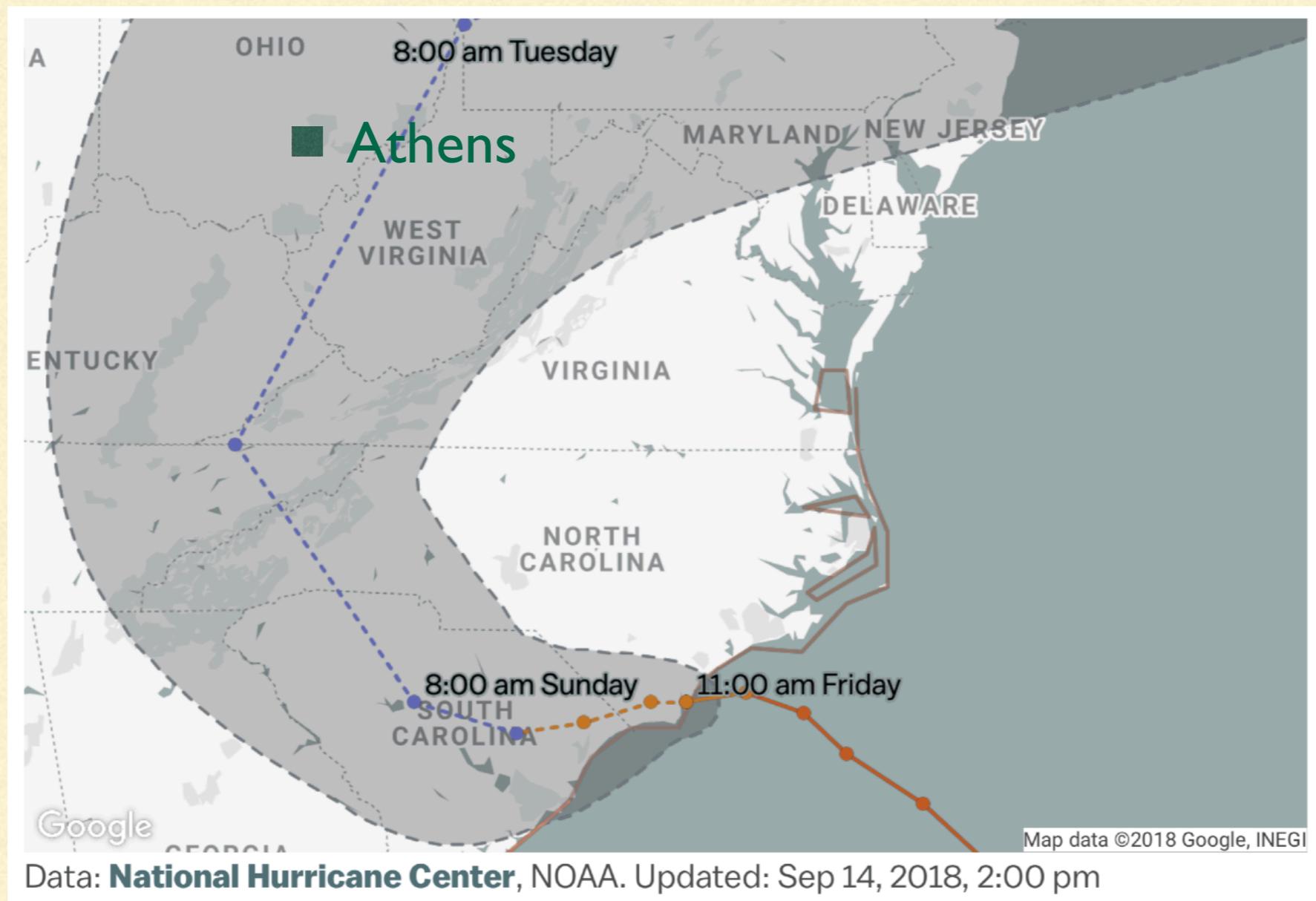
# Hurricane forecasting



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- Need to know initial state accurately (computing!)
- Evolve state forward in time (more computing!)
- Uncertainty quantification

# Hurricane forecasting



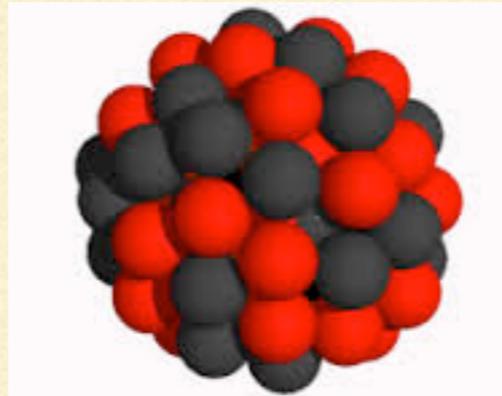
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# Nuclear reactions

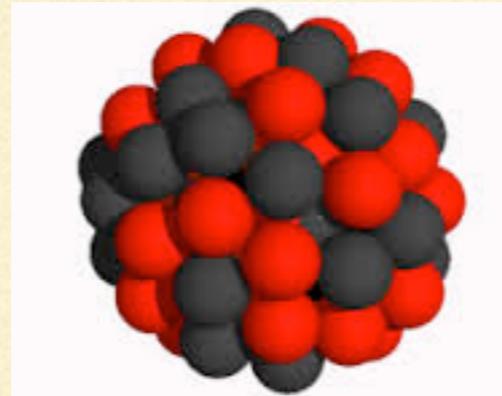
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$$i\hbar \frac{\partial |\Psi\rangle}{\partial t} = (\hat{T} + \hat{V}) |\Psi\rangle$$

# Nuclear reactions

---



$$i\hbar \frac{\partial |\Psi\rangle}{\partial t} = (\hat{T} + \hat{V}) |\Psi\rangle$$

- Forces: electromagnetic, strong nuclear
  - Conservation laws, e.g., probability, energy, momentum
  - Some parameterizations
  - Accurate knowledge of initial state (nuclear structure)
  - Computing to evolve state forward in time
  - Uncertainty quantification
-

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# Outline

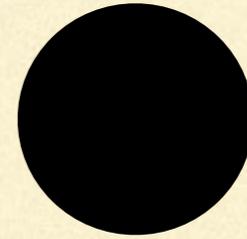
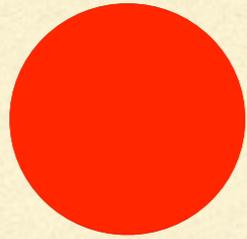
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- What we do and don't know about the strong nuclear force
  - EFT: organizing what we know, constraining what we don't
  - EFT truncation errors from a Bayesian analysis: NN scattering
  - EFT for halo nuclei: universal formula for  $\gamma + {}^A_Z \rightarrow {}^{A-1}_Z + n$
  - Uncertainty quantification for fusion:  ${}^7\text{Be}(p,\gamma)$  at solar energies
  - Conclusion
-

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# Potentials from particle exchange

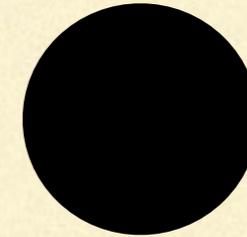
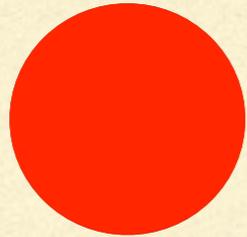
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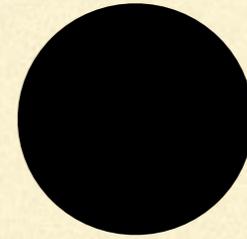
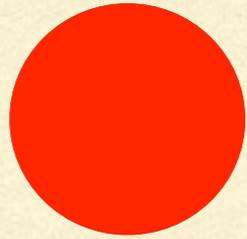


- At the sub-atomic level, forces generated by exchange of particles

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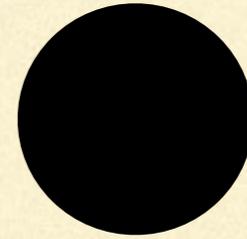
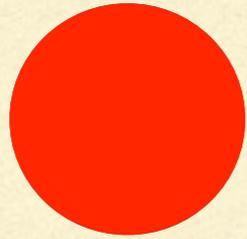


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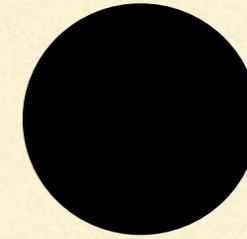
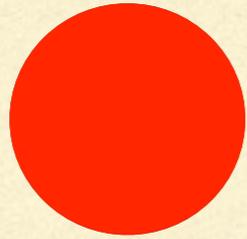


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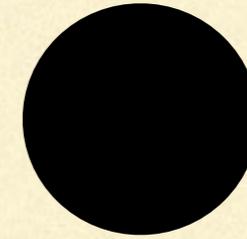
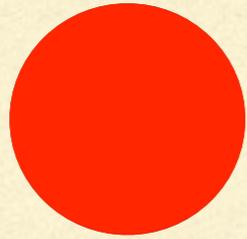


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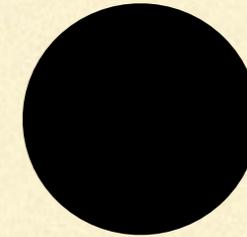
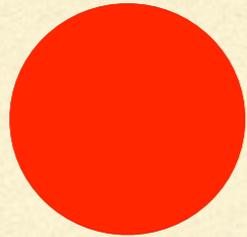


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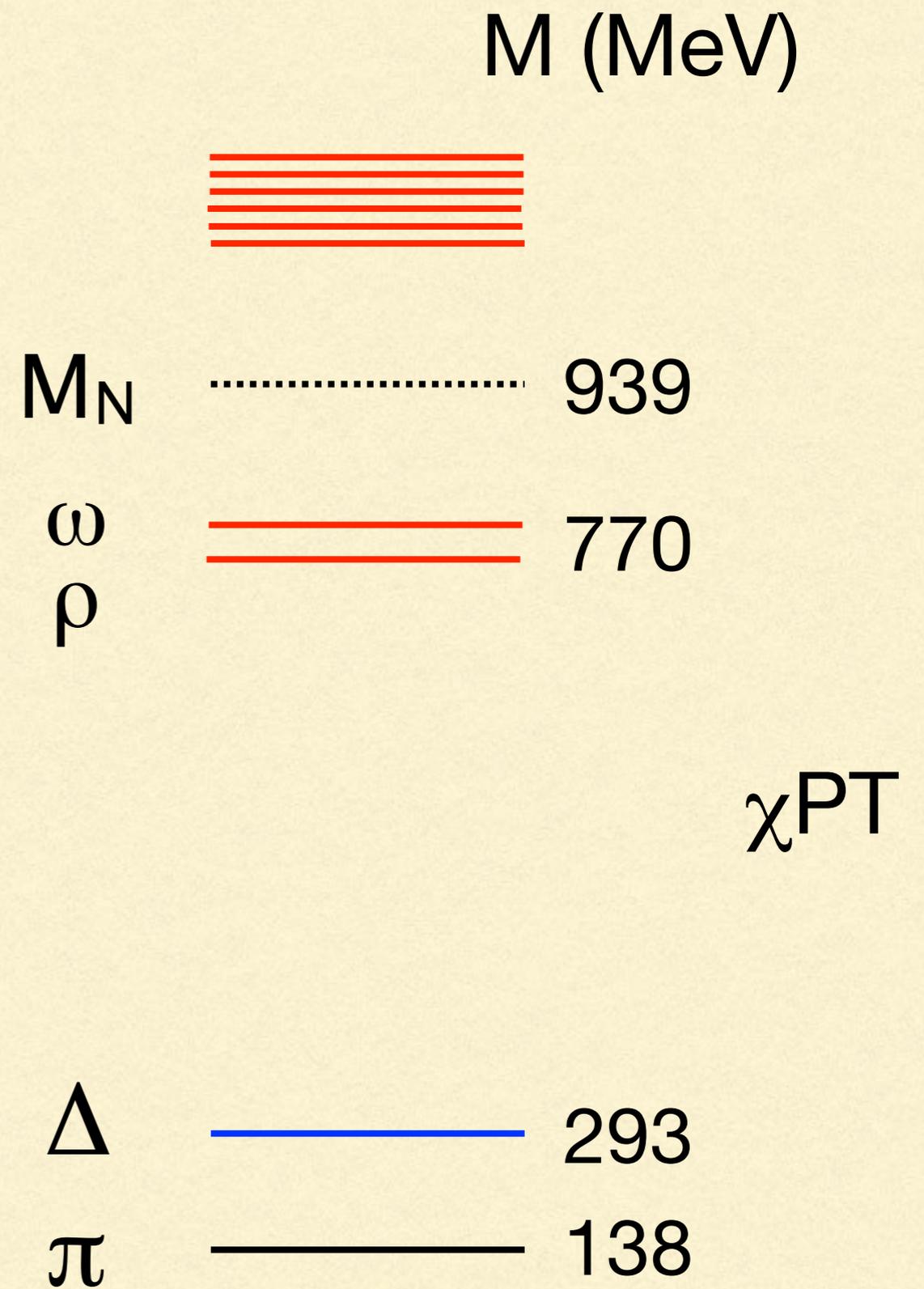
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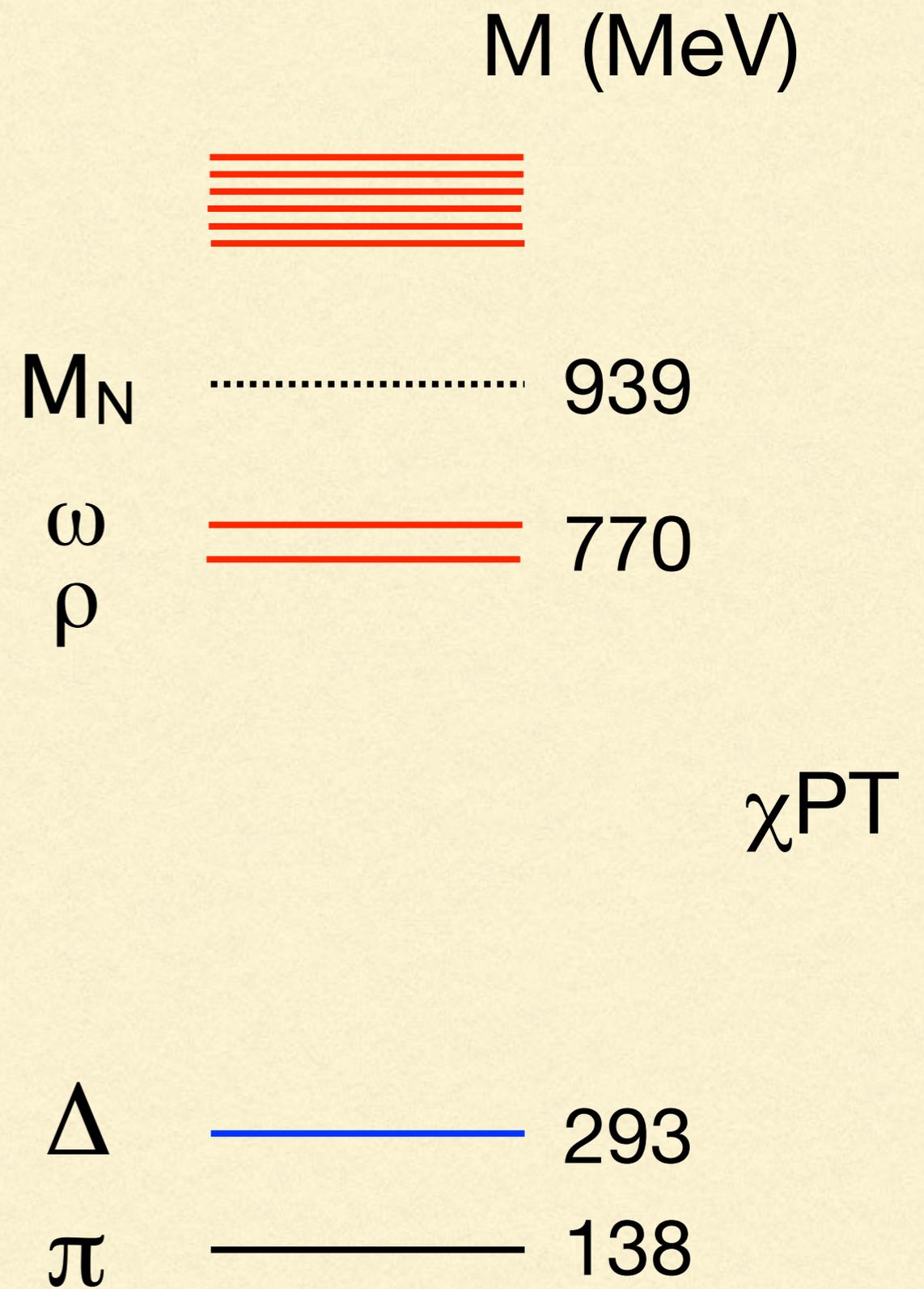
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  - Longest range forces generated by lightest particles
-

# The long and the short of hadron physics



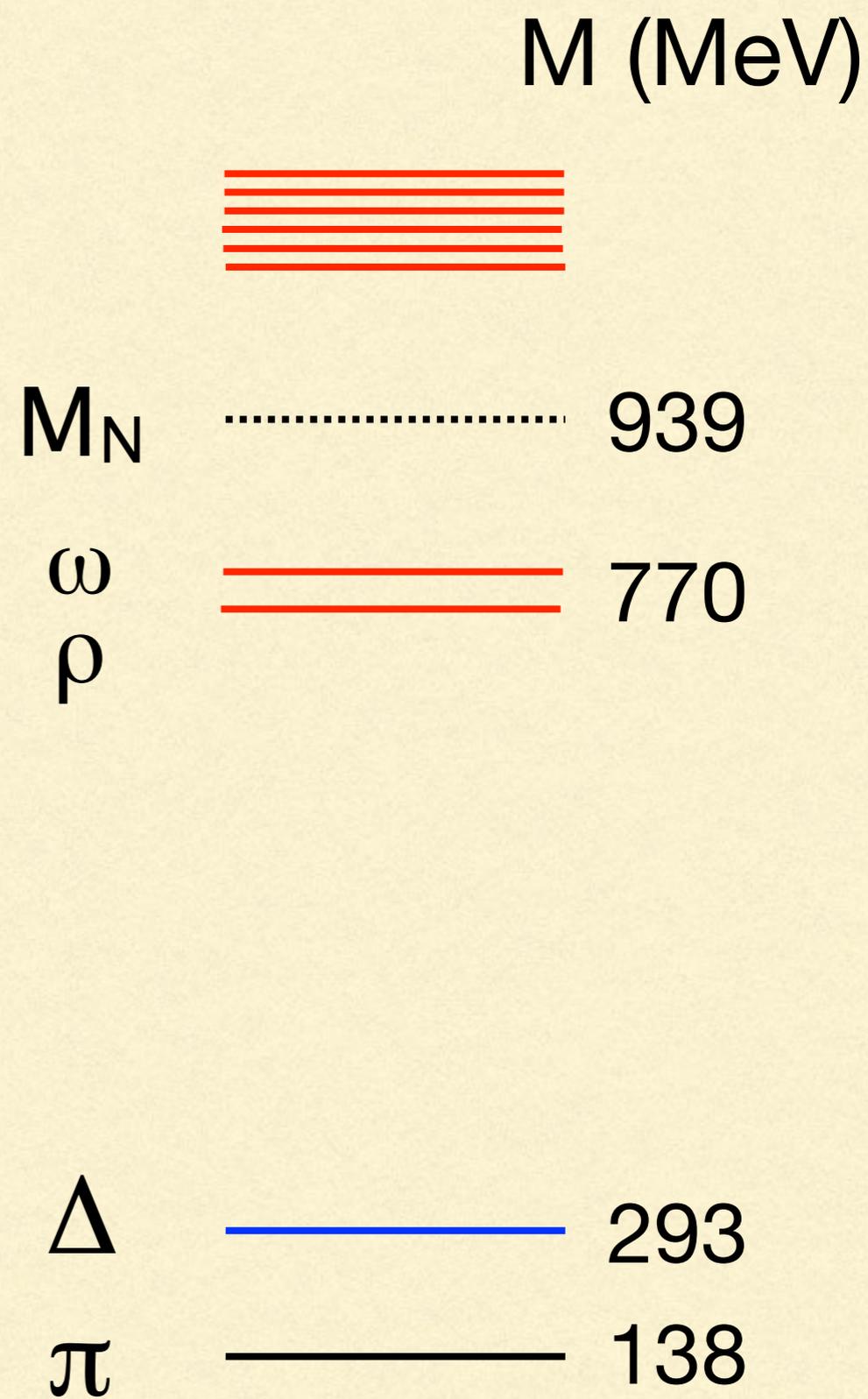
# The long and the short of hadron physics

- Spectrum of QCD bound states



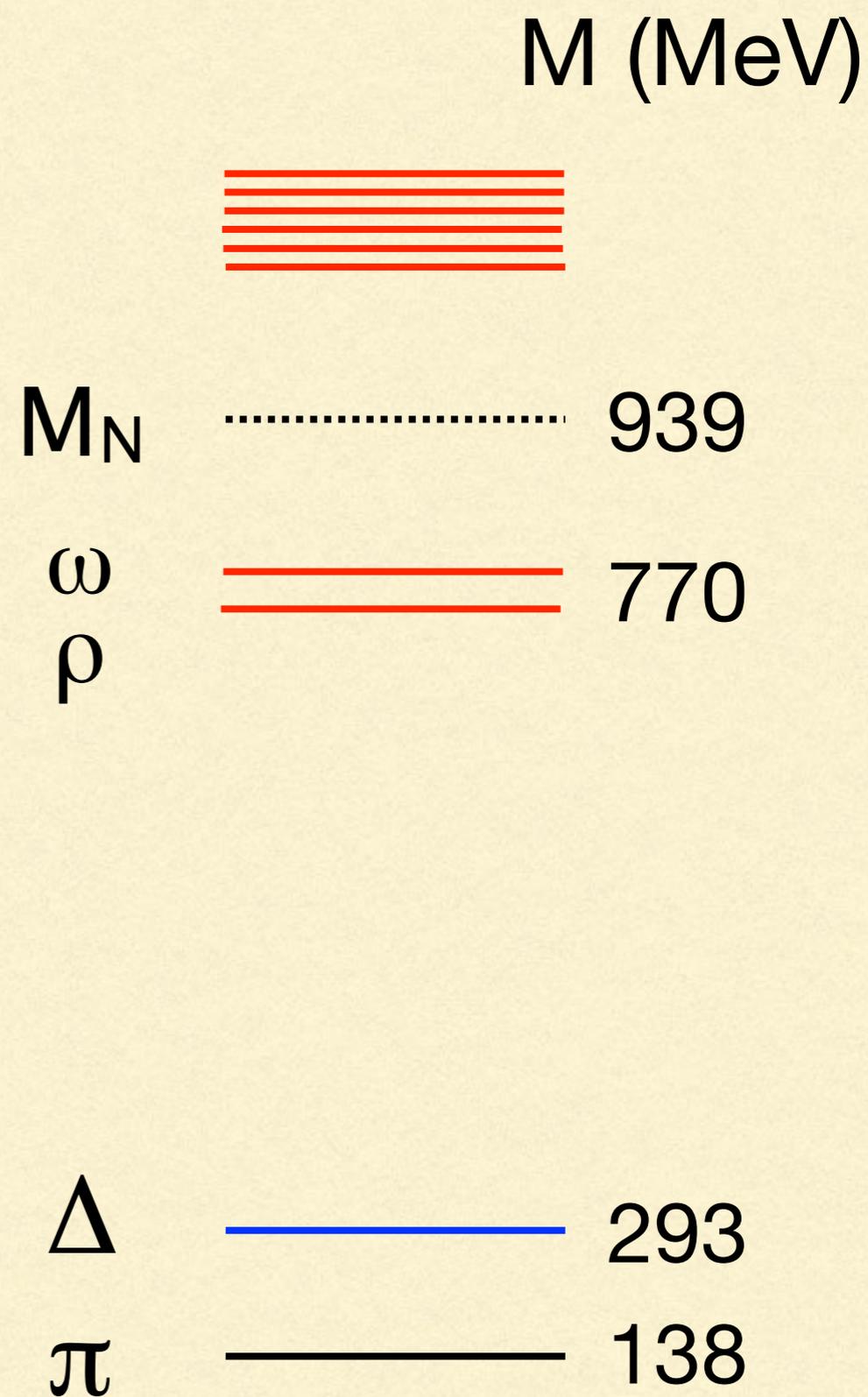
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- For probe energies  $\sim$ a hundred MeV, simplifications of the rich QCD dynamics emerge: processes dominated by  $\pi$ s (and  $\Delta$ s)

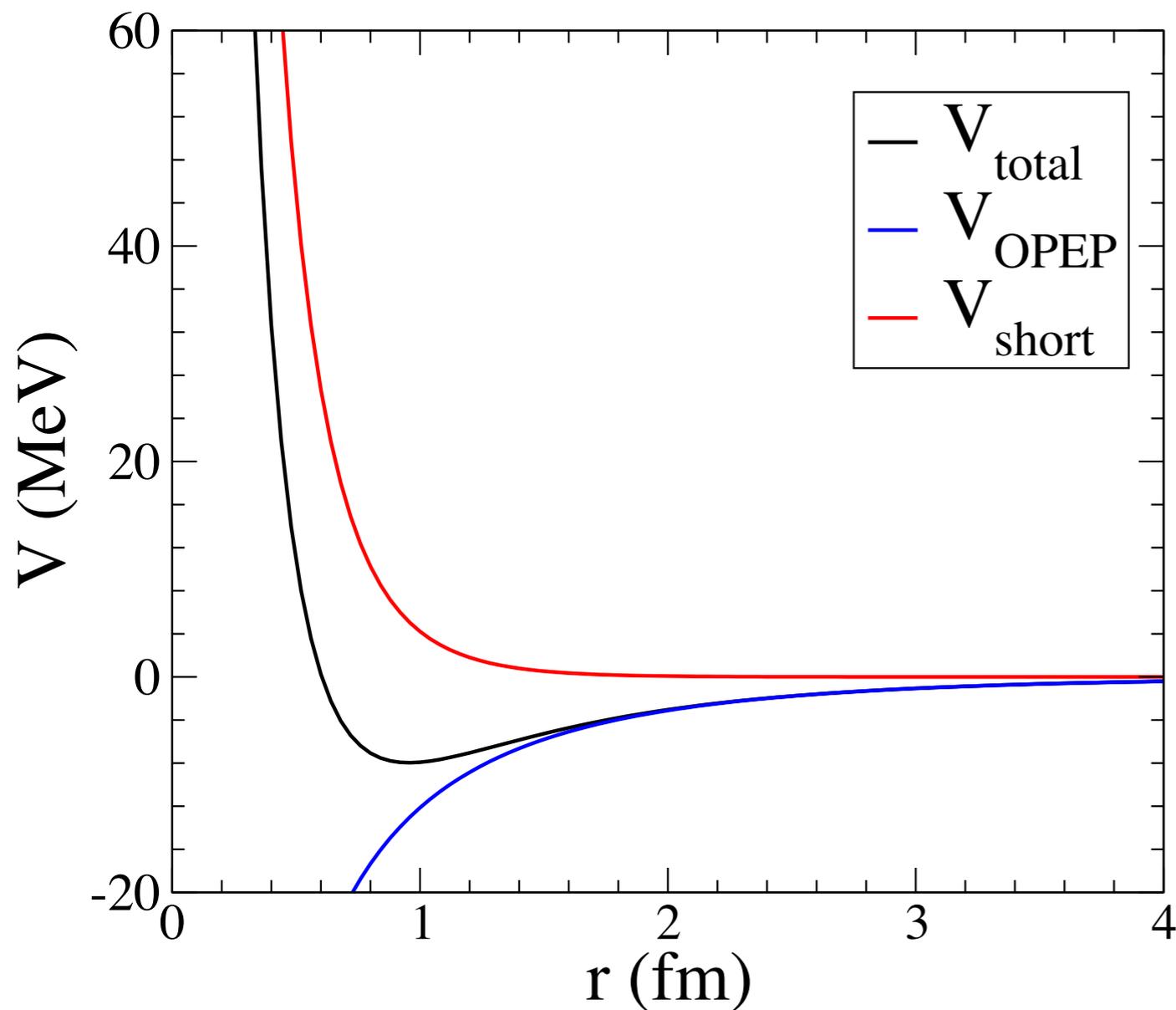


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- For probe energies  $\sim$ a hundred MeV, simplifications of the rich QCD dynamics emerge: processes dominated by  $\pi$ s (and  $\Delta$ s)
- Pion exchange generates longest-range part of NN force
- But short-distance dynamics too



# The NN potential: a cartoon



- Long-range part generated by one-pion exchange
- Intermediate ranges: multiple pion exchange
- Short ranges: “other stuff” exchange
- Needs to be parameterized, then fit to NN scattering data

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# Effective Field Theory

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- Simpler theory that reproduces results of full theory at long distances
  - Short-distance details irrelevant for long-distance (low-momentum) physics, e.g. multipole expansion
  - Expansion in ratio of physical scales:  $p/\Lambda_b = \lambda_b/r$
  - Symmetries of underlying theory limit possibilities: all possible terms up to a given order present in EFT
  - Short distances: unknown coefficients at a given order in the expansion need to be determined. Symmetry relates their impact on different processes
  - Examples: standard model, chiral perturbation theory, Halo EFT
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Monet (1881)



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**Error grows as first omitted term in expansion**

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# $\chi$ EFT for nuclear forces

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- $\chi$ PT  $\Rightarrow$  pion interactions are weak at low energy.

Weinberg (1990), apply  $\chi$ PT to  $V$ , i.e. expand it in  $P=p/\Lambda_b$

$$(E - H_0)|\psi\rangle = V|\psi\rangle$$

$$V = V^{(0)} + V^{(2)} + V^{(3)} + \dots$$

Ordóñez, Ray, van Kolck (1996); Epelbaum, Meissner, Gloeckle (1999); Entem, Machleidt (2001)

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- Leading-order  $V$ :

$$V^{(0)} = \text{[diagram: contact vertex]} + \text{[diagram: pion exchange]} ;$$

$$\langle \mathbf{p}' | V | \mathbf{p} \rangle = C^{3S1} P_{3S1} + C^{1S0} P_{1S0} + V_{1\pi}(\mathbf{p}' - \mathbf{p})$$

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	Two-nucleon force	Three-nucleon force	Four-nucleon force
$P^0$		—	—
$P^2$		<b>CONSISTENT 3NFS, 4NFS</b> —	—
$P^3$			—
$P^4$			

work in progress...

Figure courtesy  
E. Epelbaum

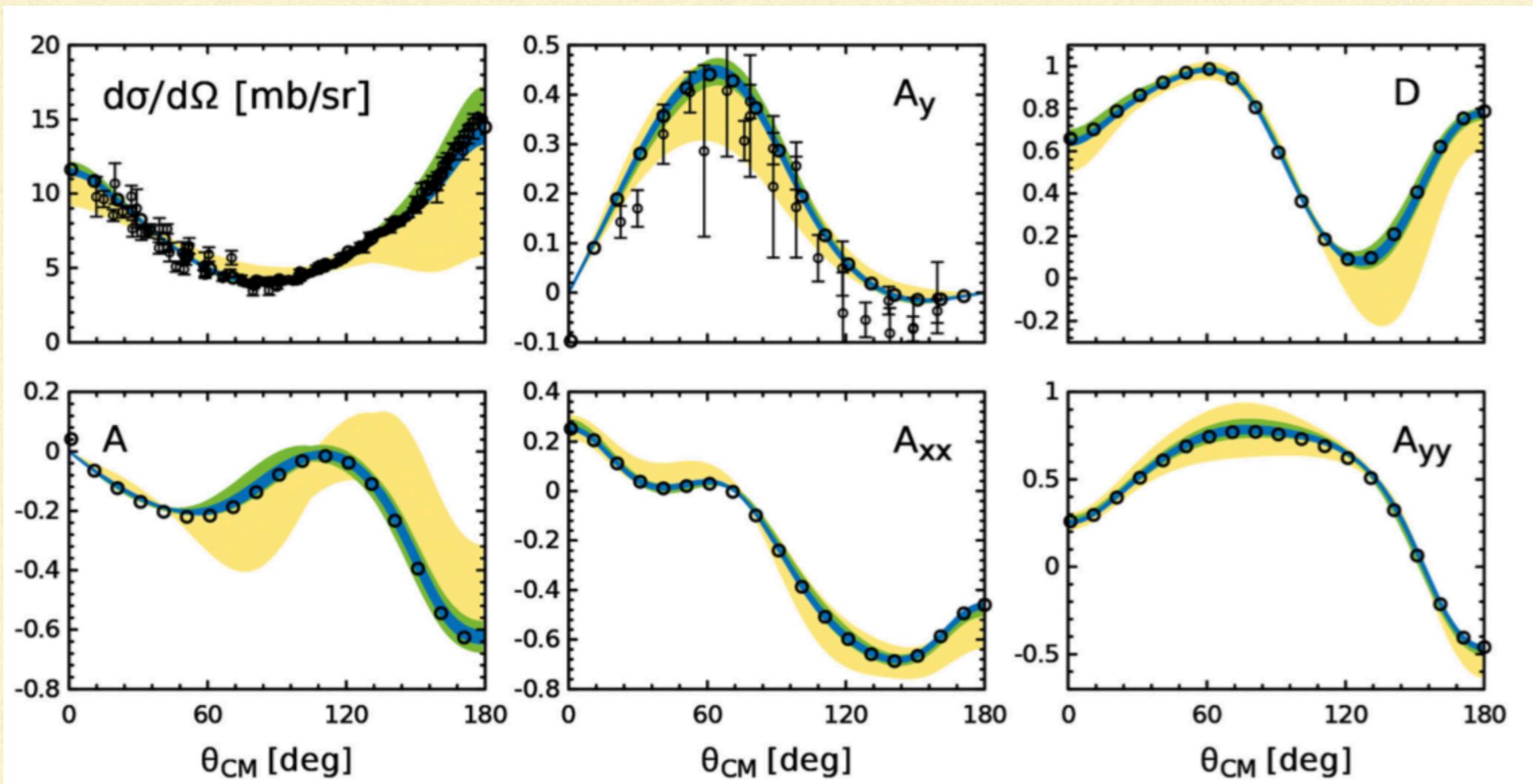
# NN scattering

Epelbaum, Krebs, Meissner, PRL (2015); EPJA (2015)

- Potential regulated by local function, parameterized by R

$$\sigma_{np}(E_{\text{lab}}) = \sigma_{\text{LO}} \sum_{n=0}^k c_n(p_{\text{rel}}) \left( \frac{p_{\text{rel}}}{\Lambda_b} \right)^n$$

EKM state  
 $\Lambda_b = 600 \text{ MeV}$



R=0.9 fm  
 here

np observables at  $E_{\text{lab}} = 96 \text{ MeV}$  at NLO,  $N^2\text{LO}$ ,  $N^3\text{LO}$  ( $k=2, 3, 4$ )

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# Successes in $A=3-12$

Epelbaum et al. (LENPIC), arXiv:1807.02848

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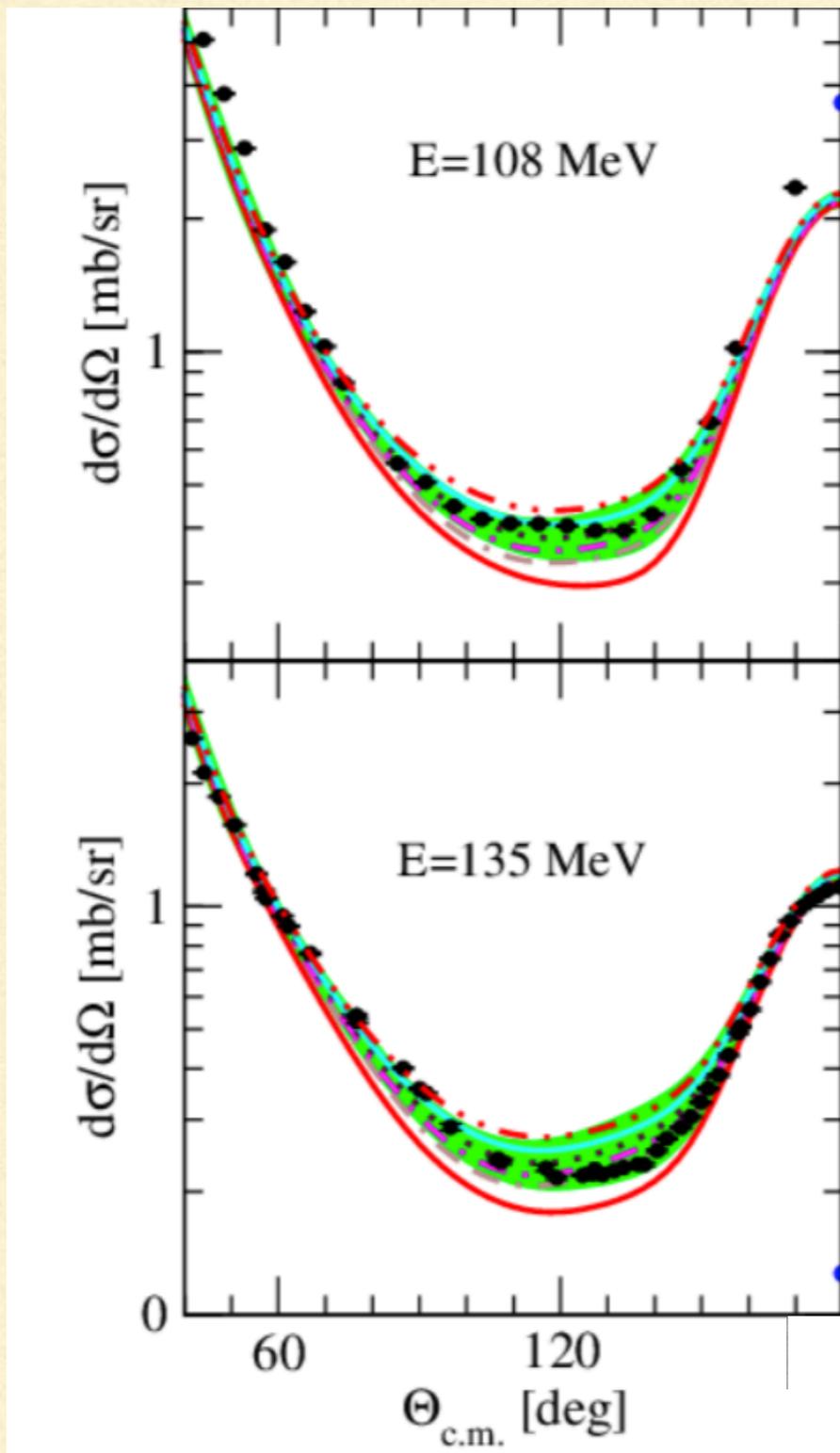
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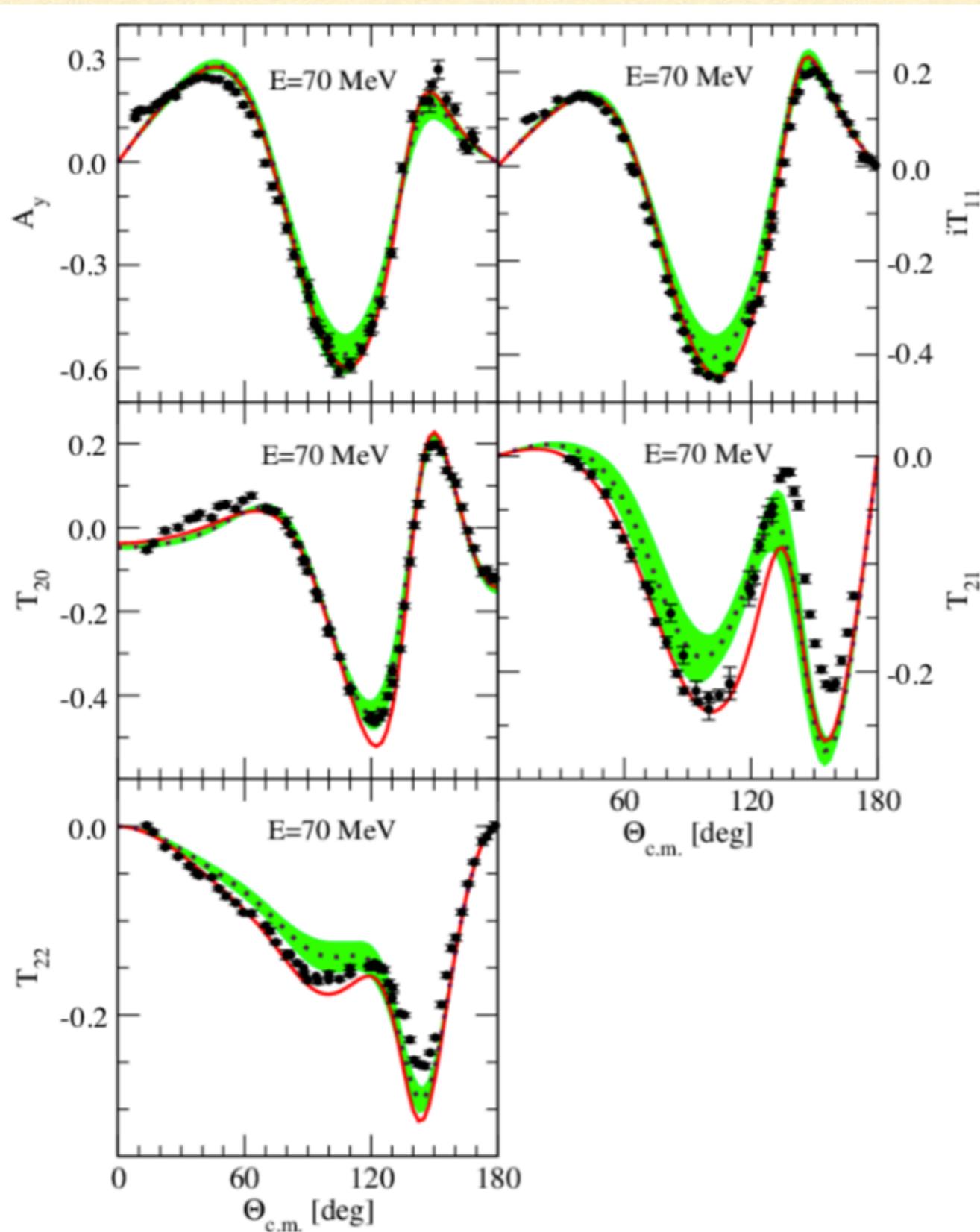
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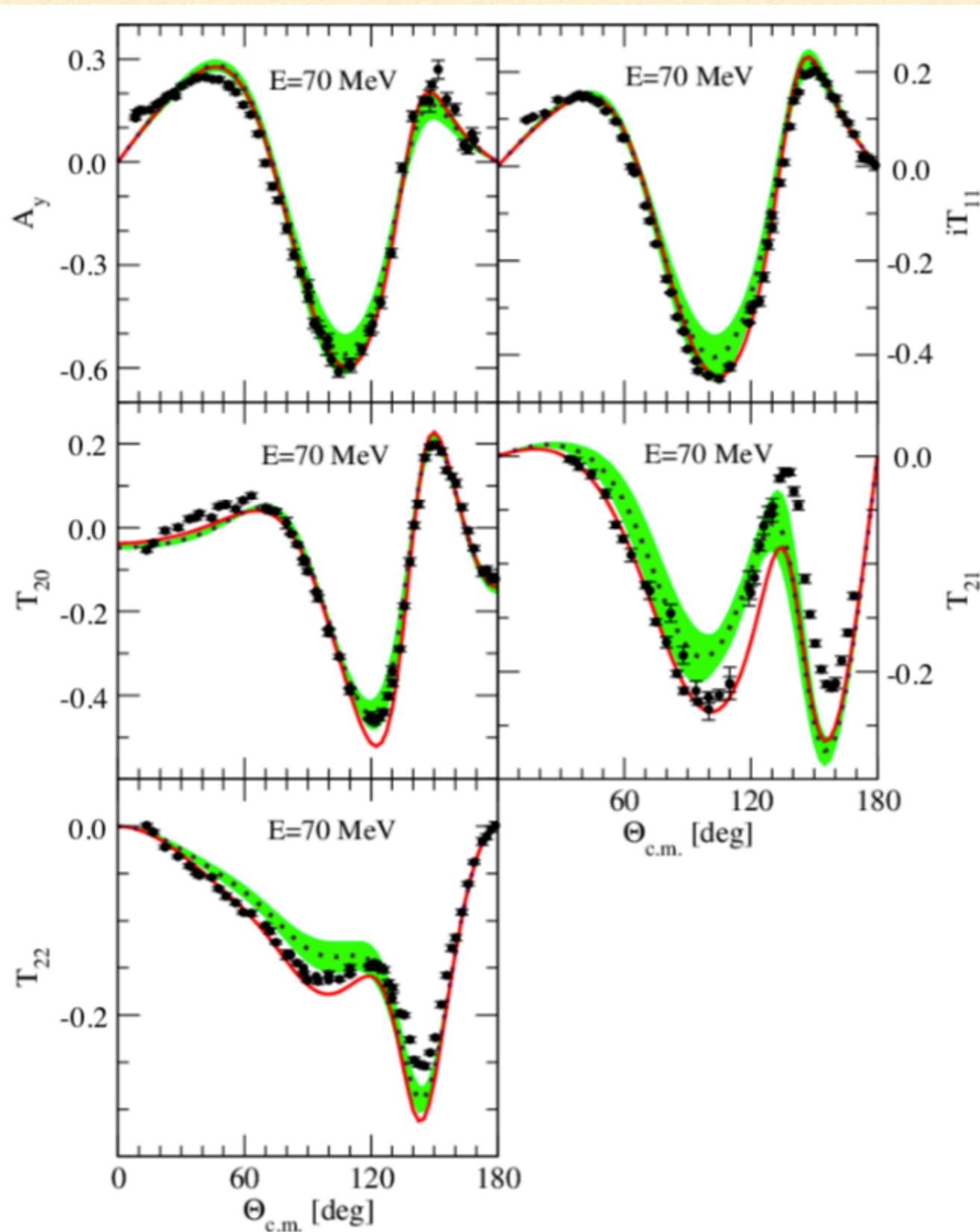
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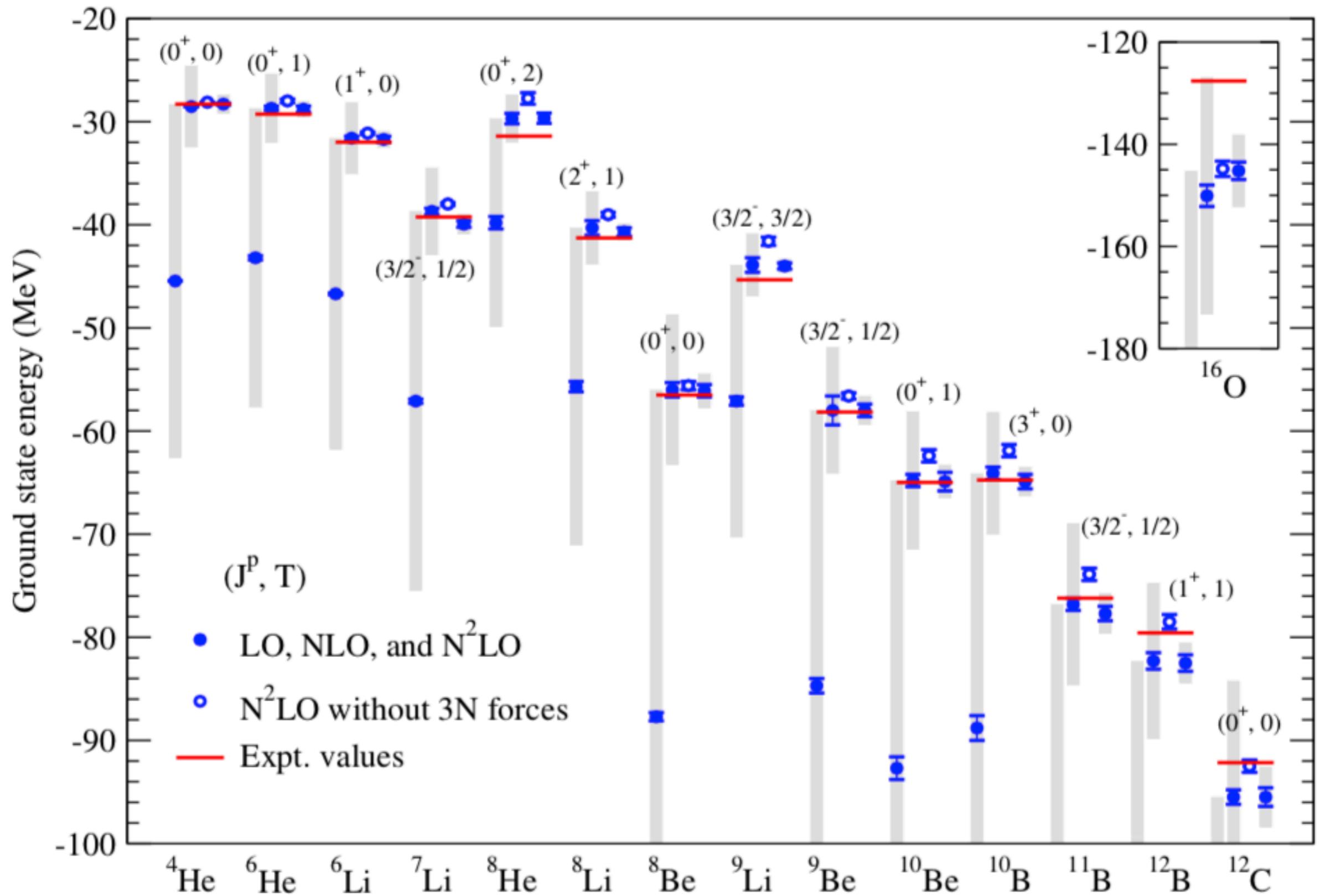
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- $\chi\text{EFT}$  at  $N^2LO$  reproduces binding energies of light nuclei reasonably well

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# Behavior of a $\chi$ EFT series

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  - In fact  $\{c_n\} = \{1, -0.46, 0.75\}$ . What then is expectation for  $c_3$ ?
- One possibility:  $c_3 = \max\{c_0, c_1, c_2\}$

Epelbaum, Krebs, Meissner (2014)

cf. McGovern, Griesshammer, Phillips (2013); many others.

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# Bayesian tools

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Thomas Bayes (1701?-1761)



<http://www.bayesian-inference.com>

$$\text{pr}(A|B, I) = \frac{\text{pr}(B|A, I)\text{pr}(A|I)}{\text{pr}(B|I)}$$

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Likelihood

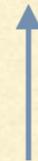


Prior



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Posterior



Normalization



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Posterior



Normalization



Marginalization:  $\text{pr}(x|\text{data}, I) = \int dy \text{pr}(x, y|\text{data}, I)$

Allows us to integrate out “nuisance” (e.g. higher-order) parameters

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# Probability for EFT coefficients

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Furnstahl, Kleo, DP, Wesolowski, PRC,2015 after Cacciari and Houdeau, JHEP, 2011

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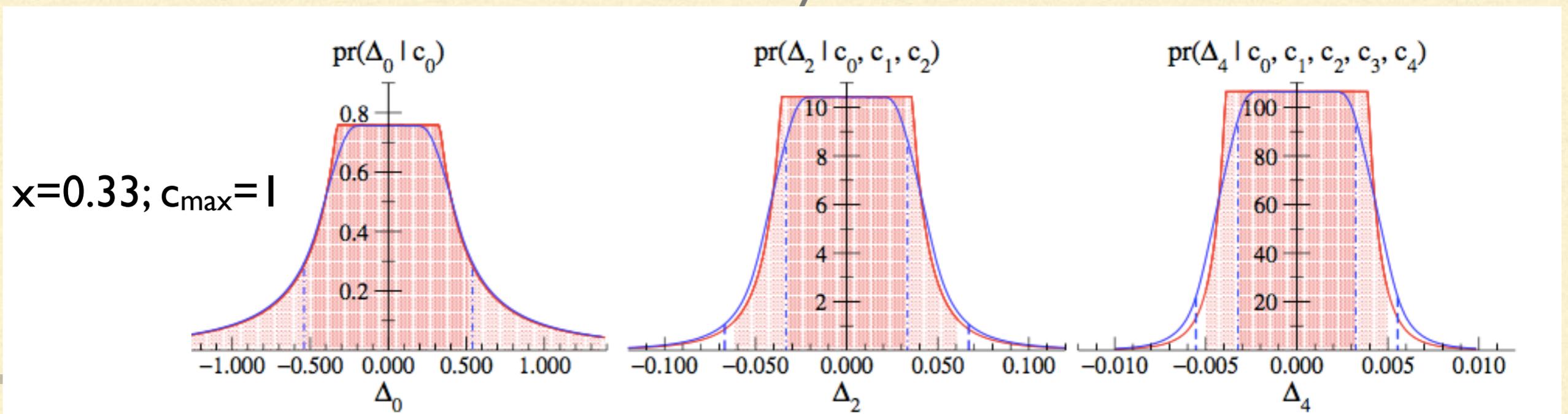
$$\text{Result: } \text{pr}(c_{k+1} | c_0, c_1, \dots, c_k) \propto \begin{cases} 1 & \text{if } c_{k+1} < c_{\max} \\ \left(\frac{c_{\max}}{c_{k+1}}\right)^{k+2} & \text{if } c_{k+1} > c_{\max} \end{cases}$$

$[-c_{\max} X_0 x^{k+1}, c_{\max} X_0 x^{k+1}]$  is a  $\frac{k+1}{k+2} * 100\%$  DoB interval

# Probability for EFT coefficients

Furnstahl, Kleo, DP, Wesolowski, PRC,2015 after Cacciari and Houdeau, JHEP, 2011

- General EFT series for observable to order  $k$ :  $X = X_0 \sum_{i=0}^k c_i x^i$
- Compute conditional probability distribution:  $\text{pr}(c_{k+1} | c_0, \dots, c_k, I)$
- $I$ =information about EFT, e.g. naturalness
- “Prior A”:  $\text{pr}(c_n | \bar{c}) = \frac{1}{2\bar{c}} \theta(\bar{c} - c_n)$ ;  $\text{pr}(\bar{c}) = \frac{1}{2 \ln(\epsilon) \bar{c}} \theta\left(\frac{1}{\epsilon} - \bar{c}\right) \theta(\bar{c} - \epsilon)$
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# NN scattering cross sections

- NN cross section at  $T_{\text{lab}}=50, 96, 143, 200$  MeV

Epelbaum, Krebs, Meissner, EPJA, 2015

- Potential regulated by local function, parameterized by R. Here:  $R=0.9$  fm data

$$\sigma_{np}(E_{\text{lab}}) = \sigma_{\text{LO}} \sum_{n=0}^k c_n(p_{\text{rel}}) \left( \frac{p_{\text{rel}}}{\Lambda_b} \right)^n$$

- Results at LO, NLO, N<sup>2</sup>LO, N<sup>3</sup>LO, N<sup>4</sup>LO ( $k=0, 2, 3, 4, 5$ )

$$x = \frac{p_{\text{rel}}}{\Lambda_b}$$

EKM state  
 $\Lambda_b=600$  MeV

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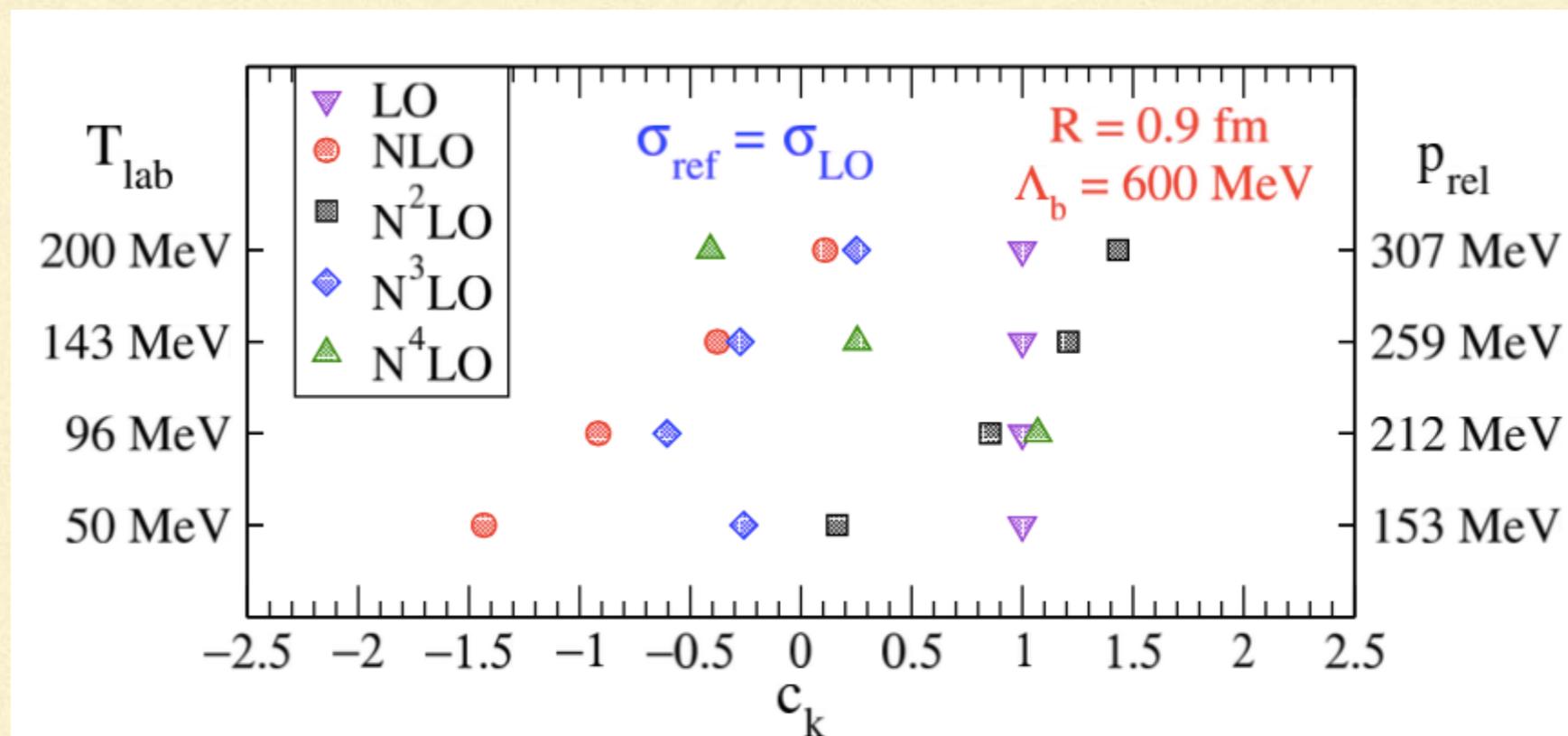
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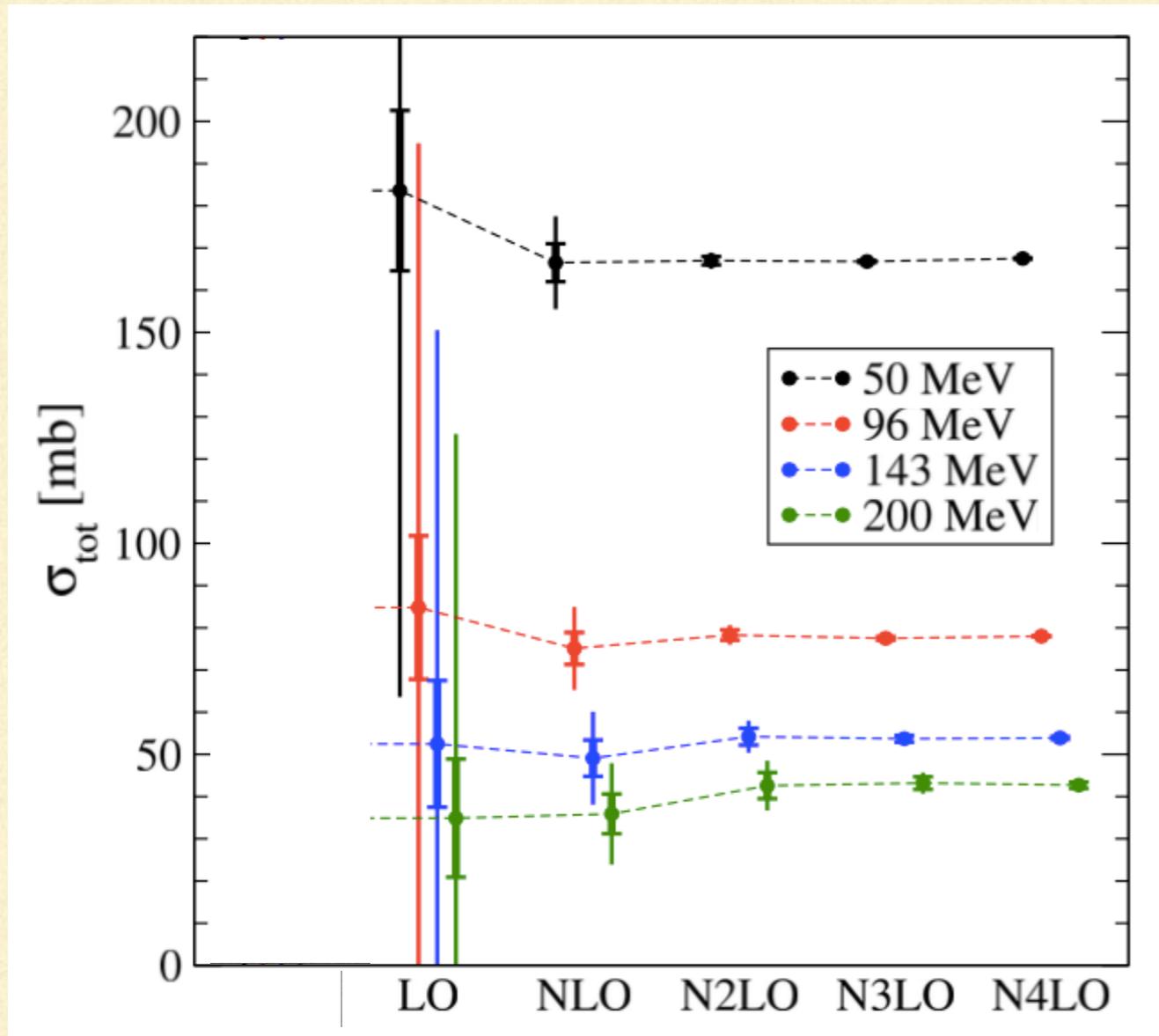
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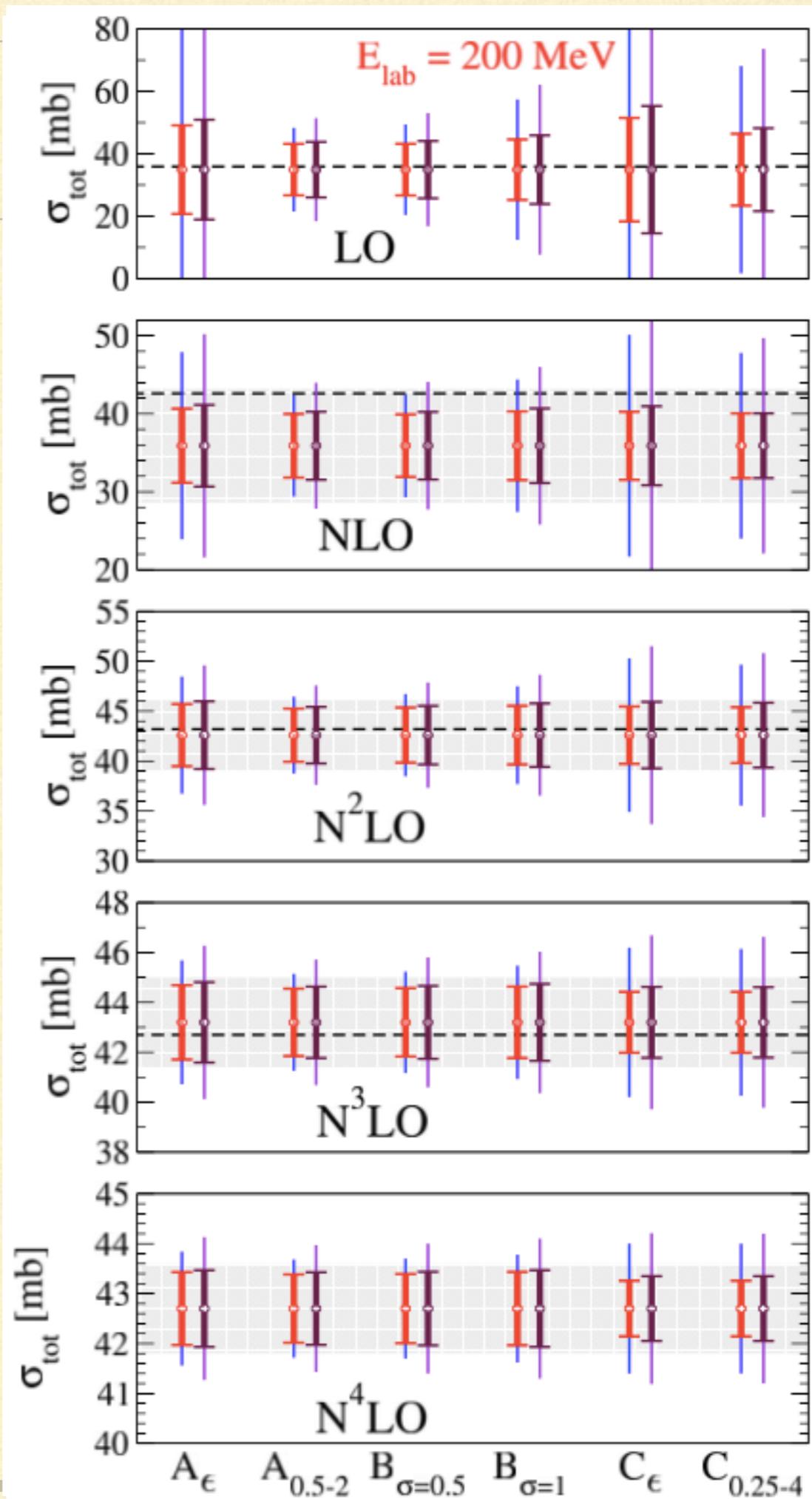
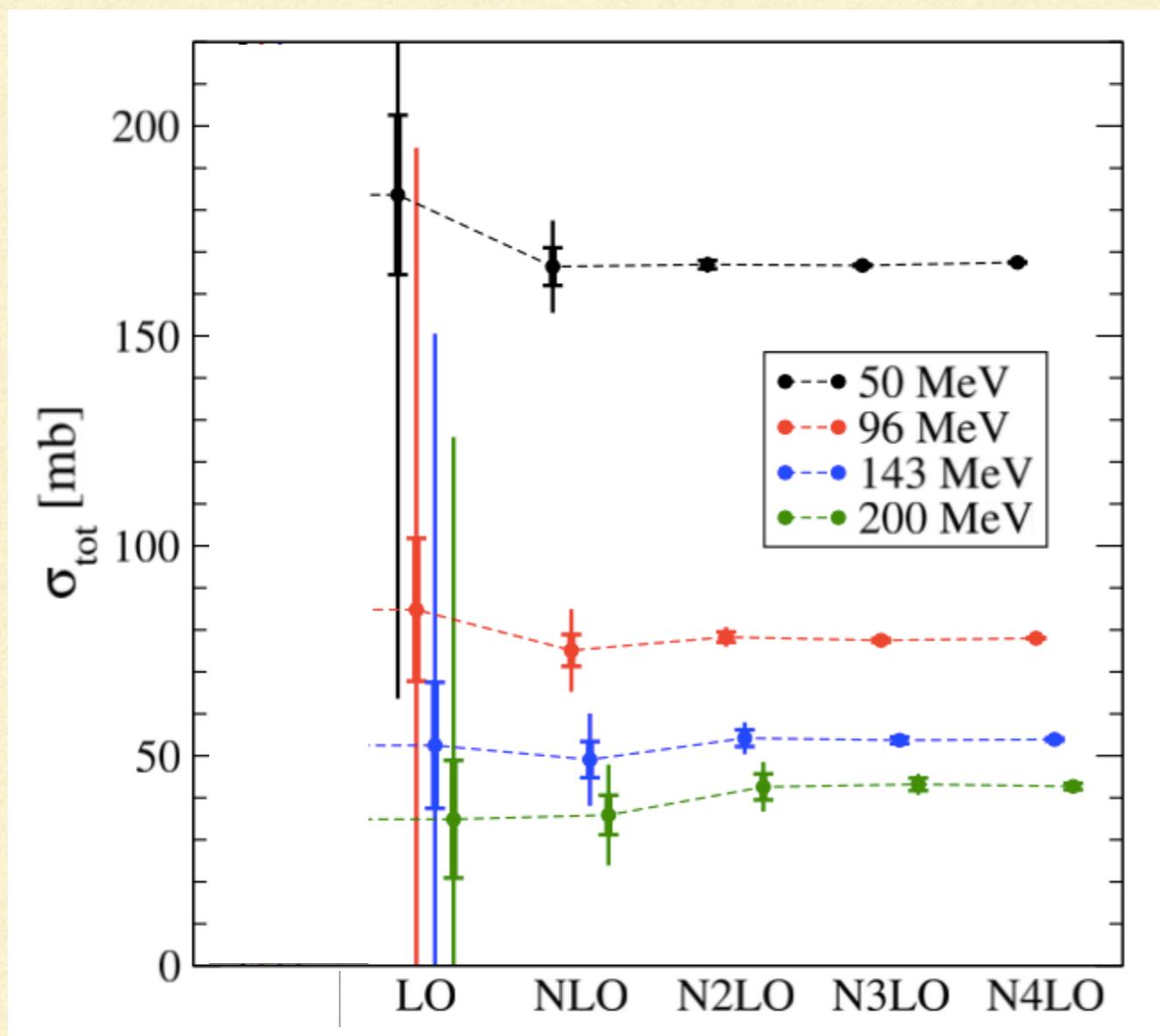


# Results





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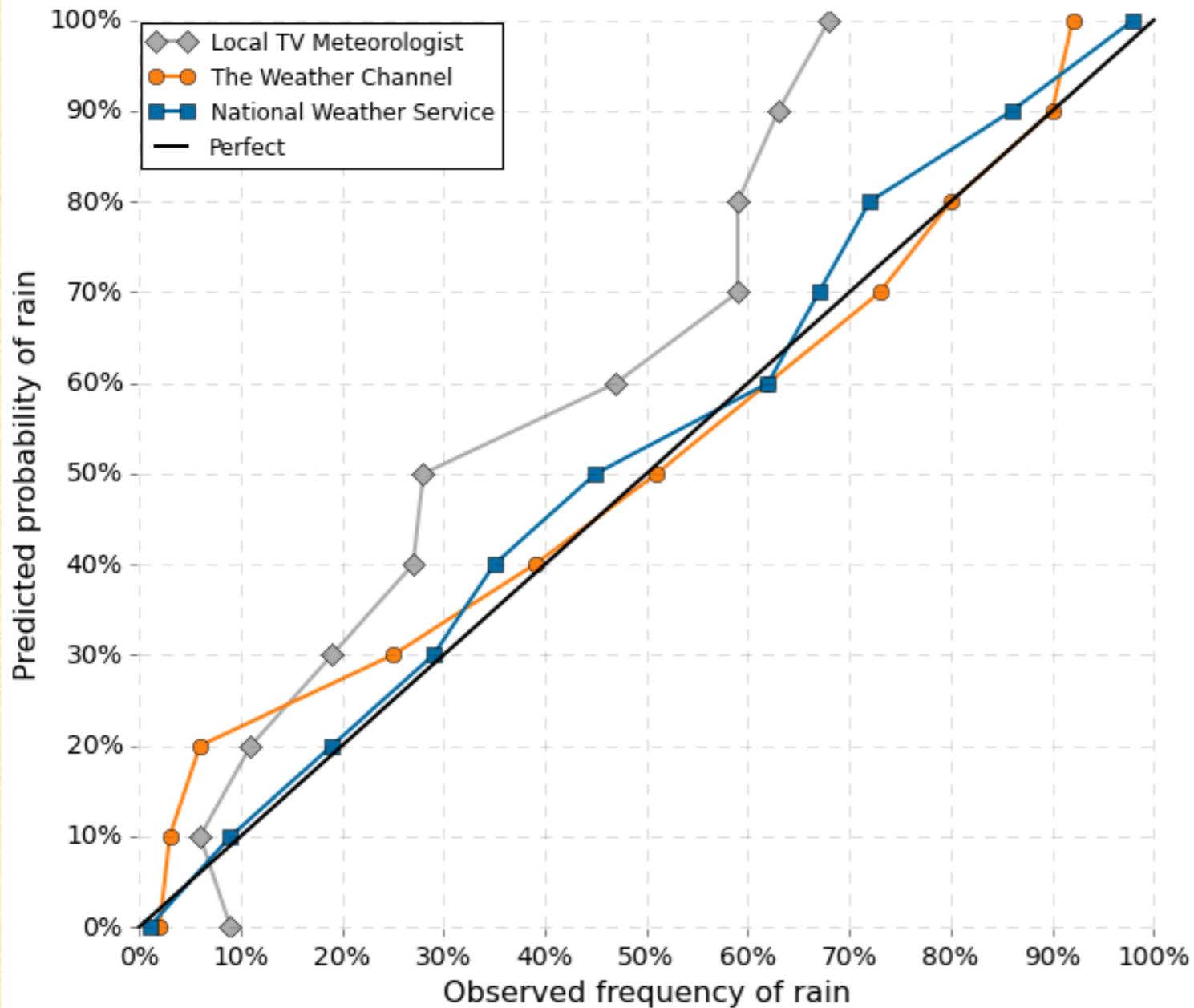
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# The well-calibrated EFTist

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# The well-calibrated EFTist

Accuracy of three weather forecasting services

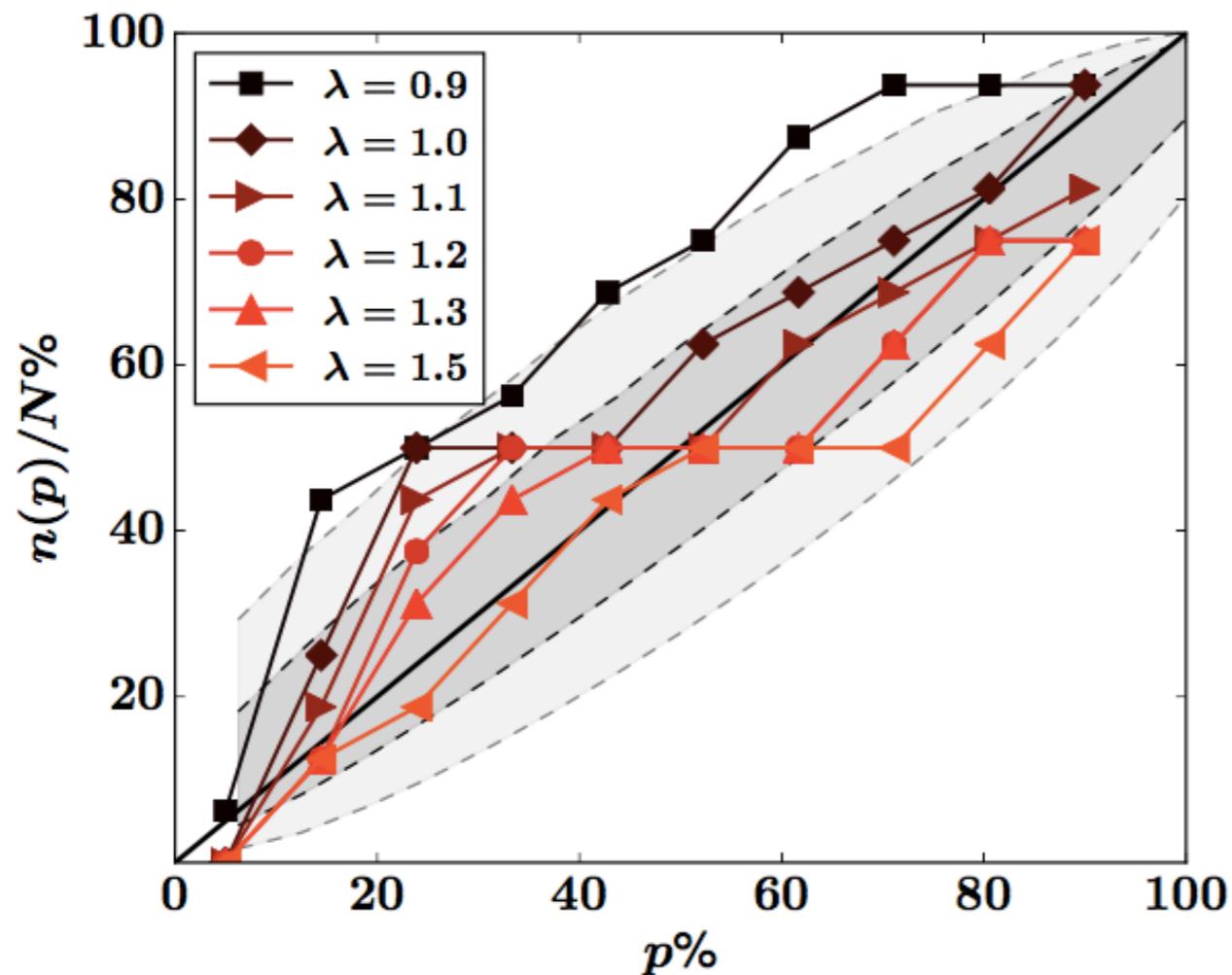


Source: "The Signal and the Noise" by Nate Silver | Author: Randy Olson (randalolson.com / @randal\_olson)

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Furnstahl, Klco, DP, Wesolowski, PRC, 2015

after: Bagnaschi, Cacciari, Guffanti, Jenniches, 2015

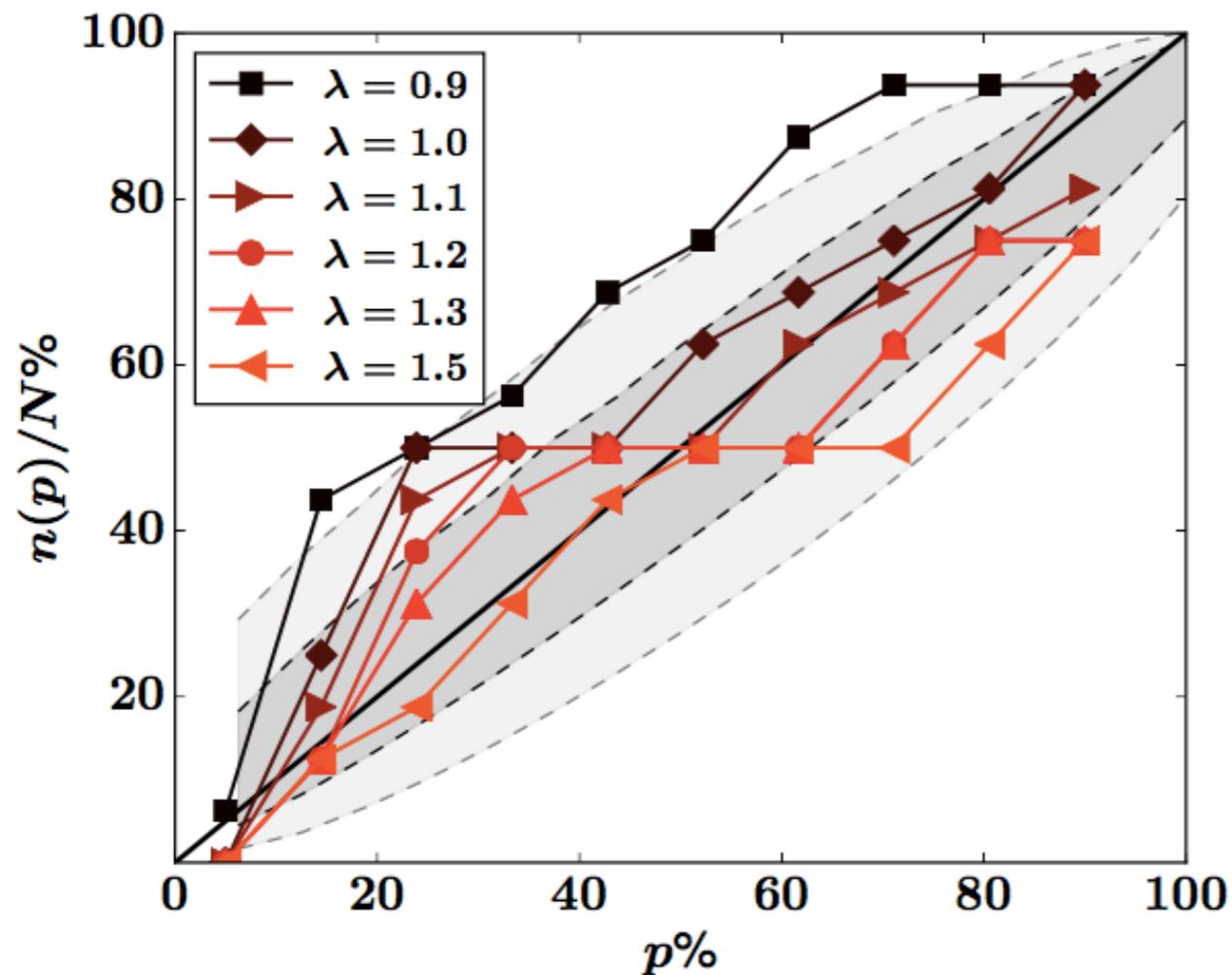


- Now we consider predictions at each order, with their error bars, as data and test them to see if the procedure is consistent
- Fix a given DOB interval, compute actual success ratio and compare
- Look at this over EKM predictions at four different orders and four different energies
- Interpret in terms of rescaling of  $\Lambda_b$  by a factor  $\lambda$

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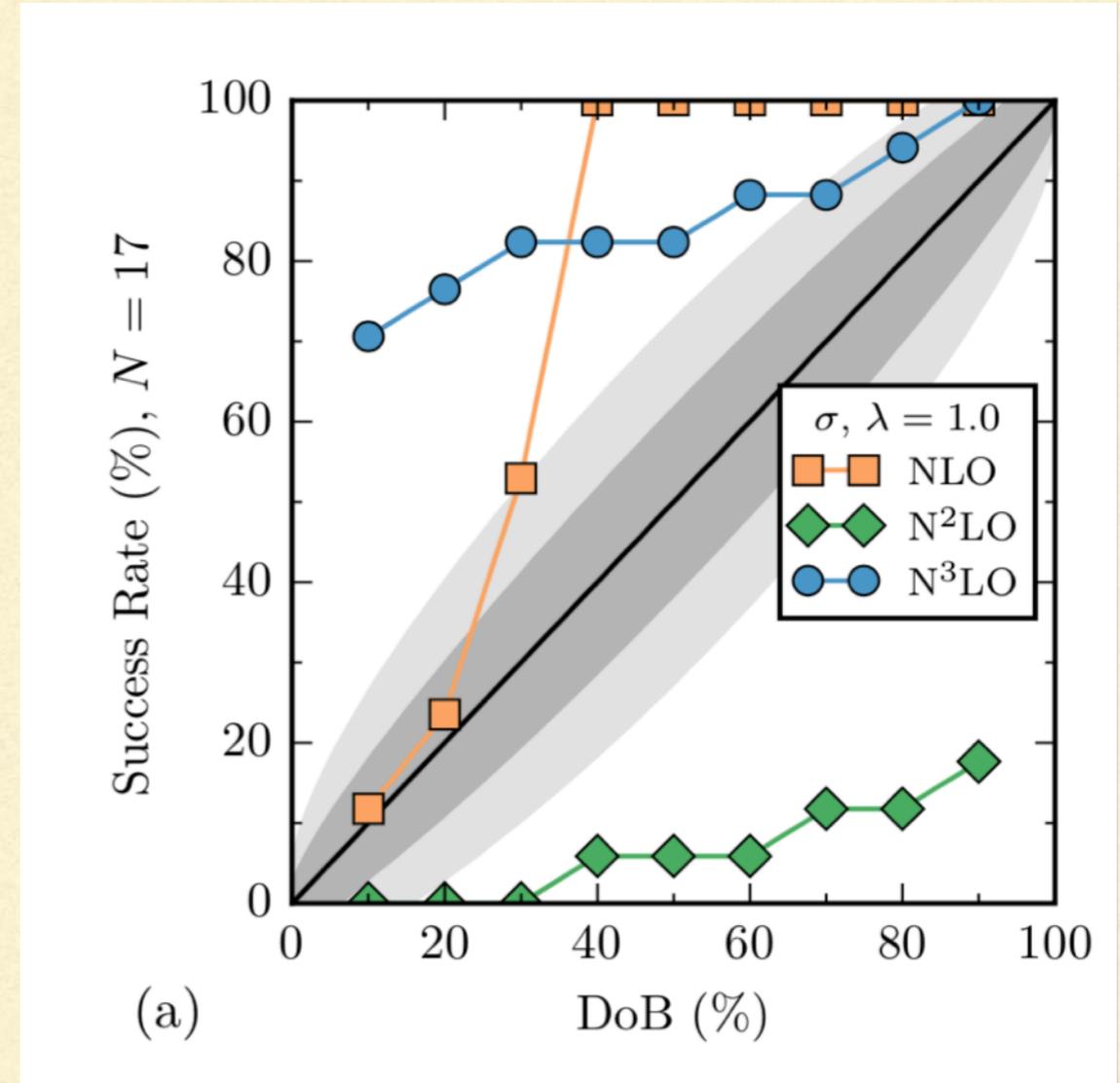
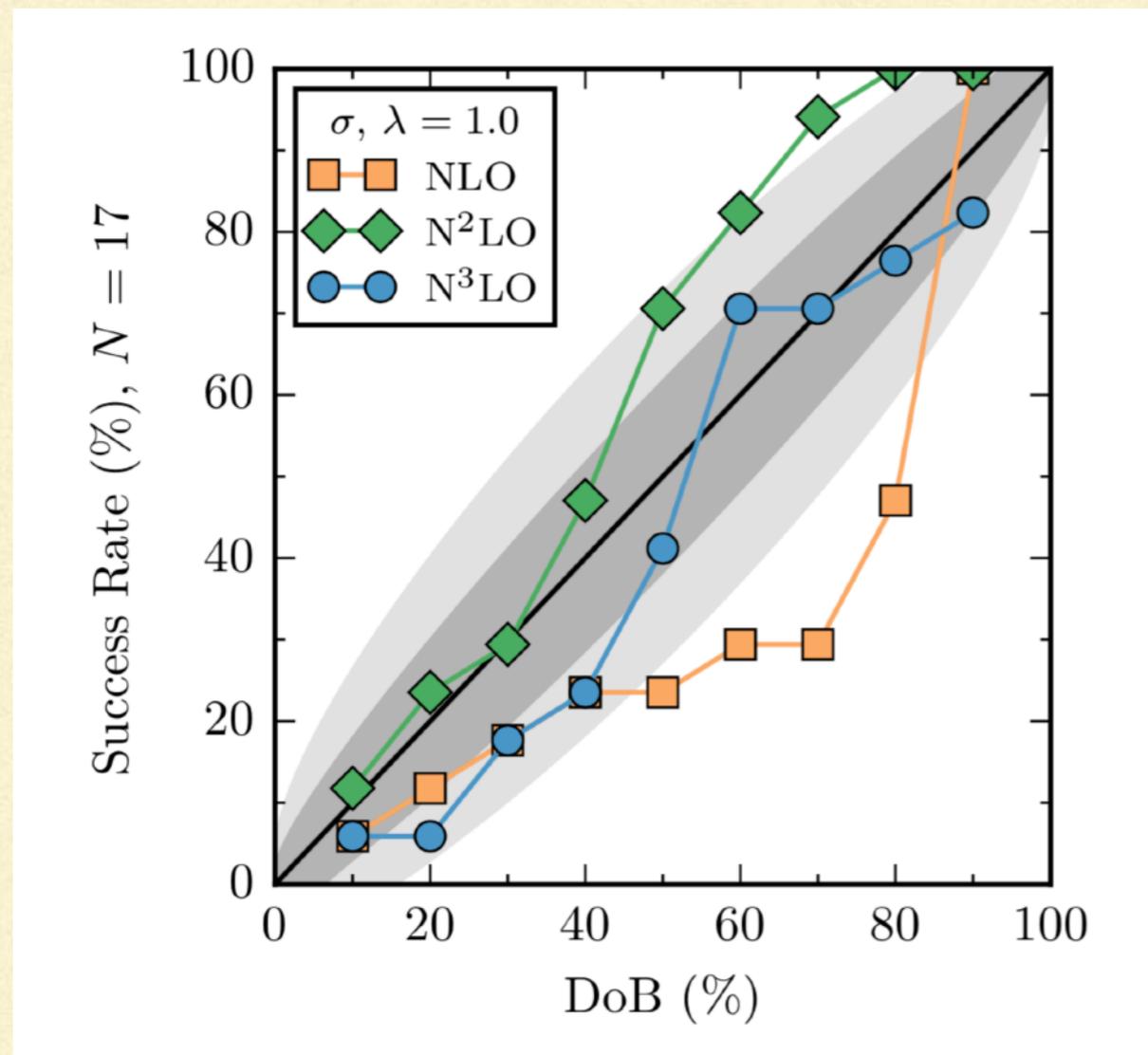
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No evidence for significant rescaling of  $\Lambda_b$

# Physics from consistency plots

**R=0.9 fm**

**R=1.2 fm**



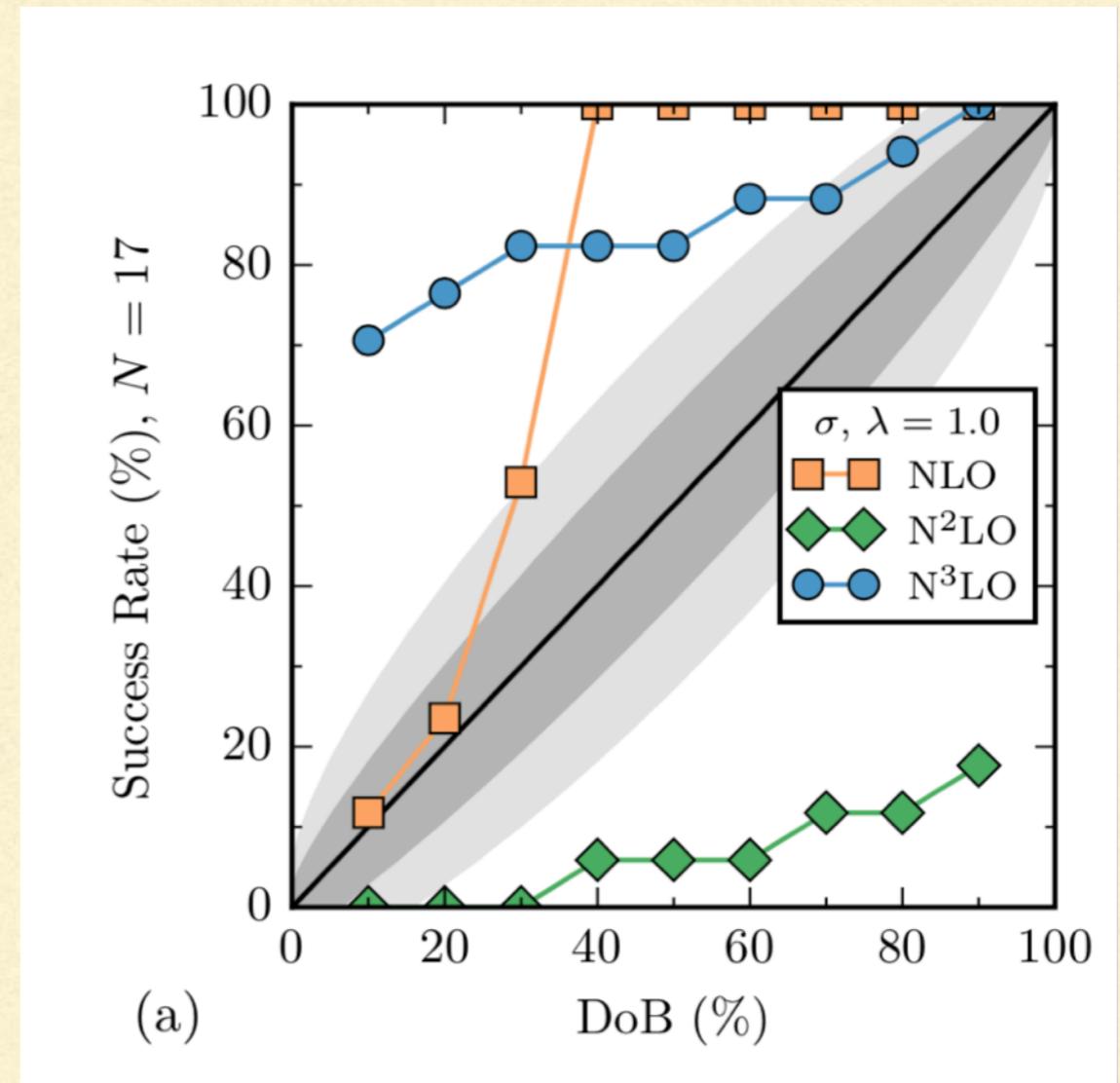
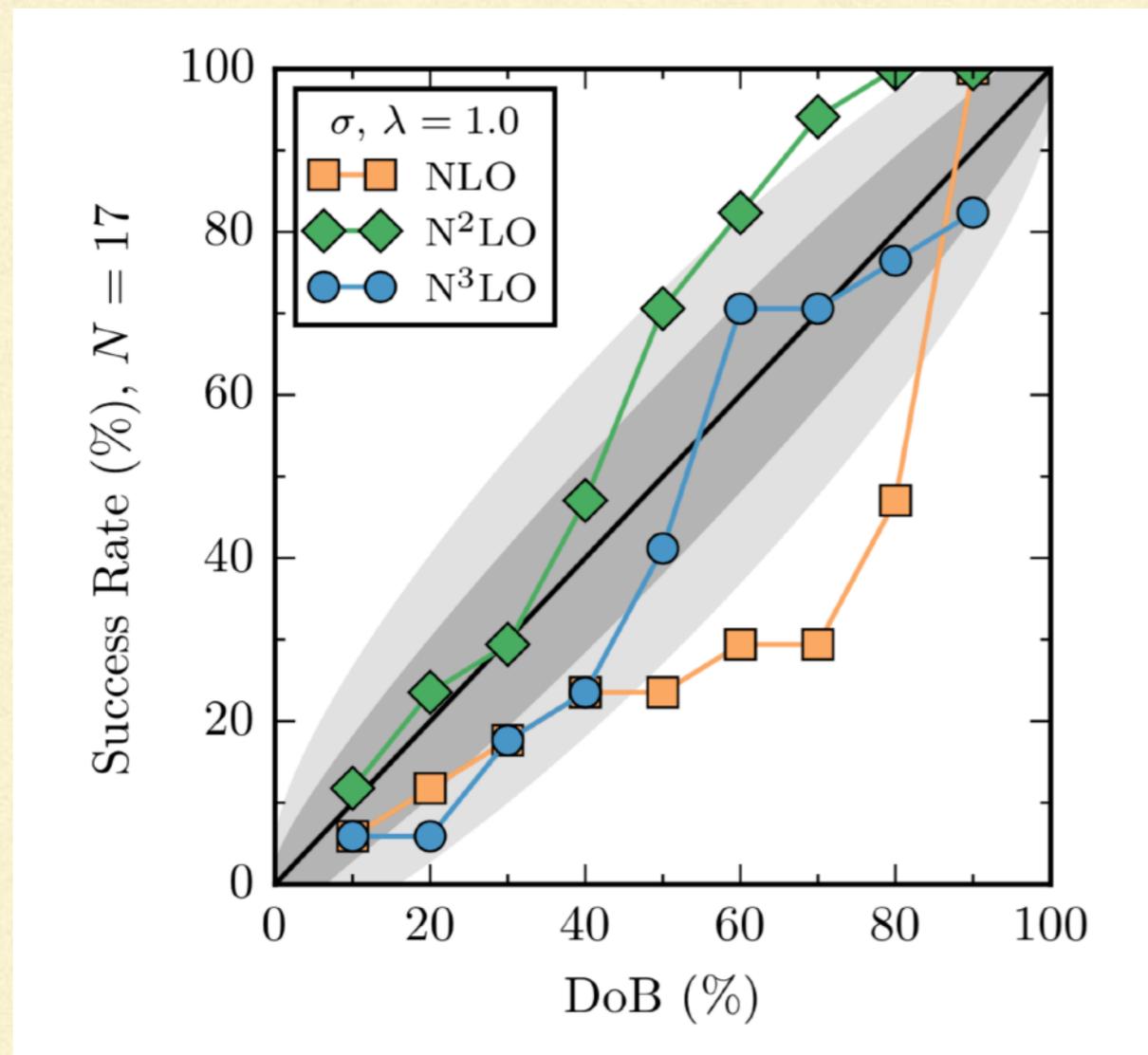
- Allows assessment of order-by-order convergence
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Melendez, Furnstahl, Wesolowski, PRC, 2017

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# Outline

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- What we do and don't know about the strong nuclear force ✓
  - EFT: organizing what we know, constraining what we don't ✓
  - EFT truncation errors from a Bayesian analysis: NN scattering ✓
  - EFT for halo nuclei: universal formula for  $\gamma + {}^A_Z \rightarrow {}^{A-1}_Z + n$
  - Uncertainty quantification for fusion:  ${}^7\text{Be}(p,\gamma)$  at solar energies
  - Conclusion
-

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# Ordinary vs. halo nuclei

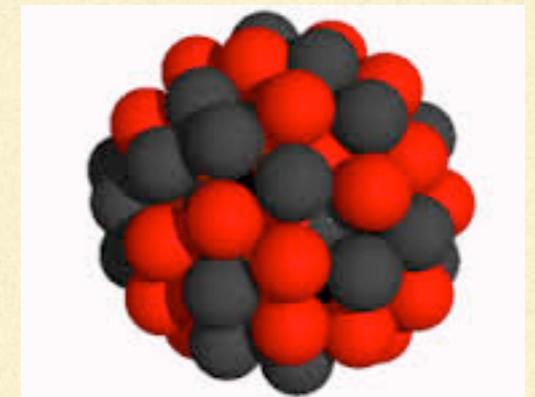
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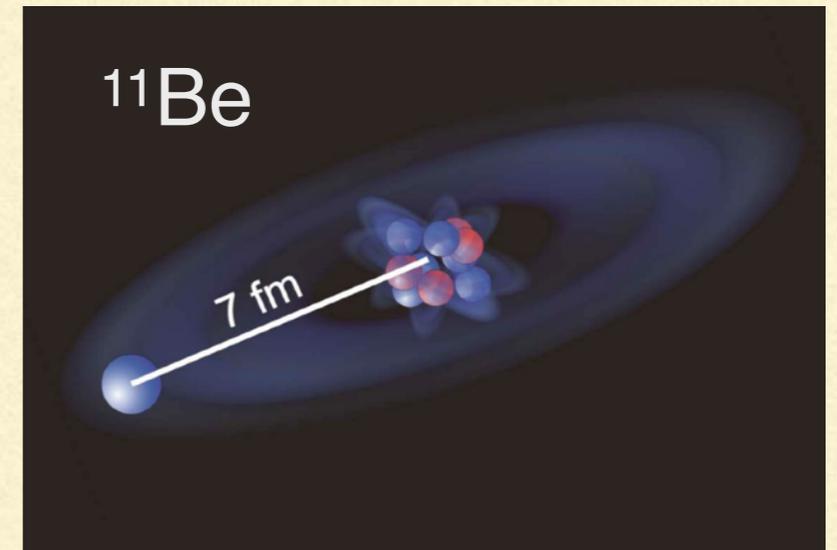
- In nuclei, each nucleon moves in the potential generated by the others
- The nuclear size grows as  $A^{1/3}$ ; cross sections like  $A^{2/3}$
- Nuclear binding energies are on the order of 8 MeV/nucleon



<http://alternativephysics.org>

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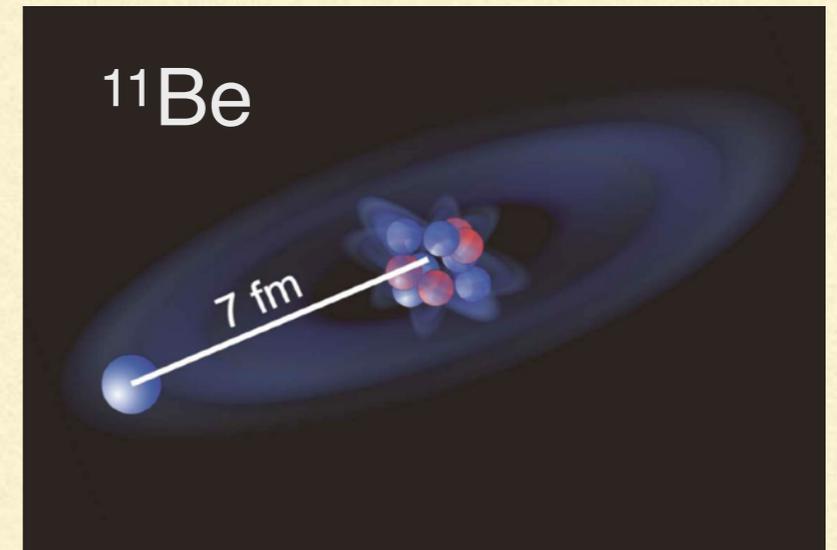
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- Halo nuclei: the last few nucleons “orbit” far from the nuclear “core”
- Characterized by small nucleon binding energies, large radii, large interaction cross sections, large EI transition strengths.

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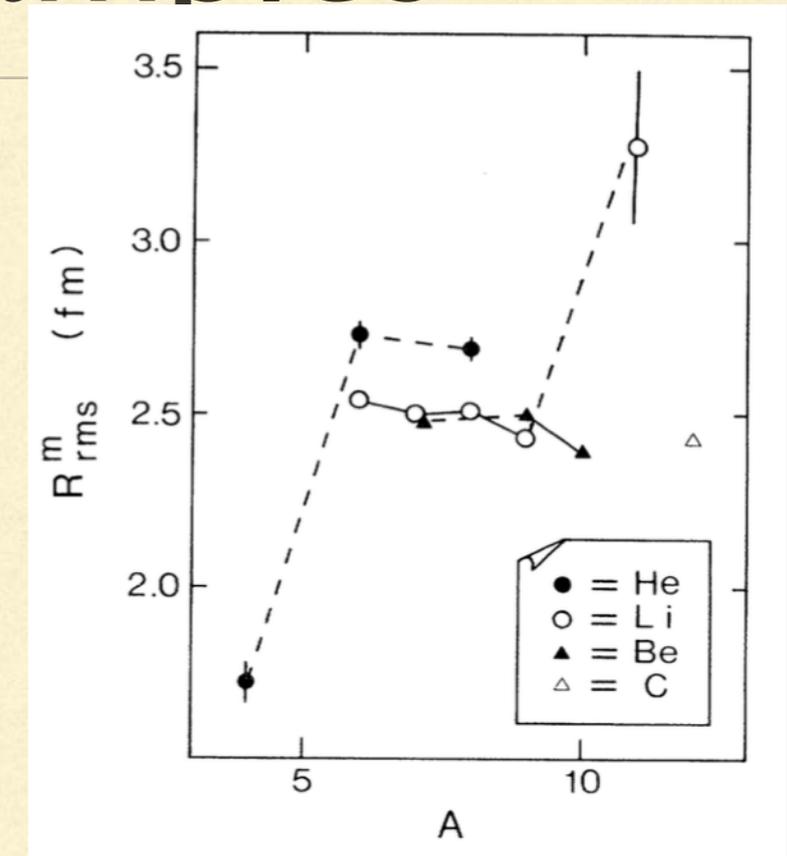
# Halo nuclei: history & examples

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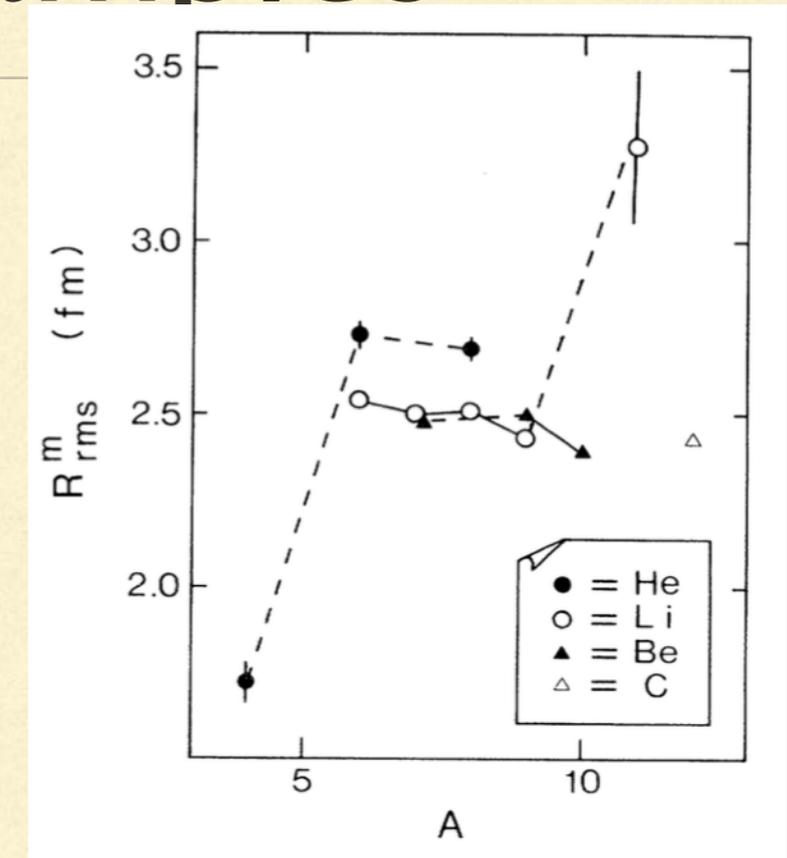
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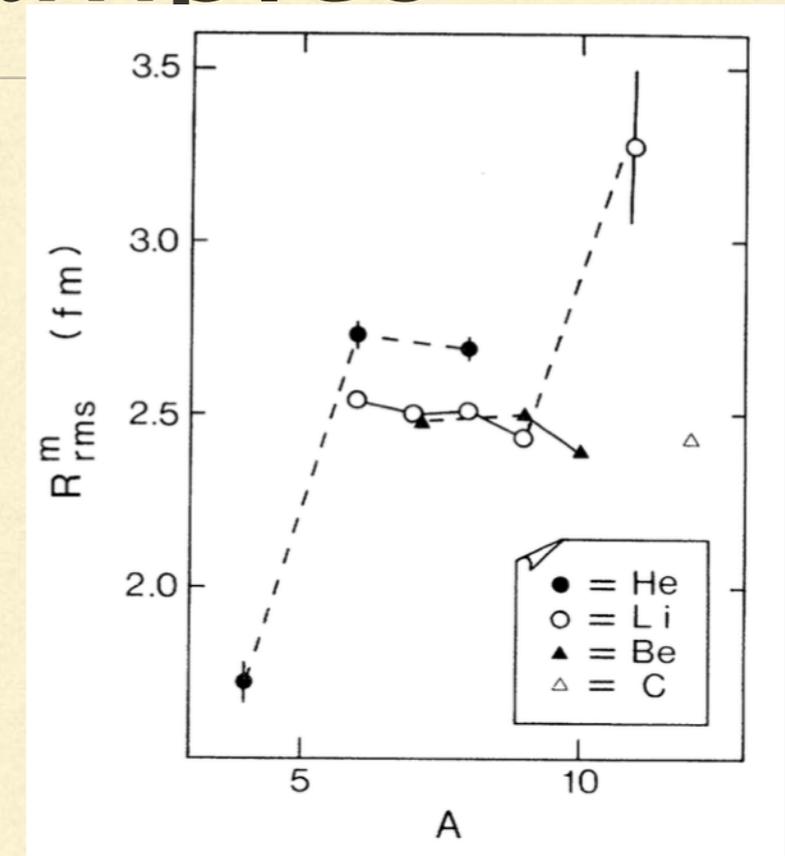
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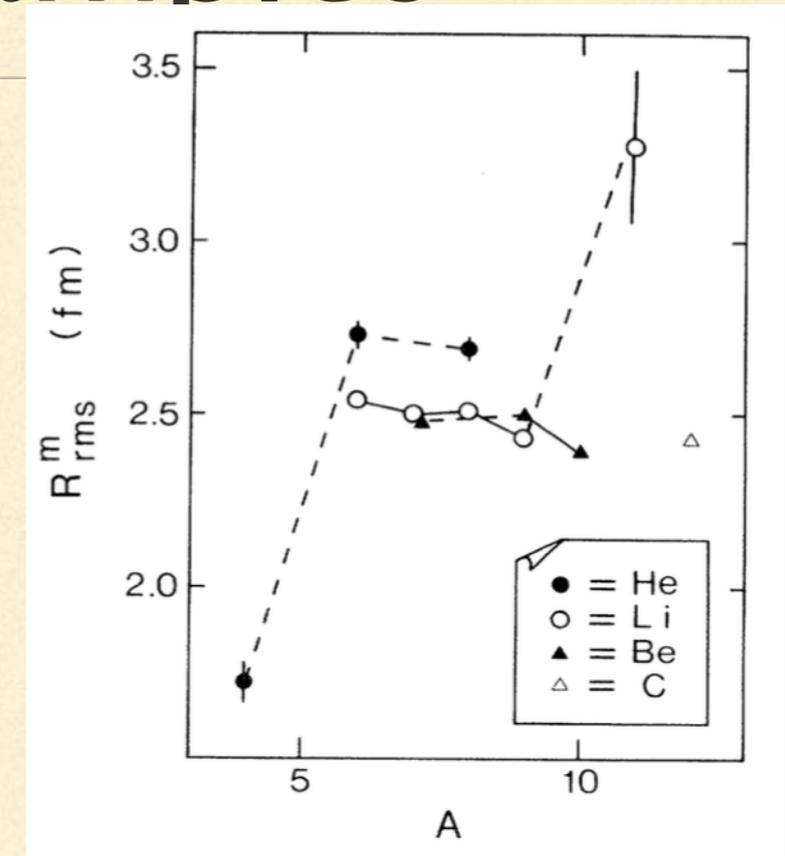
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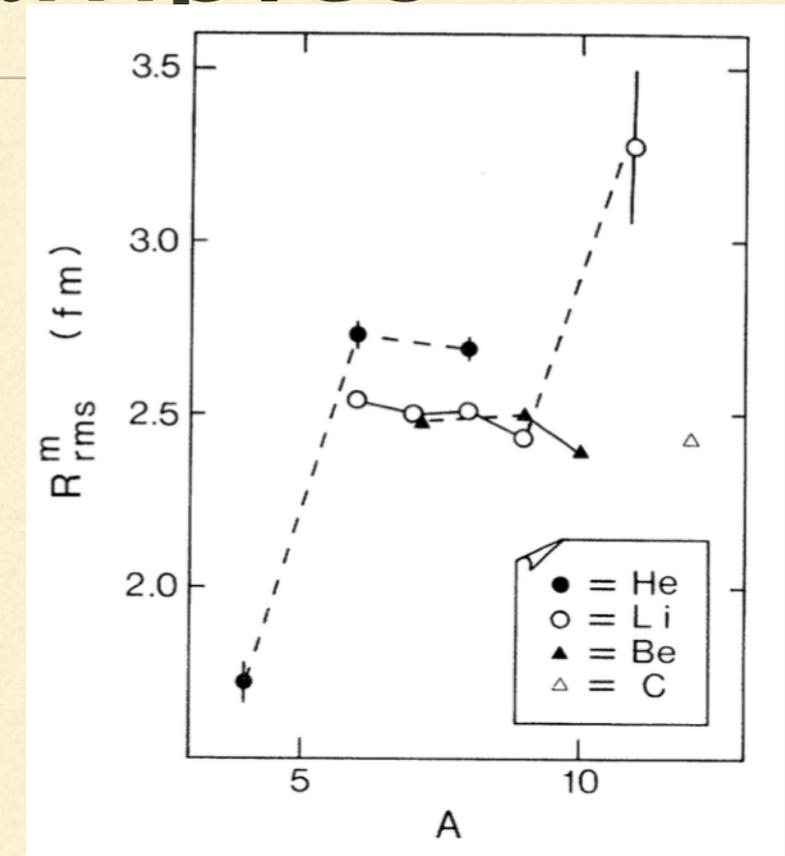
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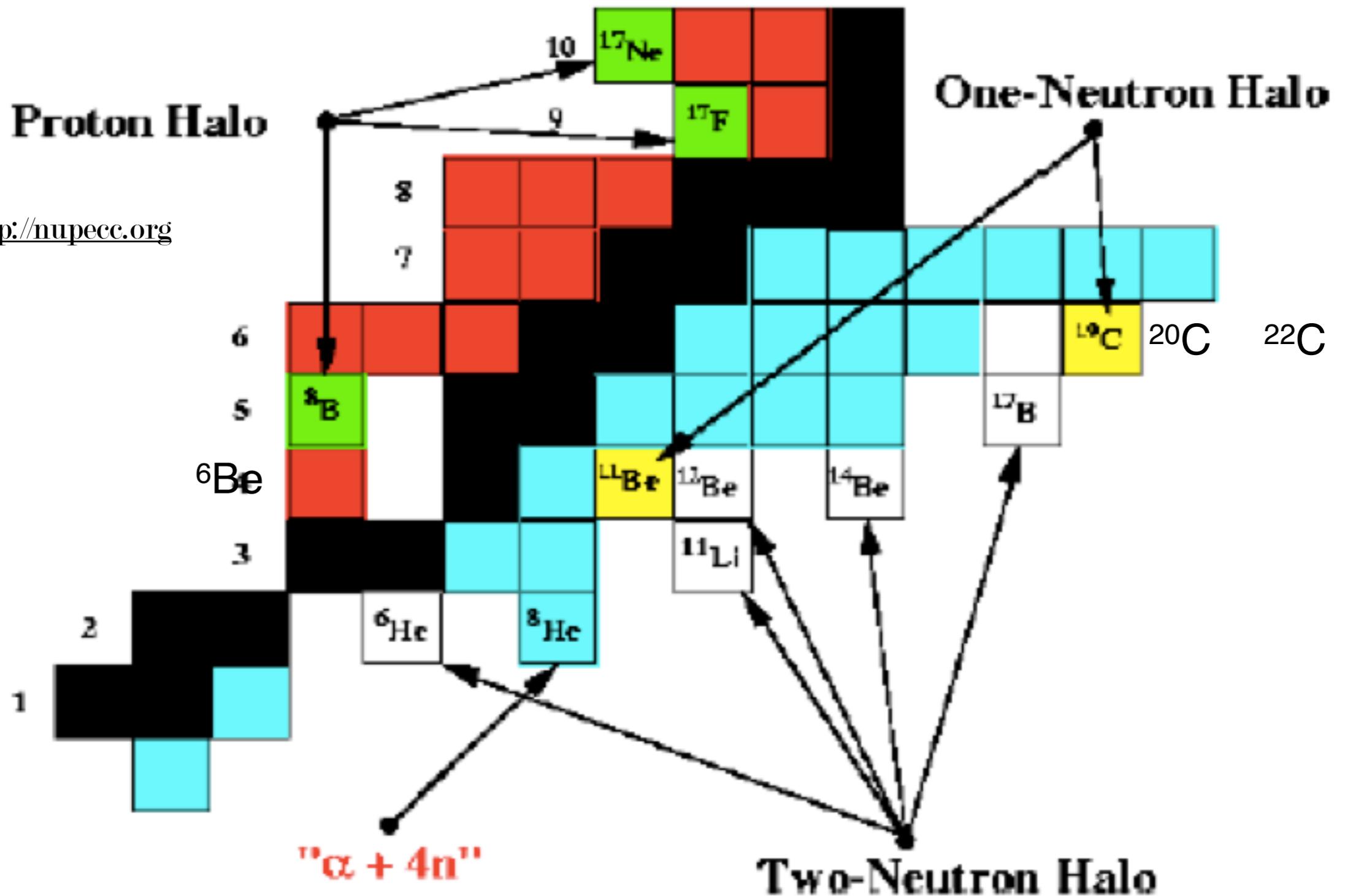
- “Open quantum systems”: physics beyond mean field

- Universality: common features of weakly-bound quantum systems



# Halo nuclei: history & examples

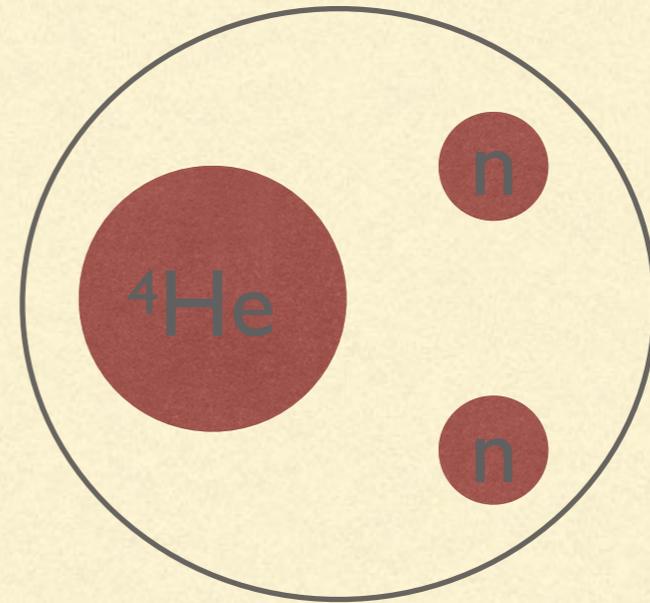
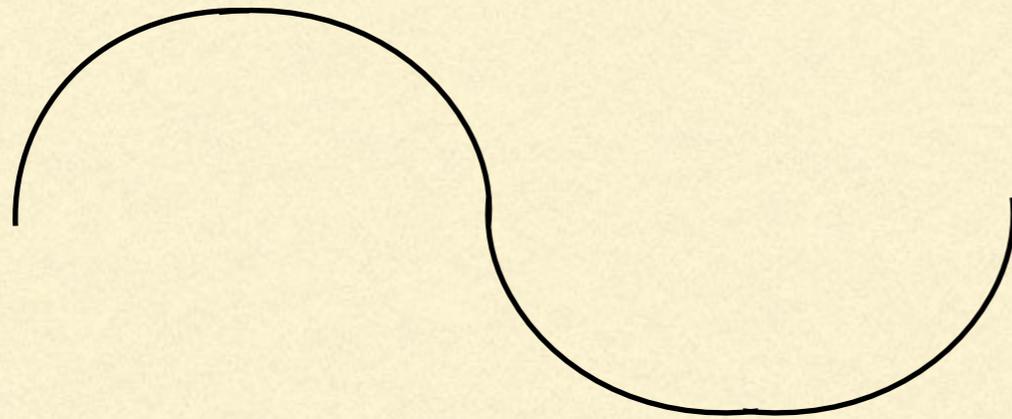
<http://nupecc.org>



# Halo EFT

Bertulani, Hammer, van Kolck, NPA (2003);  
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Review: Hammer, Ji, DP, J. Phys. G 44, 103002 (2017).

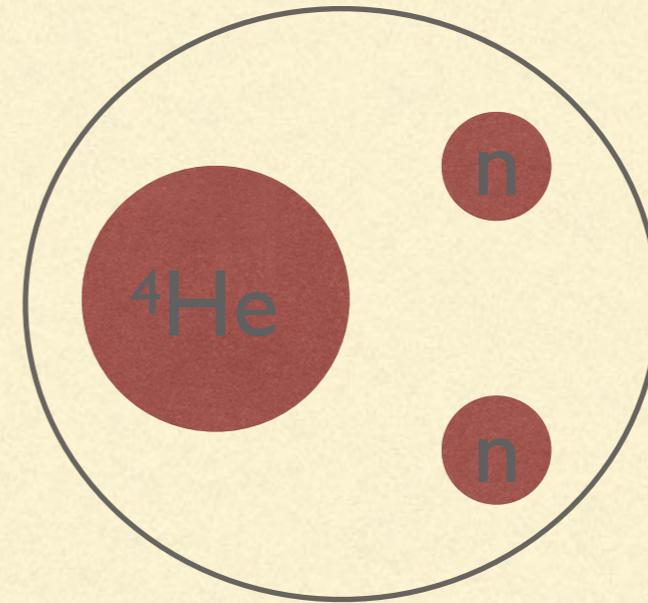
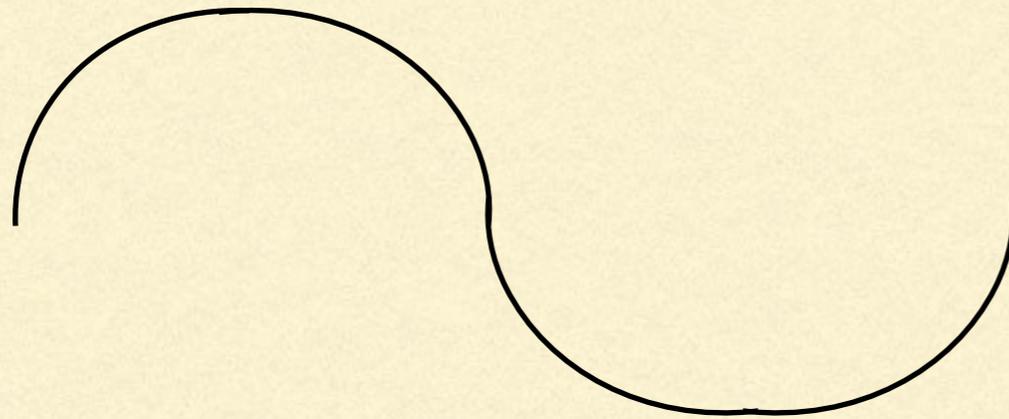
$$\lambda \gg R_{\text{core}}; \lambda \lesssim R_{\text{halo}}$$



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$$\lambda \gg R_{\text{core}}; \lambda \lesssim R_{\text{halo}}$$



- Define  $R_{\text{halo}} = \langle r^2 \rangle^{1/2}$ . Seek EFT expansion in  $R_{\text{core}}/R_{\text{halo}}$ . Valid for  $\lambda \lesssim R_{\text{halo}}$
- Typically  $R \equiv R_{\text{core}} \sim 2$  fm. And since  $\langle r^2 \rangle$  is related to the neutron separation energy we are looking for systems with neutron separation energies of order 1 MeV or less
- By this definition the deuteron is the lightest halo nucleus, and the pionless EFT for few-nucleon systems is a specific case of Halo EFT

# Predicting dissociation

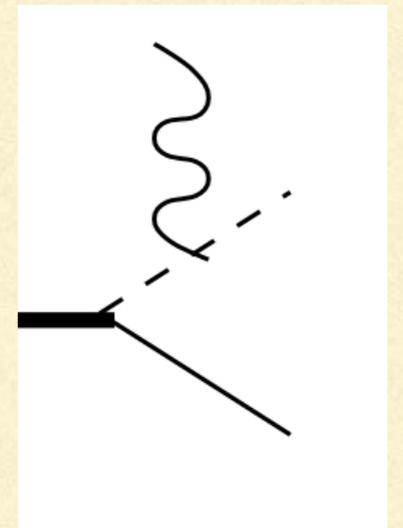
$$\mathcal{M} = \frac{eZg_0 2m_R}{\gamma_0^2 + \left(\mathbf{p} - \frac{\mathbf{k}}{A}\right)^2}$$

$$\gamma_0 = \sqrt{2m_R S_{1n}}$$

$$p = \sqrt{2m_R E}$$

$$E1 \propto \int_0^\infty dr j_1(pr) r u_0(r);$$

$$u_0(r) = A_0 e^{-\gamma_0 r}$$



Chen, Savage (1999)

# Predicting dissociation

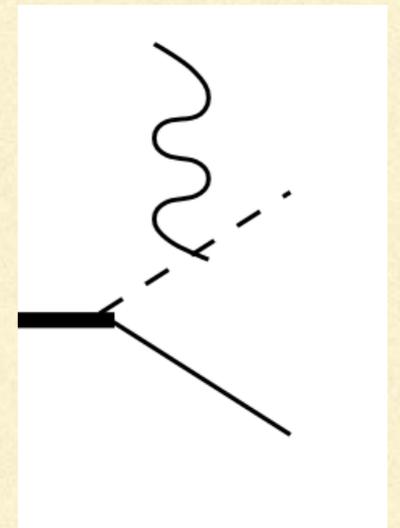
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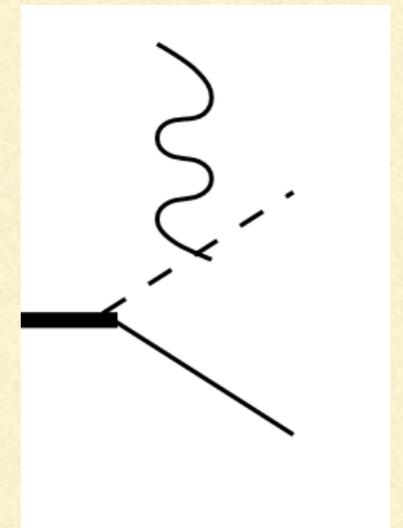
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Chen, Savage (1999)

$$\frac{dB(E1)}{e^2 dE} = \frac{6m_R}{\pi^2} \frac{Z^2}{A^2} A_0^2 \frac{p^3}{(\gamma_0^2 + p^2)^2}$$

Universal E1 strength formula for S-wave halos

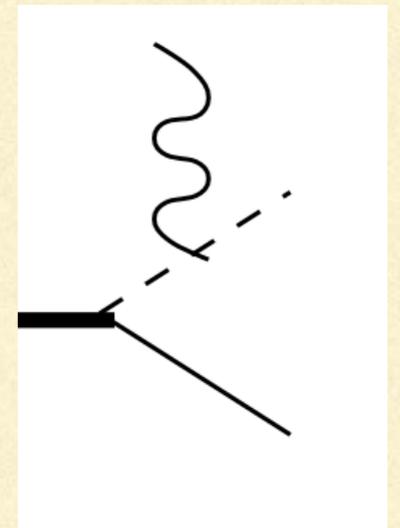
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Universal E1 strength formula for S-wave halos

- Final-state interactions suppressed by  $(R_{\text{core}}/R_{\text{halo}})^3$

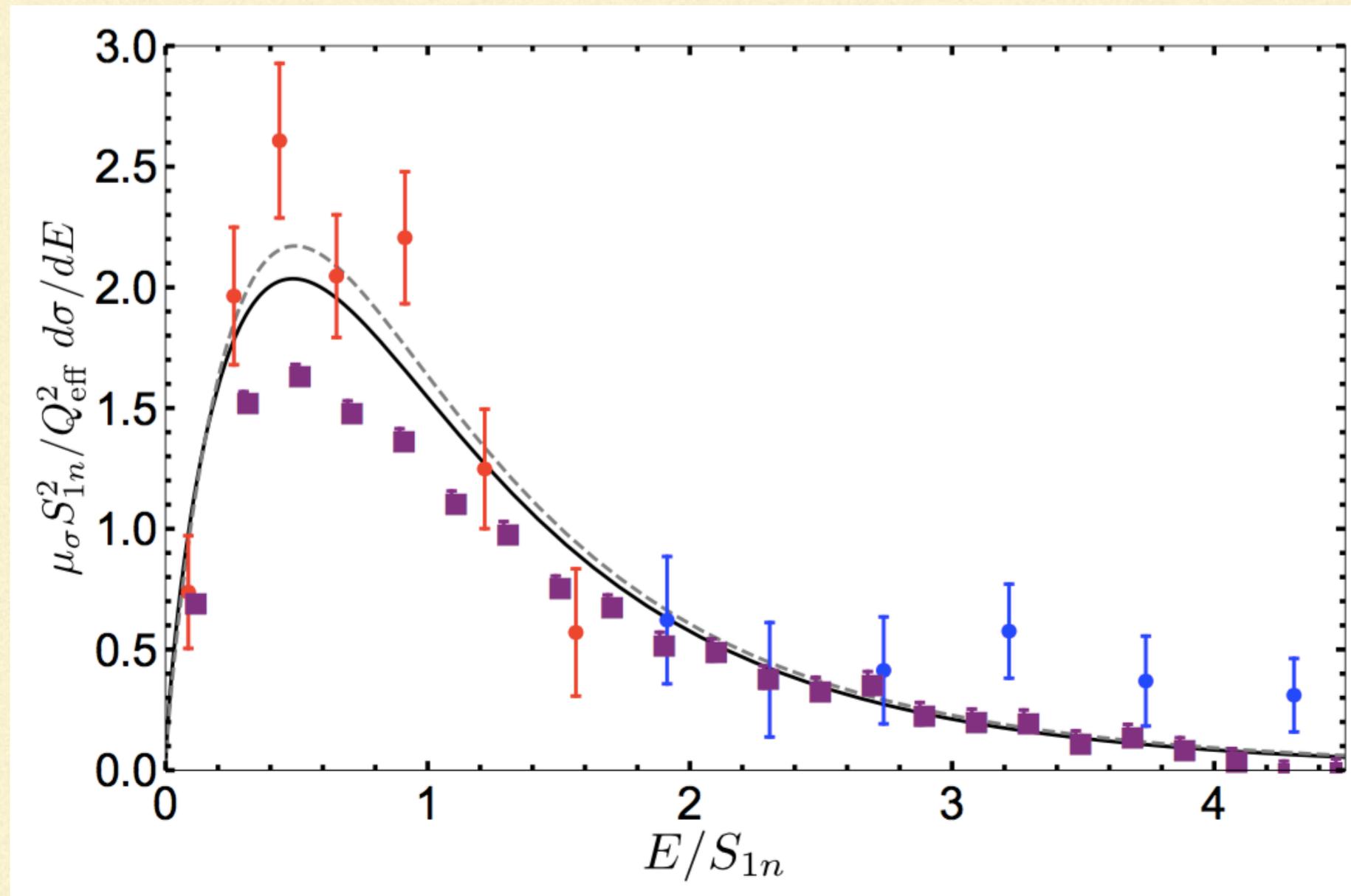
- Short-distance piece of E1 m.e.:  $L_{E1} \sigma^\dagger \mathbf{E} \cdot (n \overleftrightarrow{\nabla} c) + \text{h.c.} \sim \left(\frac{R_{\text{core}}}{R_{\text{halo}}}\right)^4$

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# Results

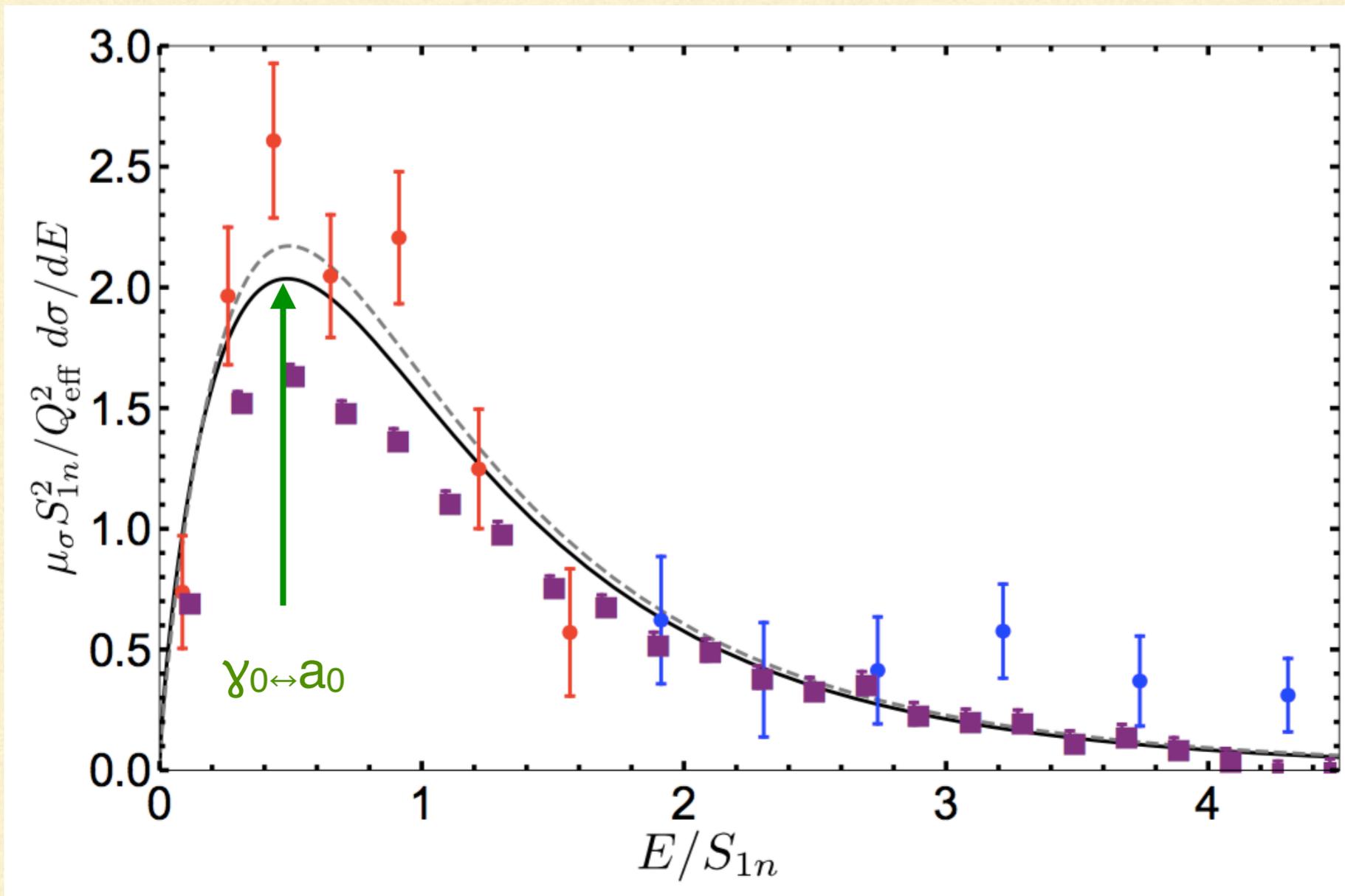
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# Results



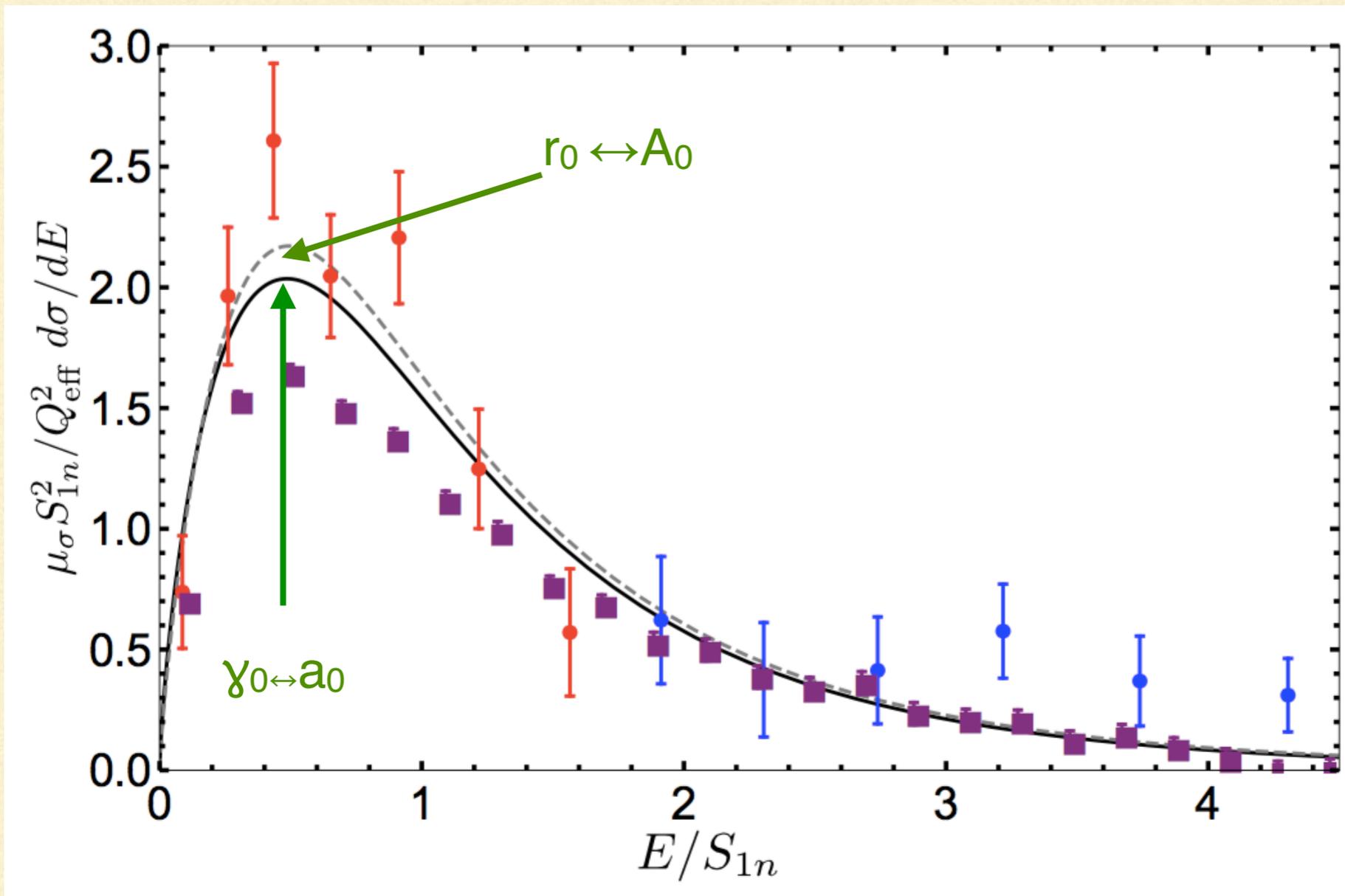
Data: Nakamura et al., 1999, 2003;  
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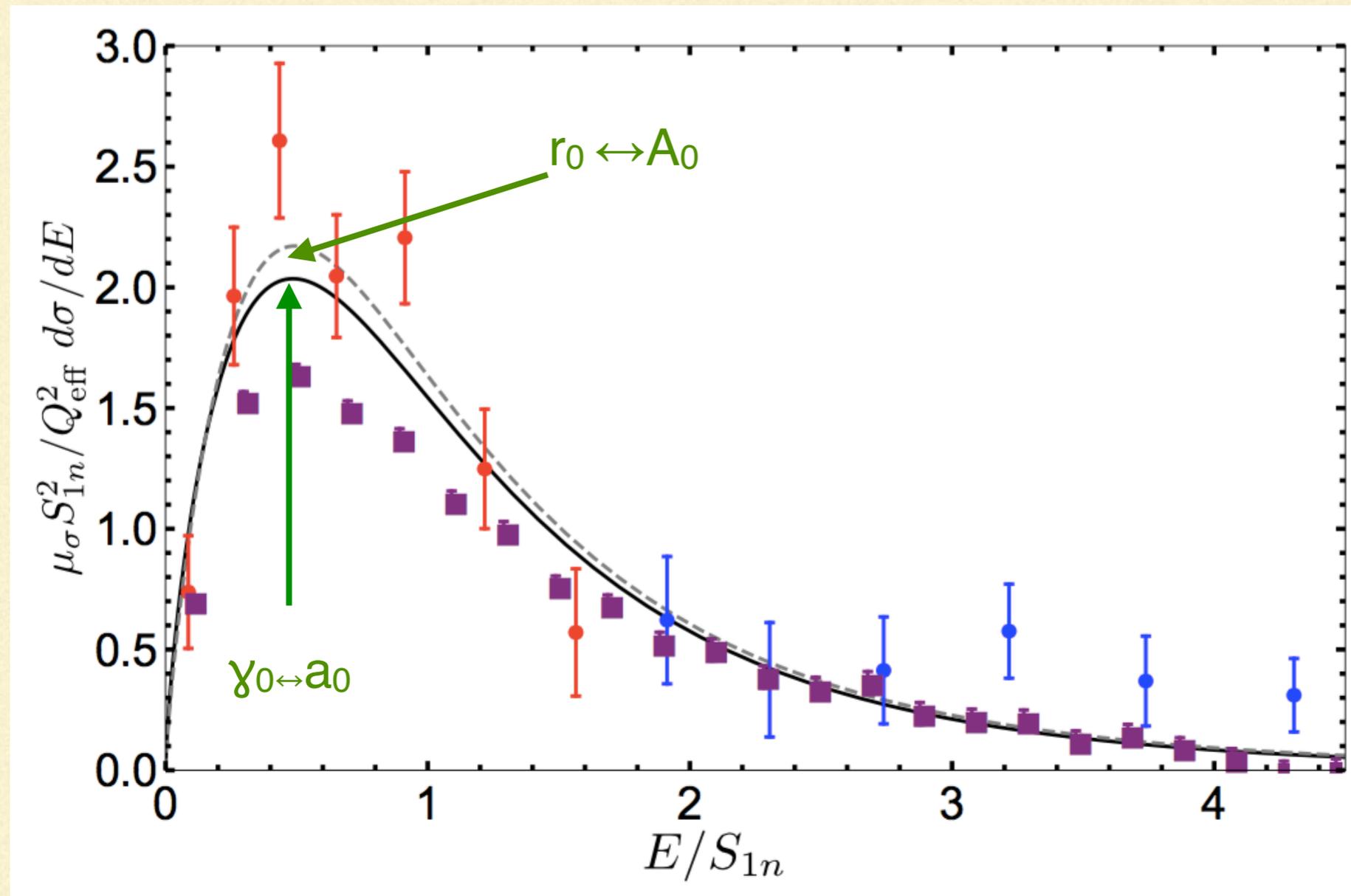
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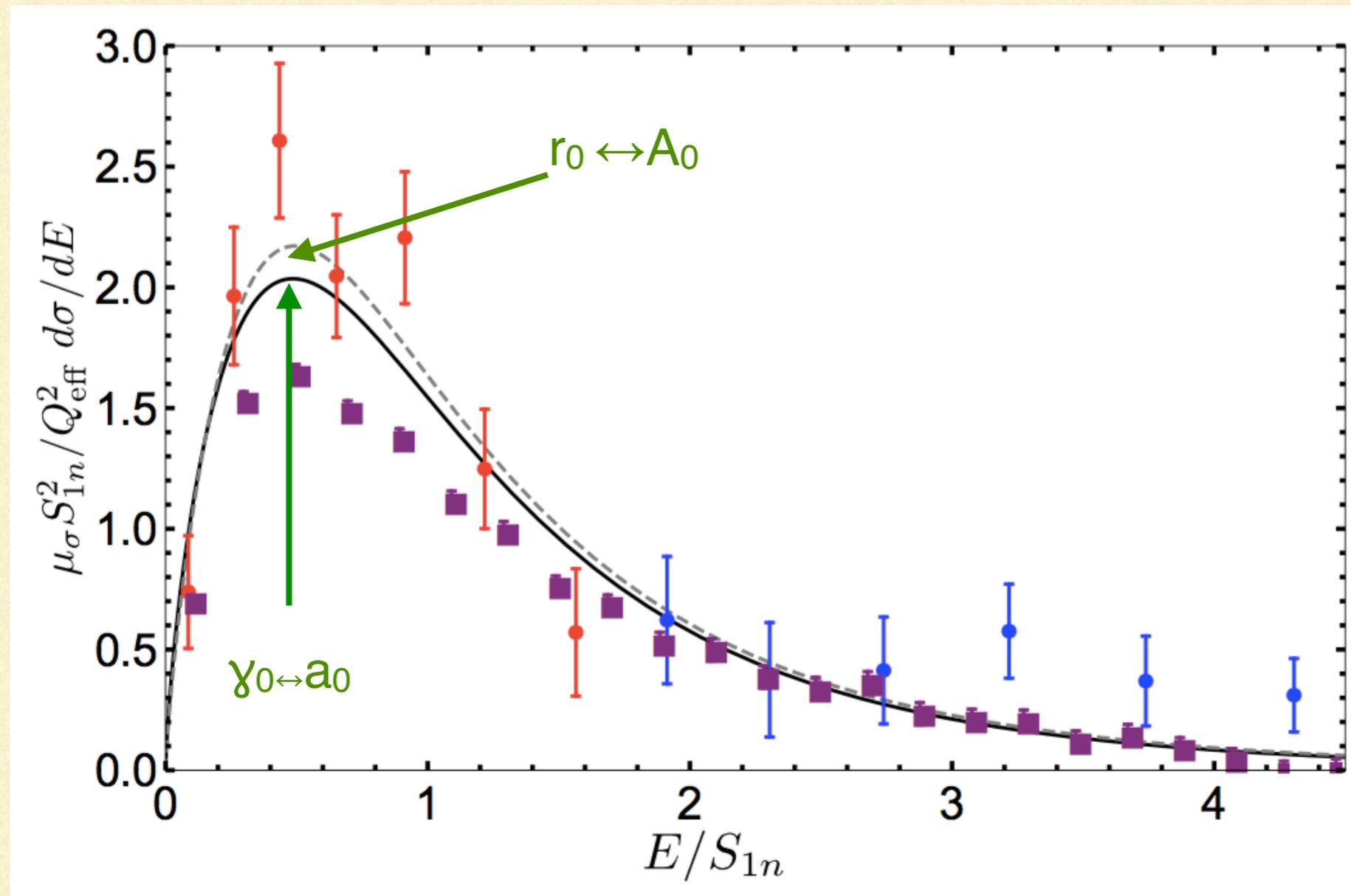
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Determine S-wave  $^{18}\text{C}$ -n scattering parameters  $\Leftrightarrow$   $^{19}\text{C}$  ANC from dissociation data.

# Results



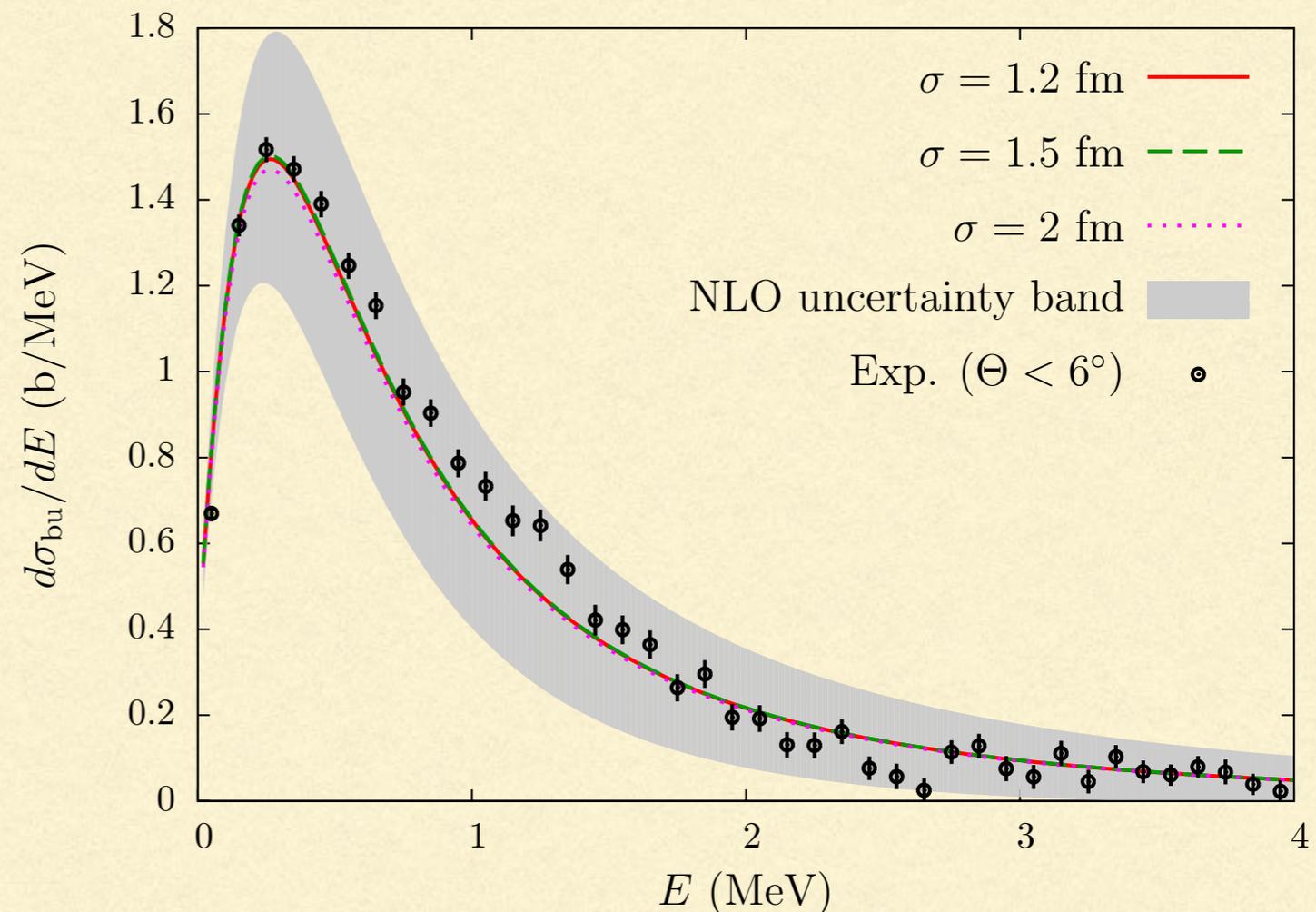
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■ For  $^{19}\text{C}$ :  
 $a = (7.75 \pm 0.35(\text{stat.}) \pm 0.3(\text{EFT})) \text{ fm};$   
 $r_0 = (2.6_{-0.9}^{+0.6}(\text{stat.}) \pm 0.1(\text{EFT})) \text{ fm}.$

# *Ab initio* → Halo EFT → Reaction theory

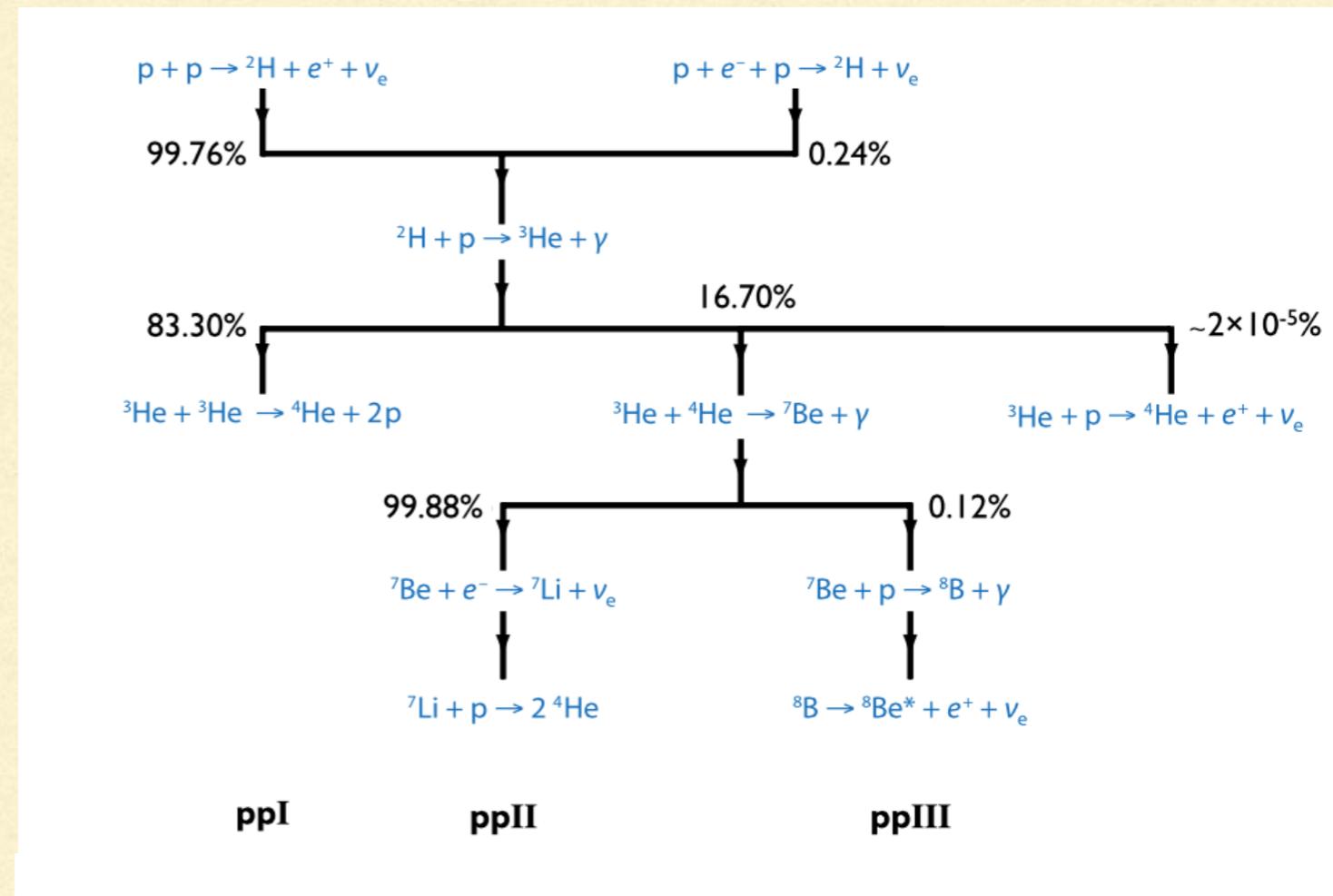
- $^{11}\text{Be}$  is a halo nucleus: last neutron only bound to  $^{10}\text{Be}$  by 503 keV. Has a p-wave halo state with  $S_{1n}=184$  keV.
- Model Coulomb dissociation of  $^{11}\text{Be}$  via sophisticated “Dynamical Eikonal Approximation”: includes nuclear and Coulomb  $^{208}\text{Pb}$ - $^{10}\text{Be}$ -n potentials
- Use Halo EFT to identify important  $^{10}\text{Be}$ -n inputs for reaction-theory calculation: s- and p-wave phase shifts
- Take those from *ab initio* calculation of Calci et al. based on modern nuclear forces and NCSMC (PRL **117**, 242501)

Capel, DP, Hammer, Phys. Rev. C 98, 034610 (2018)  
Data: Fukuda et al., Phys. Rev. C 70, 054606 (2004).



No dependence on interior of  $^{10}\text{Be}$ -n potential used

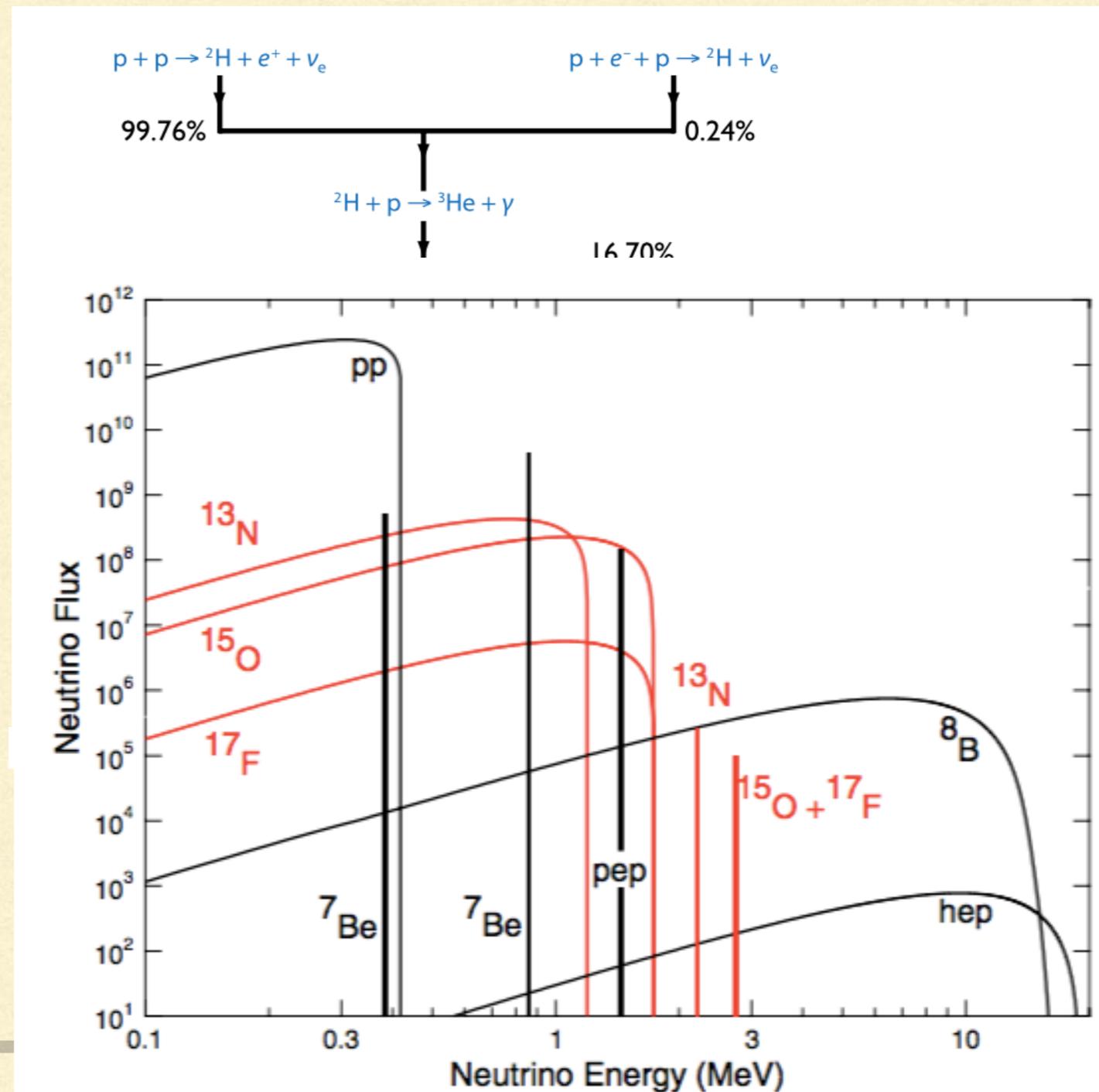
# Why is ${}^7\text{Be}(p,\gamma)$ important?



# Why is ${}^7\text{Be}(p,\gamma)$ important?

- Part of pp chain (ppIII)
- Key for predictions flux of solar neutrinos, especially high-energy ( ${}^8\text{B}$ ) neutrinos
- Accurate knowledge of  ${}^7\text{Be}(p,\gamma)$  needed for inferences from solar-neutrino flux regarding chemical composition of Sun  $\rightarrow$  solar-system formation history
- $S(0) = 20.8 \pm 0.7 \pm 1.4 \text{ eV b}$

“SFII”: Adelberger et al. (2010)



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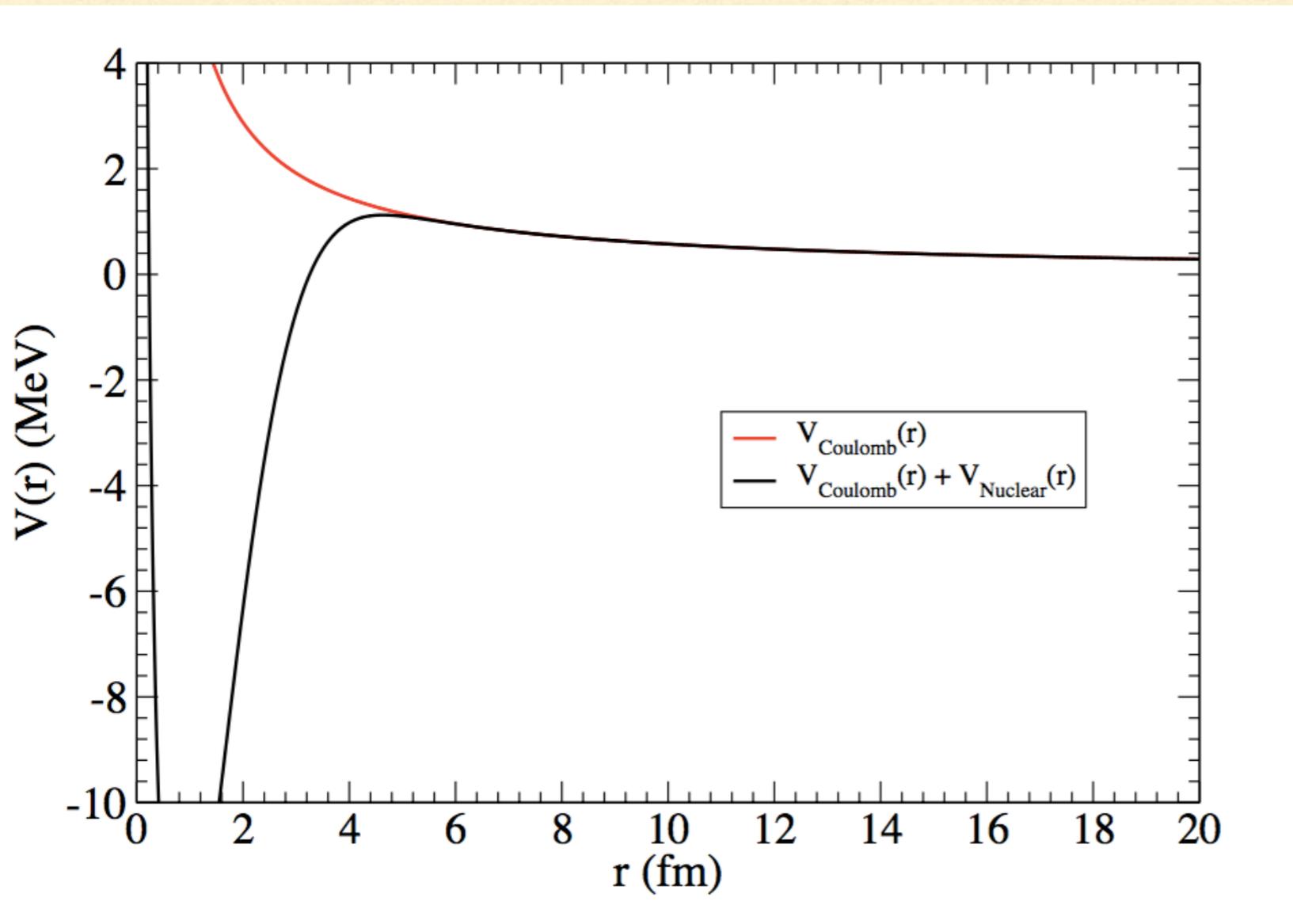
# This is an extrapolation problem

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Thermonuclear  
reaction rate  $\propto \langle v\sigma \rangle \propto \int_0^\infty dE \exp\left(-\frac{E}{k_B T}\right) E \sigma(E)$

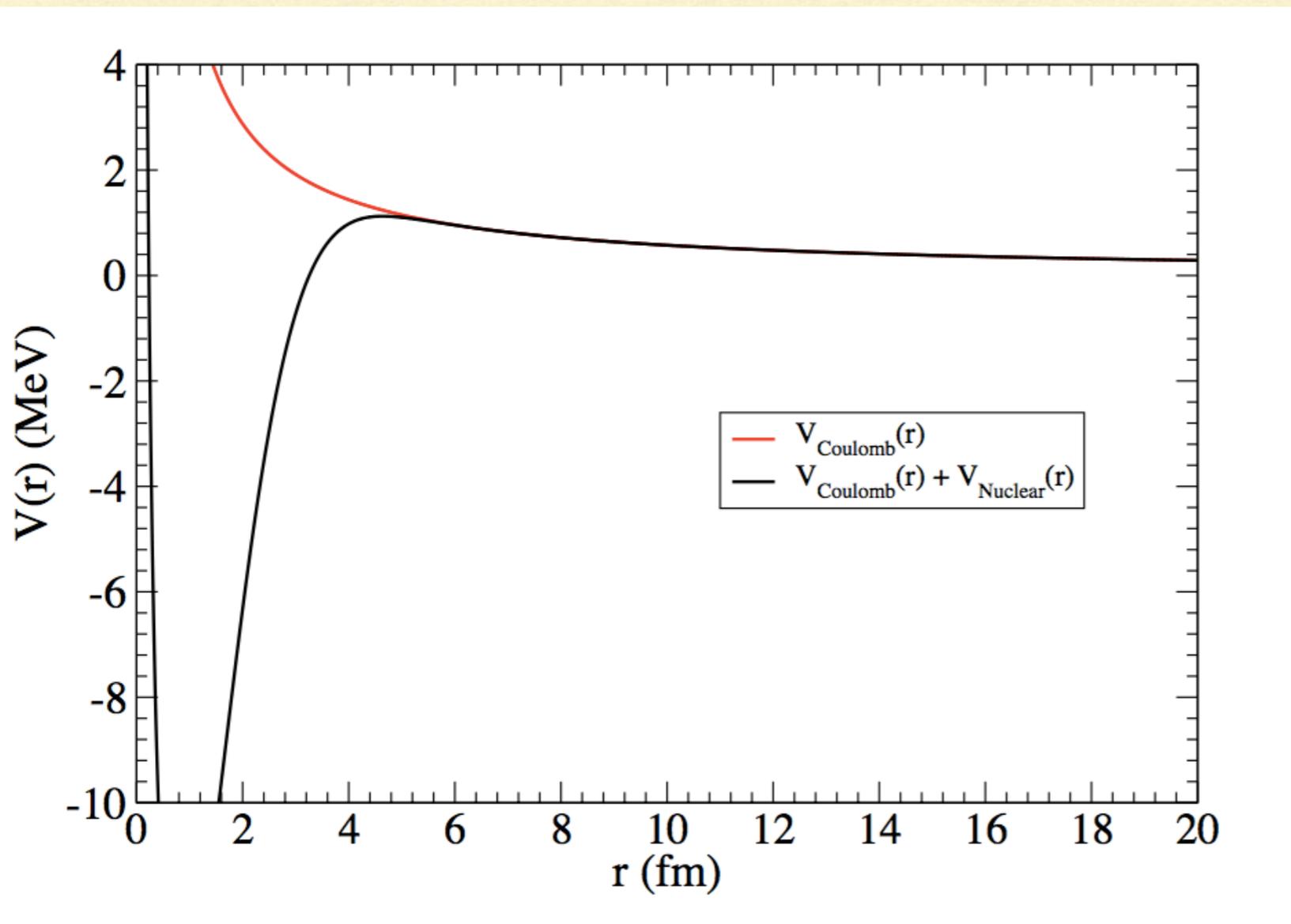
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Thermonuclear reaction rate  $\propto \langle v\sigma \rangle \propto \int_0^\infty dE \exp\left(-\frac{E}{k_B T}\right) E \sigma(E)$



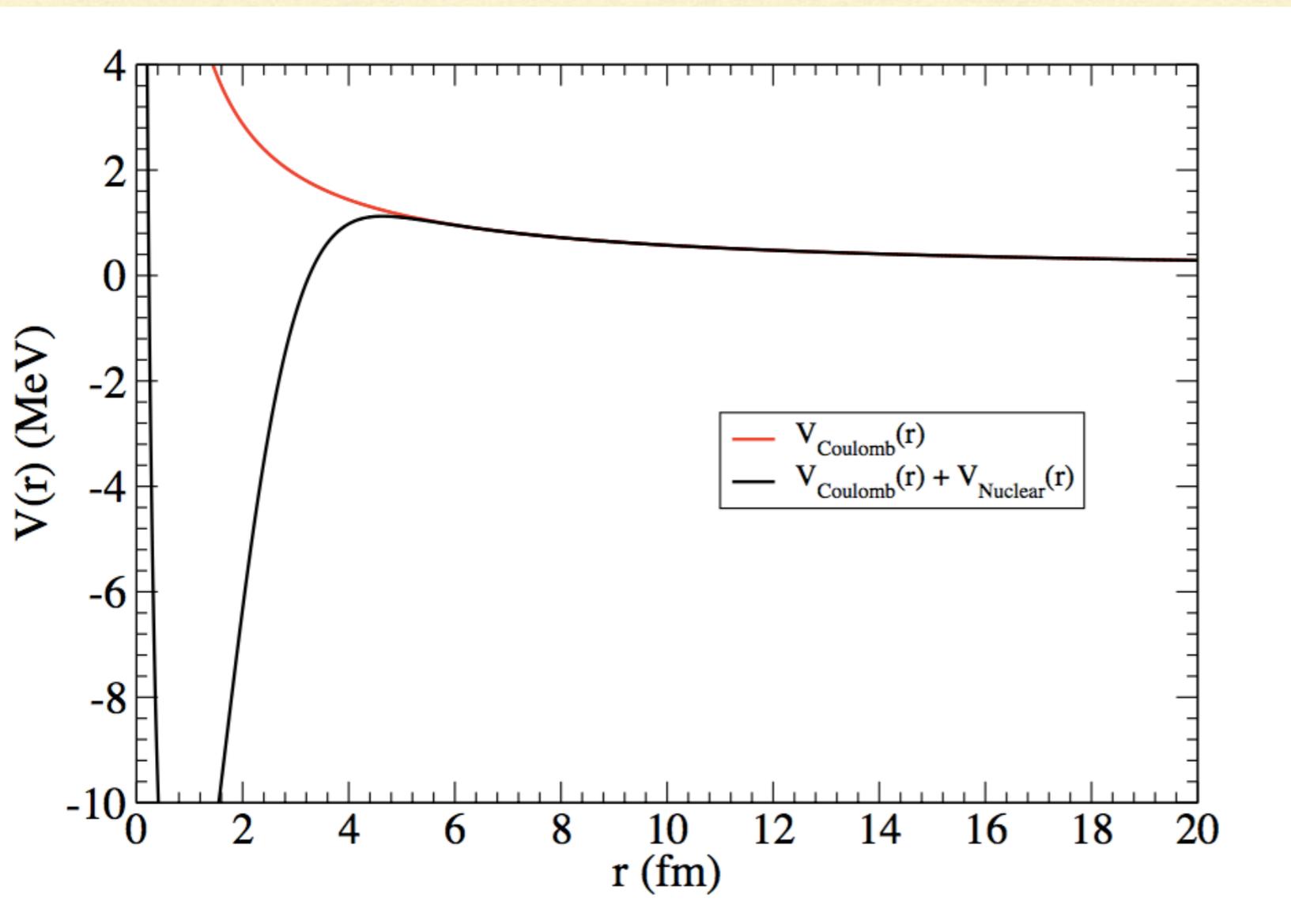
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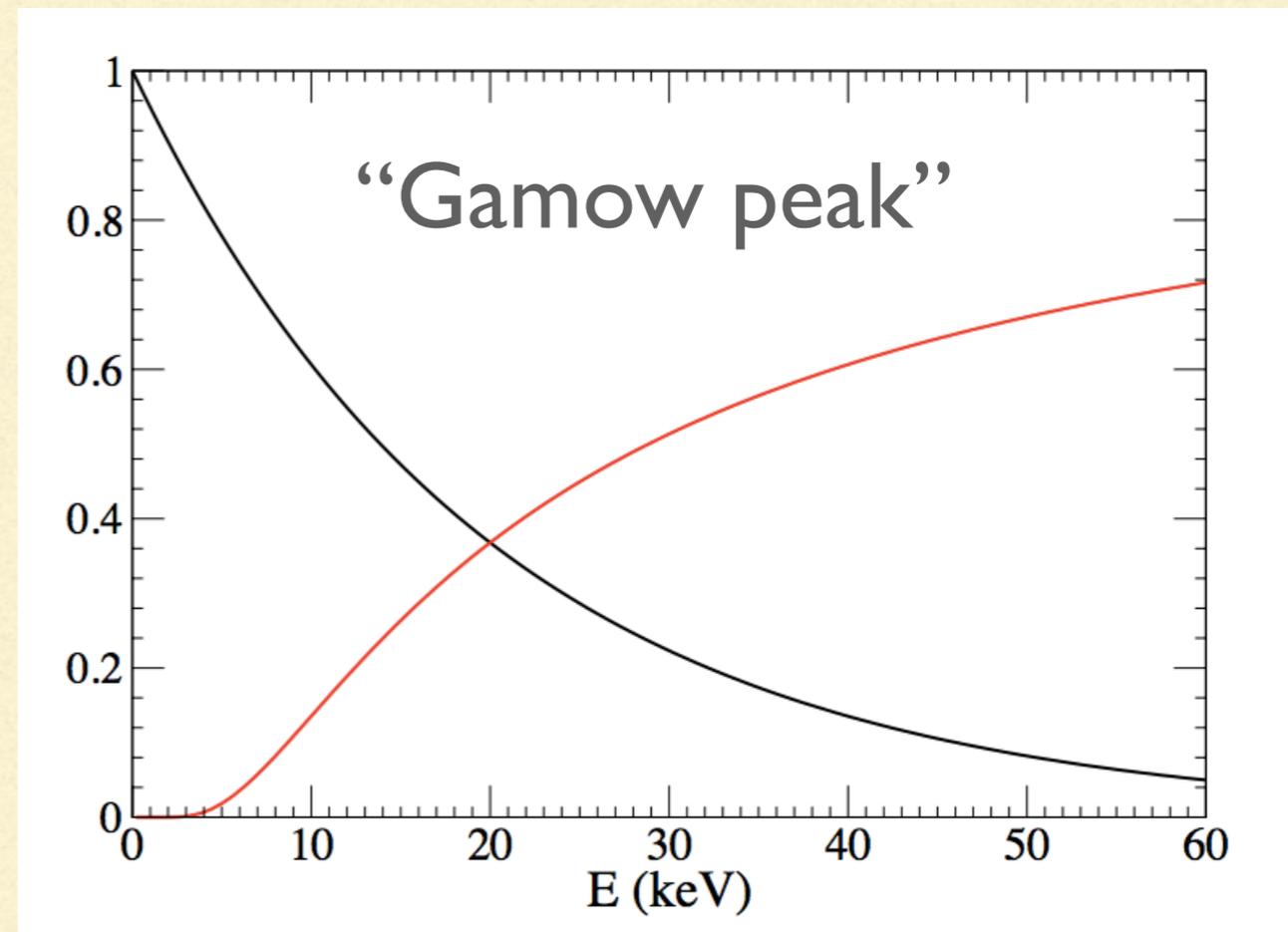
$$\sigma(E) = \frac{S(E)}{E} \exp\left(-\pi Z_1 Z_2 \alpha_{\text{em}} \sqrt{\frac{m_R}{2E}}\right)$$



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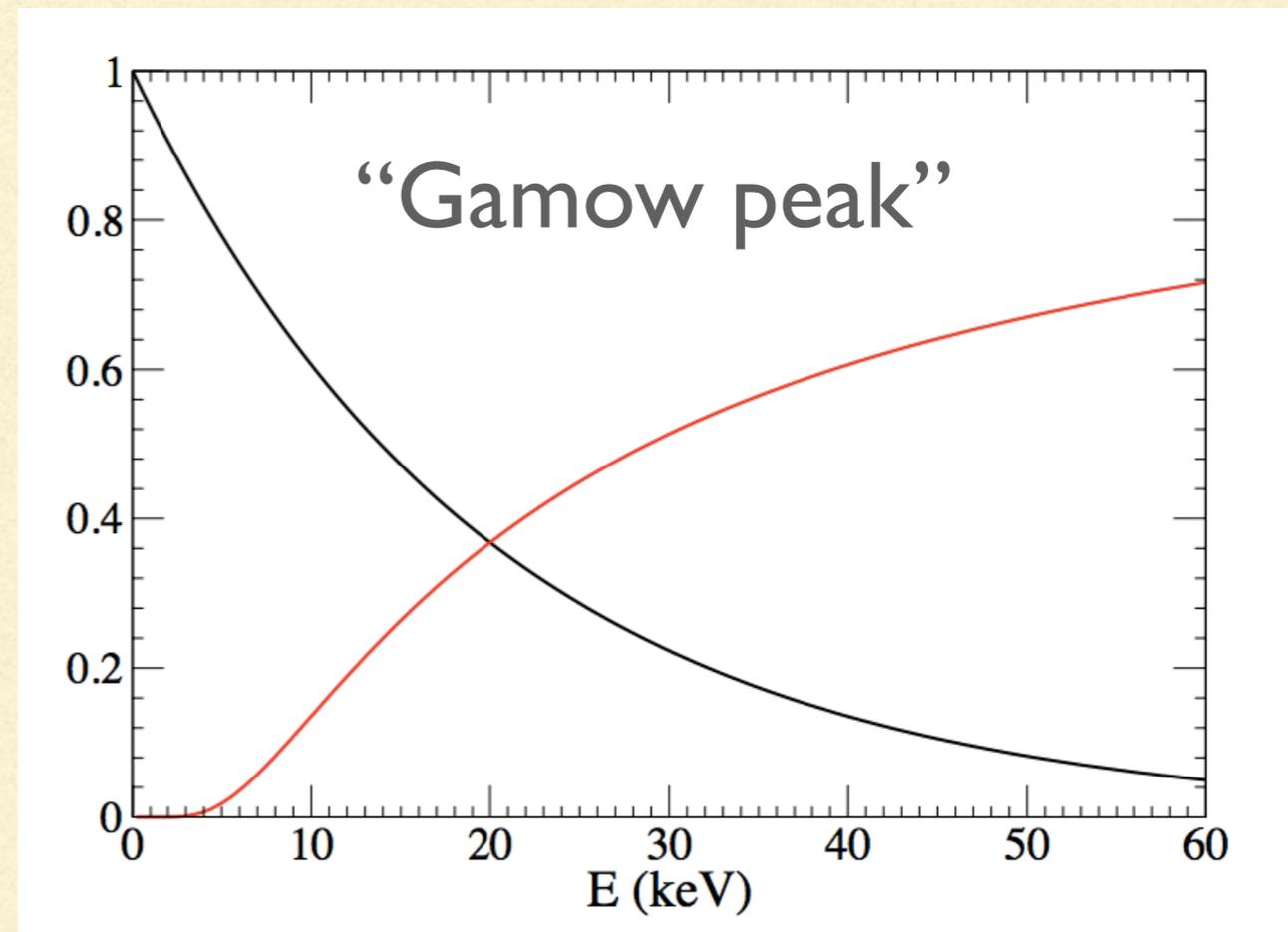


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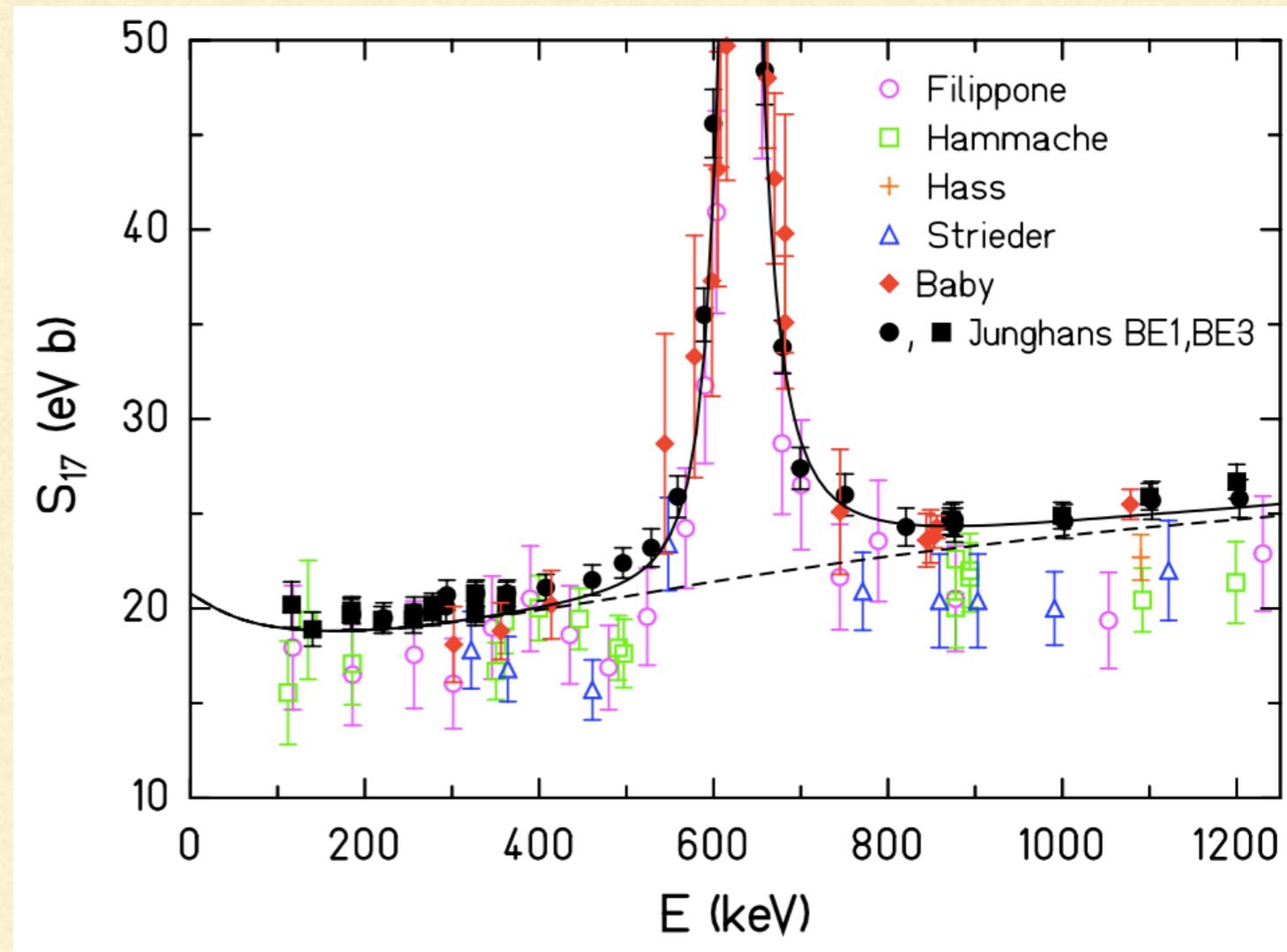
$$\sigma(E) = \frac{S(E)}{E} \exp\left(-\pi Z_1 Z_2 \alpha_{\text{em}} \sqrt{\frac{m_R}{2E}}\right)$$

- EI capture:  ${}^7\text{Be} + \text{p} \rightarrow {}^8\text{B} + \gamma$
- Energies of relevance 20 keV



# Status as in “Solar Fusion II”

- Energies of relevance  $\approx 20$  keV
- There dominated by  ${}^7\text{Be}$ -p separations  $\sim 10$ s of fm
- Below narrow  $1^+$  resonance proceeds via s- and d-wave direct capture
- Energy dependence due to interplay of bound-state properties, Coulomb, strong ISI



- SF II central value used energy-dependence from Descouvemont's ab initio eight-body calculation. Errors from consideration of energy-dependence in a variety of “reasonable models”

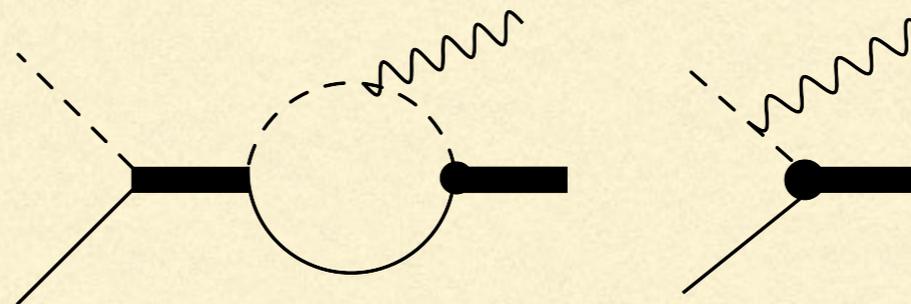
# Capture to p-wave halo in EFT

Hammer & DP, NPA (2011)

- At LO: p-wave In halo described solely by its ANC and binding energy

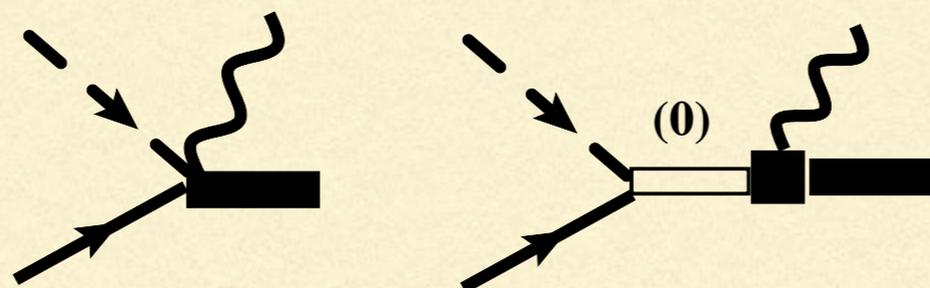
$$u_1(r) = A_1 \exp(-\gamma_1 r) \left( 1 + \frac{1}{\gamma_1 r} \right)$$

- Capture to the p-wave state proceeds via the one-body E1 operator: “external direct capture”



$$E1 \propto \int_0^\infty dr u_0(r) r u_1(r); \quad u_0(r) = 1 - \frac{r}{a}$$

- NLO: piece of the amplitude representing capture at short distances, represented by a contact operator  $\Rightarrow$  there is an LEC that must be fit



# NLO for ${}^7\text{Be}(p,\gamma)$

Zhang, Nollett, Phillips, PRC (2014)

cf. Ryberg, Forsen, Hammer, Platter, EPJA (2014)

Zhang, Nollett, Phillips, PLB (2015); PRC (2018)

- LO calculation: ISI in  $S=2$  &  $S=1$  into p-wave bound state. Scattering wave functions are linear combinations of Coulomb wave functions  $F_0$  and  $G_0$ . Bound state wave function = the appropriate Whittaker function
- We also incorporate a low-lying excited state ( $1/2^-$ ) in  ${}^7\text{Be}$
- NLO: piece of the amplitude representing capture at short distances, represented by a contact operator  $\Rightarrow$  there is an LEC that must be fit

$$S(E) = f(E) \sum_s C_s^2 \left[ |\mathcal{S}_{\text{EC}}(E; \delta_s(E)) + \bar{L}_s \mathcal{S}_{\text{SD}}(E; \delta_s(E)) + \epsilon_s \mathcal{S}_{\text{CX}}(E; \delta_s(E))|^2 + |\mathcal{D}(E)|^2 \right]$$

- ANCs in  ${}^5\text{P}_2$  and  ${}^3\text{P}_2$ :  $A_{5\text{P}_2}$  and  $A_{3\text{P}_2}$

Four parameters at LO;  
five more at NLO

- Scattering lengths and effective ranges in both  ${}^5\text{S}_2$  and  ${}^3\text{S}_1$ :  $a_2, r_2$  and  $a_1, r_1$
- Core excitation: determined by ratio of  ${}^8\text{B}$  couplings of  ${}^7\text{Be}^*p$  and  ${}^7\text{Be}-p$  states:  $\epsilon_1$
- LECs associated with contact interaction, one each for  $S=1$  and  $S=2$ :  $L_1$  and  $L_2$

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# Extrapolation to zero energy

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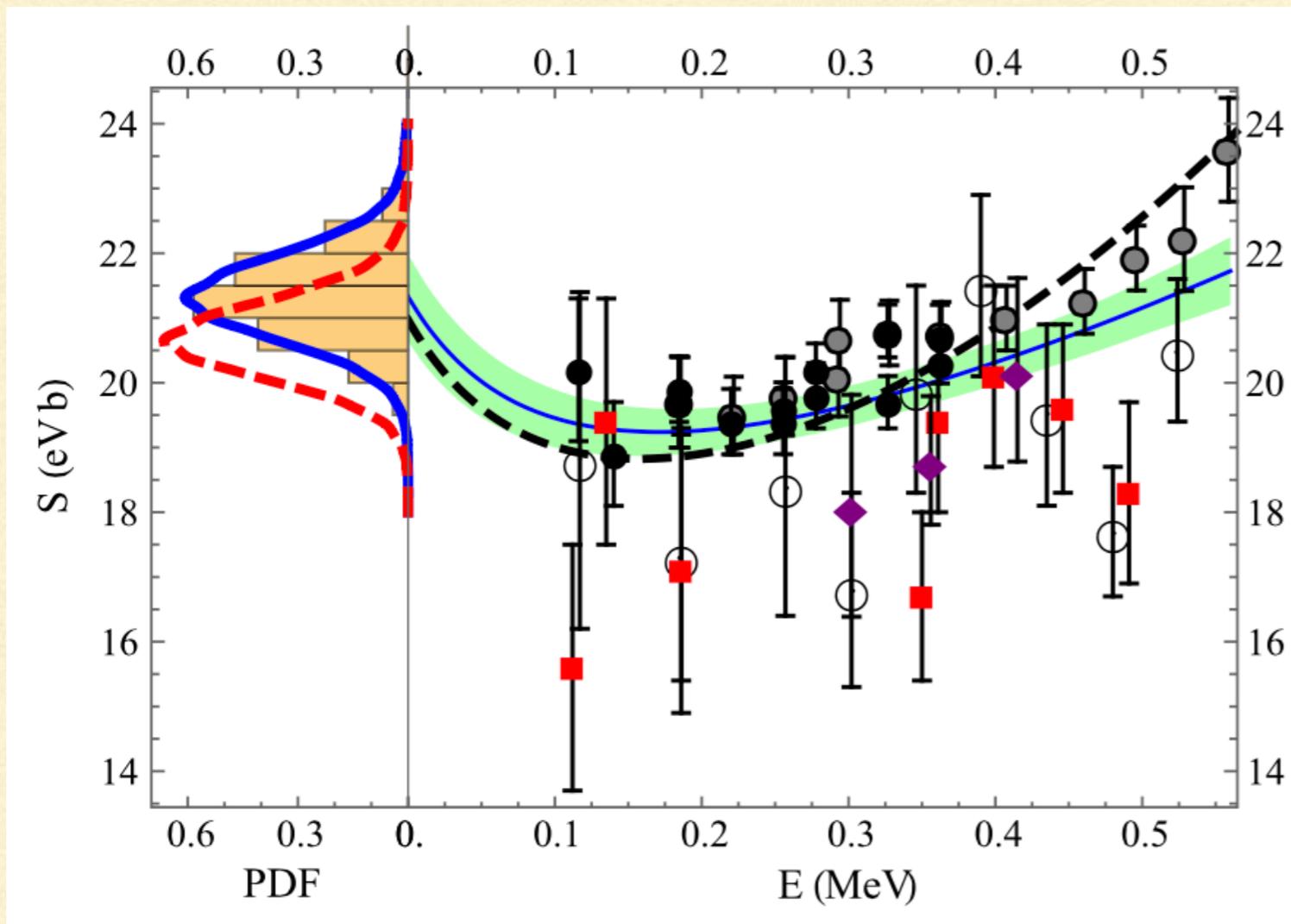
Zhang, Nollett, DP, PLB, 2015; arXiv:1708.04017

$$\text{pr}(\bar{F}|D;T;I) = \int \text{pr}(\vec{g}, \{\xi_i\}|D;T;I) \delta(\bar{F} - F(\vec{g})) d\xi_1 \dots d\xi_5 d\vec{g}$$

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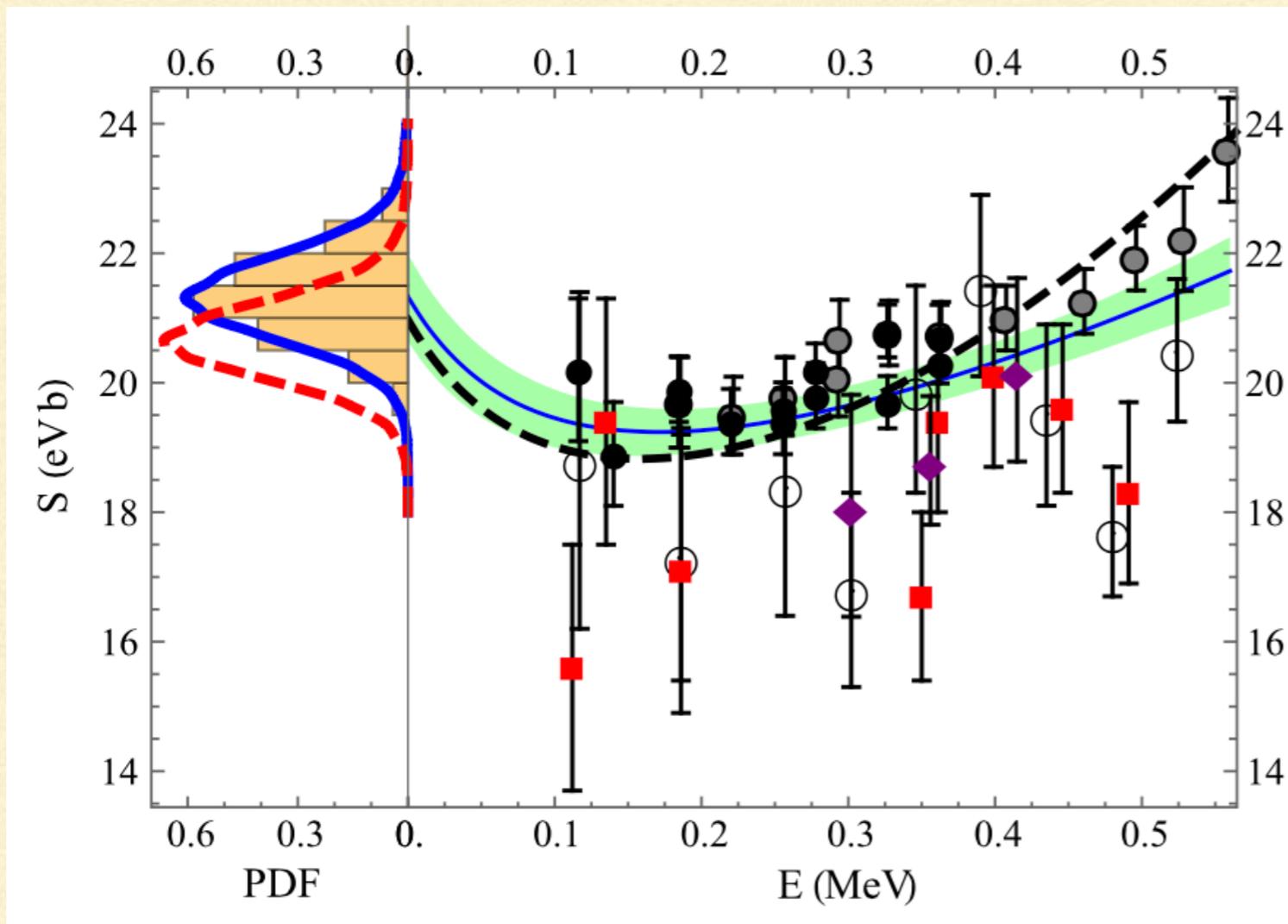
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$$S(0) = 21.33^{+0.66}_{-0.69} \text{ eV b}$$

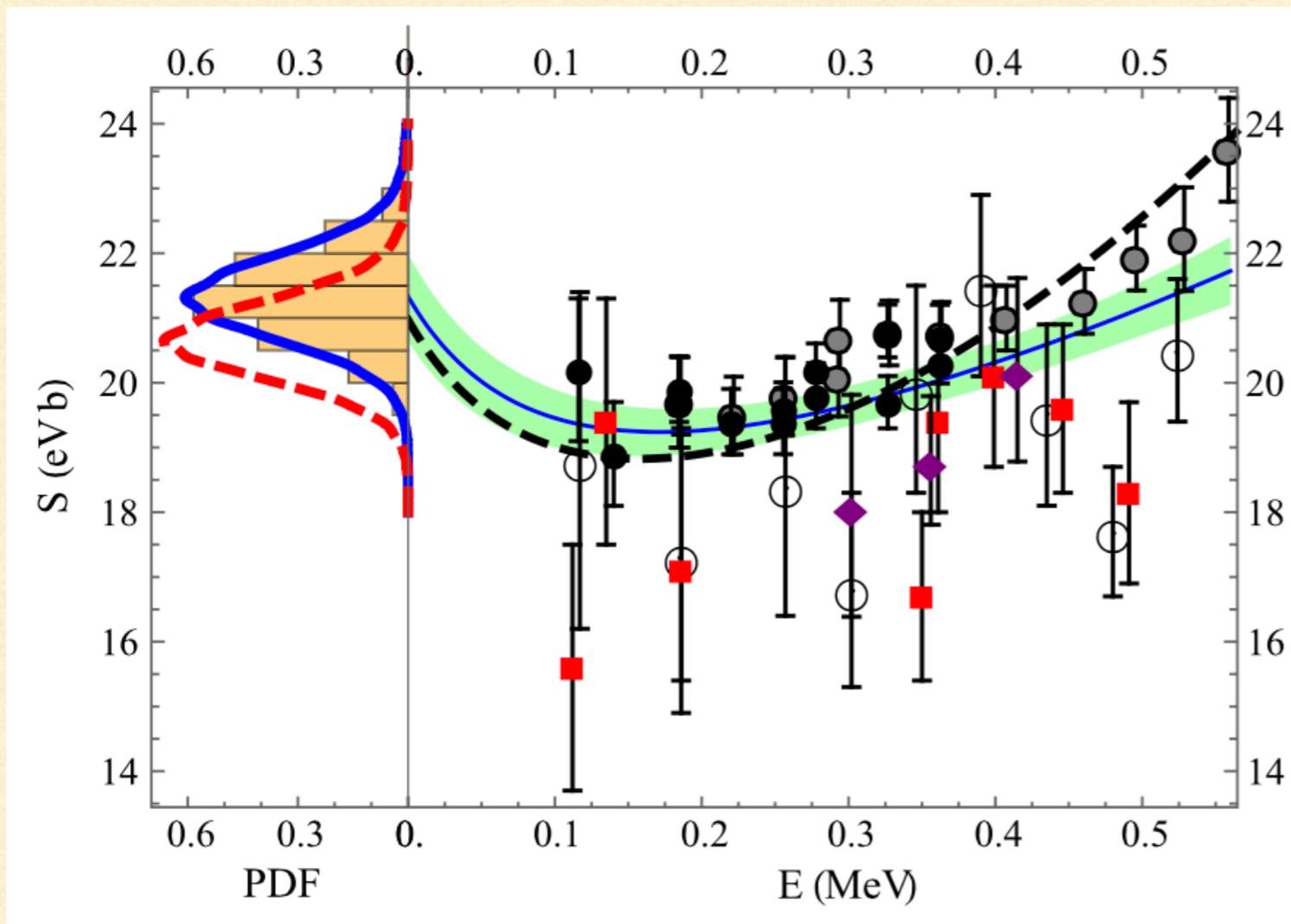
No  $N^2\text{LO}$  corrections

Also assessed impact of  $N^3\text{LO}$  contact operator

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No N<sup>2</sup>LO corrections

Also assessed impact of  
N<sup>3</sup>LO contact operator

Some remaining  
uncertainty due to <sup>8</sup>B S<sub>1p</sub>

Uncertainty reduced by factor of two:  
model selection

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# Ongoing work along these lines

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- Simultaneous fit to  ${}^7\text{Be}+p$  scattering data: requires inclusion of resonances (TRIUMF experiment) Poudel, Zhang, DP
  - Same techniques applied to  ${}^3\text{He}({}^4\text{He},\gamma)$  Vaghani, Higa, Rupak Zhang, Nollett, DP
  - Coulomb dissociation: better reaction theory and connection to *ab initio* structure Capel, Hammer, DP
  - Rotational states as explicit degrees of freedom Coello Pérez, Papenbrock Alnamlah, Coello Perez, DP
  - Gaussian process models for EFT truncation errors Melendez, Furnstahl, DP, Wesolowski
  - $\chi\text{EFT}$  truncation errors in nuclear & neutron matter Drischler, Melendez, Furnstahl, DP
  - Parameter estimation for 3NFs in  $\chi\text{EFT}$
-

# One thing is certain....

The purpose of this Editorial is to discuss the importance of including uncertainty estimates in papers involving theoretical calculations of physical quantities.

It is not unusual for manuscripts on theoretical work to be submitted without uncertainty estimates for numerical results. In contrast, papers presenting the results of laboratory measurements would usually not be considered acceptable for publication in *Physical Review A* without a detailed discussion of the uncertainties involved in the measurements....

The question is to what extent can the same high standards be applied to papers reporting the results of theoretical calculations....There are many cases where it is indeed not practical to give a meaningful error estimate for a theoretical calculation....However, there is a broad class of papers where estimates of theoretical uncertainties can and should be made.

Papers presenting the results of theoretical calculations are expected to include uncertainty estimates for the calculations whenever practicable, and especially under the following circumstances:

1. If the authors claim high accuracy, or improvements on the accuracy of previous work.
2. If the primary motivation for the paper is to make comparisons with present or future high precision experimental measurements.
3. If the primary motivation is to provide interpolations or extrapolations of known experimental measurements.

*Physical Review A Editorial, 29 April 2011*

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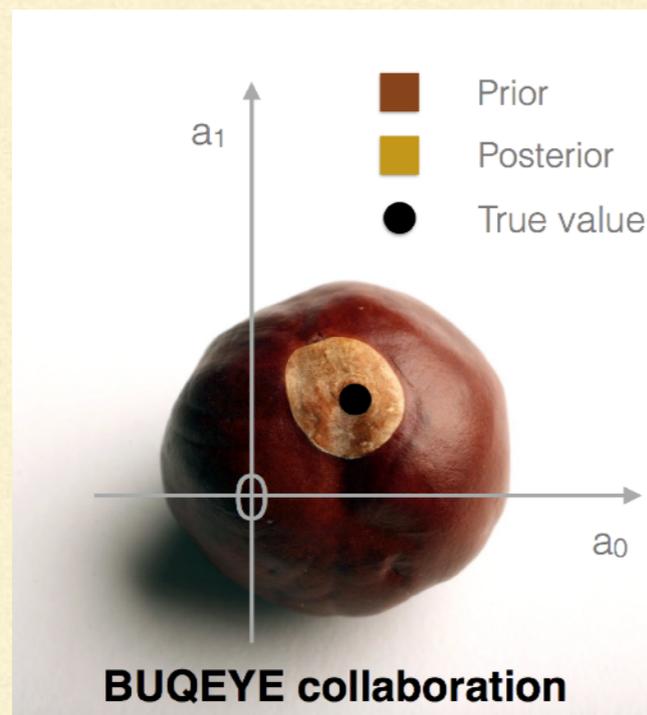
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## Bayesian Uncertainty Quantification: Errors for Your EFT

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# Theorists Anonymous

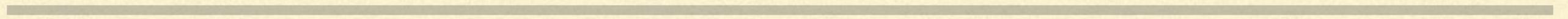
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# Theorists Anonymous

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- Admit that you have a problem: your theory has uncertainties



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  - Make amends for those mistakes
  - Help others who must deal with the same issues
-

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# References

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  - “Bayesian parameter estimation for effective field theories”, S. Wesolowski, N. Klco, R. J. Furnstahl, D. R. Phillips, A. Thapaliya, *J. Phys. G* **43**, 074001 (2016); “Exploring Bayesian Parameter Estimation for ChiEFT using NN phase shifts”, S. Wesolowski, R. J. Furnstahl, J. Melendez, D. R. Phillips, arXiv:1808.08211
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  - “Effective field theory for halo nuclei”, H.-W. Hammer, C. Ji, D. R. Phillips, *Topical Review for J. Phys. G* **44**, 103002 (2017).
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# Backup Slides

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# A Generic EFT

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$$g(x) = \sum_{i=0}^k \mathcal{A}_i(x) x^i$$

$$x = \frac{p}{\Lambda_b}$$

---

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- Suppose we are interested in a quantity as a function of a momentum,  $p$ , that is small compared to some high scale,  $\Lambda_b$ .

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$$\mathcal{A}_i(x) = a_i(\mu) + f_i(x, \mu) \quad a_i, f_i = \mathcal{O}(1) \text{ for } \mu \sim \Lambda_b, x \sim 1$$

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- $f_i(x, \mu)$  is a calculable function, that encodes IR physics at order  $i$
  - $a_i$  is a low-energy constant (LEC): encodes UV physics at order  $i$ . Must be fit to data
  - Complications: multiple light scales, multiple functions at a given order, skipped orders, .....
-

---

# Bayes $\rightarrow$ Result

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- Bayes theorem: 
$$\begin{aligned}\text{pr}(\bar{c}|c_0, c_1, \dots, c_k) &= \frac{\text{pr}(c_0, c_1, \dots, c_k|\bar{c})\text{pr}(\bar{c})}{\text{pr}(c_0, c_1, \dots, c_k)} \\ &= \mathcal{N}\text{pr}(\bar{c})\prod_{n=0}^k\text{pr}(c_n|\bar{c})\end{aligned}$$

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- Bayes theorem: 
$$\text{pr}(\bar{c} | c_0, c_1, \dots, c_k) = \frac{\text{pr}(c_0, c_1, \dots, c_k | \bar{c}) \text{pr}(\bar{c})}{\text{pr}(c_0, c_1, \dots, c_k)}$$
$$= \mathcal{N} \text{pr}(\bar{c}) \prod_{n=0}^k \text{pr}(c_n | \bar{c})$$
- Marginalization:

$$\text{pr}(c_{k+1} | c_0, c_1, \dots, c_k) = \int_0^\infty d\bar{c} \text{pr}(c_{k+1} | \bar{c}) \text{pr}(\bar{c} | c_0, c_1, \dots, c_k)$$

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- This is generic, but the integrals are simple in the case of “Prior A”

$$\text{pr}(\bar{c} | c_0, c_1, \dots, c_k) \propto \begin{cases} 0 & \text{if } \bar{c} < \max\{c_0, \dots, c_k\} \\ 1/\bar{c}^{k+2} & \text{if } \bar{c} > \max\{c_0, \dots, c_k\} \end{cases}$$

$$\text{pr}(c_{k+1} | c_0, c_1, \dots, c_k) \propto \begin{cases} 1 & \text{if } c_{k+1} < c_{\max} \\ \left(\frac{c_{\max}}{c_{k+1}}\right)^{k+2} & \text{if } c_{k+1} > c_{\max} \end{cases}$$

---

# I don't like THAT prior!

---

- Modify Set A to restrict  $\bar{c}$  to a finite range, e.g.  $A_{[0.25,4]}$

- Set B: give  $\bar{c}$  a log-normal prior:  $\text{pr}(\bar{c}) = \frac{1}{\sqrt{2\pi\bar{c}\sigma}} e^{-(\log \bar{c})^2 / 2\sigma^2}$

- Set C:  $\text{pr}(c_n | \bar{c}) = \frac{1}{\sqrt{2\pi\bar{c}}} e^{-c_n^2 / 2\bar{c}^2}$ ;  $\text{pr}(\bar{c}) \propto \frac{1}{\bar{c}} \theta(\bar{c} - \bar{c}_{<}) \theta(\bar{c}_{>} - \bar{c})$

- Same formulas as before can be invoked. Now numerical.

$$\text{pr}(c_{k+1} | c_0, c_1, \dots, c_k) = \int_0^\infty d\bar{c} \text{pr}(c_{k+1} | \bar{c}) \text{pr}(\bar{c} | c_0, c_1, \dots, c_k)$$

$$\text{pr}(\bar{c} | c_0, c_1, \dots, c_k) = \mathcal{N} \text{pr}(\bar{c}) \prod_{n=0}^k \text{pr}(c_n | \bar{c})$$

- You don't like these? Pick your own and follow the rules...

- First omitted term approximation
-

---

# BREAKDOWN-SCALE INFERENCE

---

- $\Lambda_b$  determines the size of the  $c_n$ 's. Choose it too big, and they'll be too big. Choose it too small, they'll be too small. And progressively so as one moves to higher and higher order.
- We have a theory for  $\text{pr}(c_n|c_0, c_1, \dots, c_k)$ : now use Bayes' theorem to see how (im)probable are the  $c_n$ 's that dimensionful EFT coefficients ( $b_n$ 's) produce for a given  $\Lambda_b$ .

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At one energy:

$$\text{pr}(\Lambda_b|b_2, \dots, b_k) \propto \frac{1}{\Lambda_b} \left( \frac{\Lambda_b^{k+2}}{(k+1)\langle b^2 \rangle} \right)^{\frac{k-1}{2}}$$

(NLO:  $k=2$ , NNLO:  $k=3$ , N<sup>3</sup>LO:  $k=4$ , etc.)

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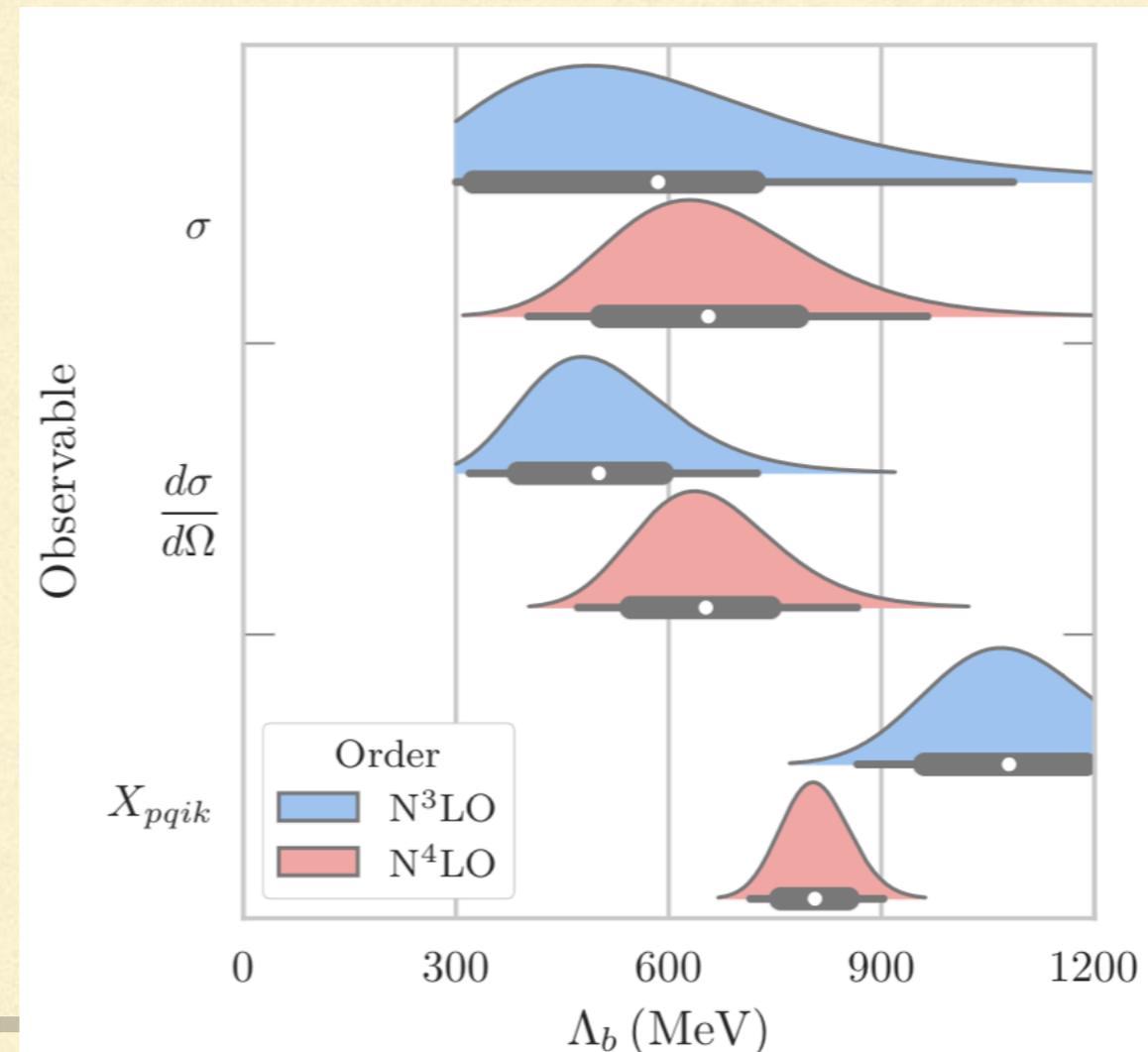
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Using 5 energies (and 2 angles):



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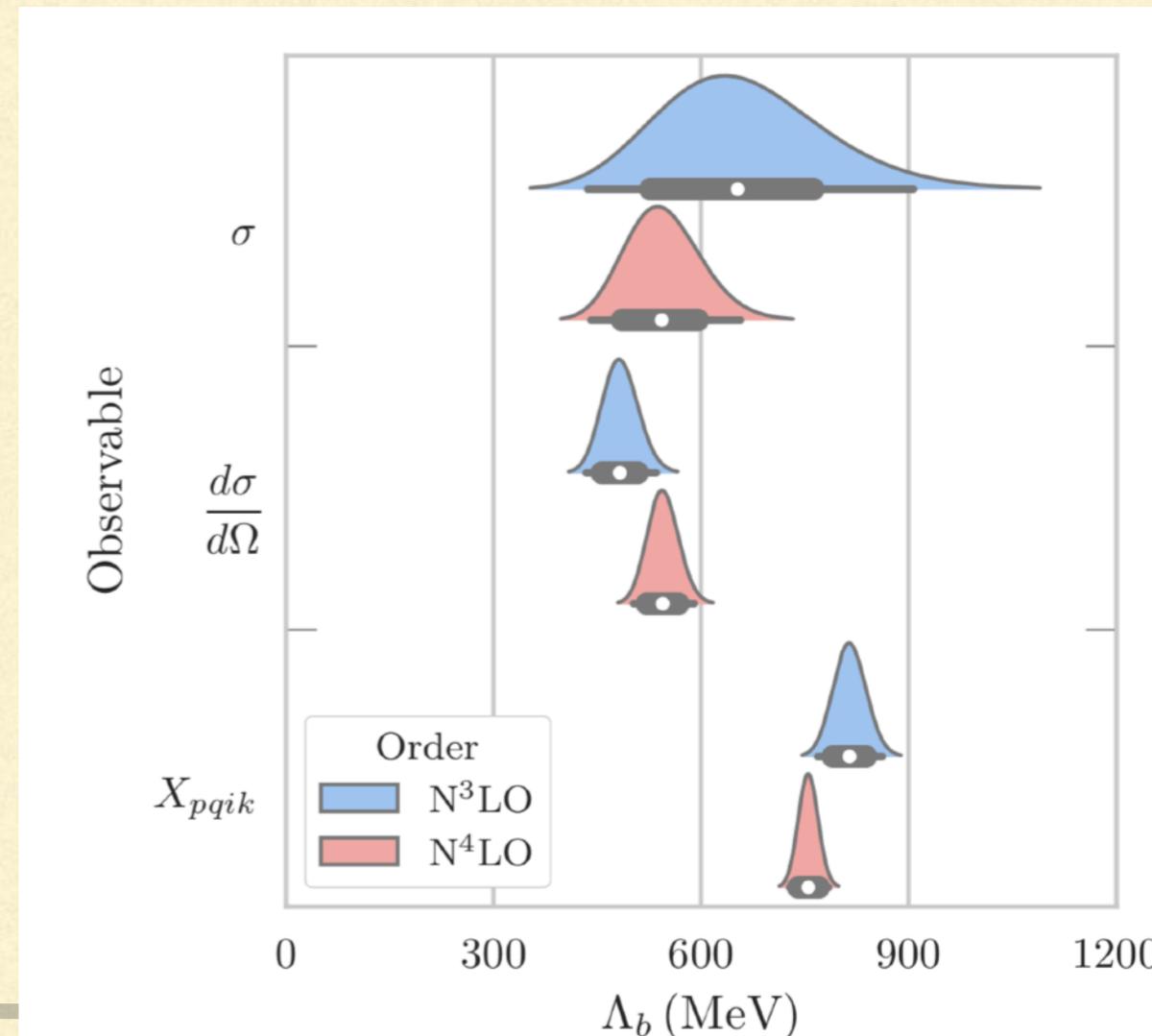
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(NLO:  $k=2$ , NNLO:  $k=3$ , N<sup>3</sup>LO:  $k=4$ , etc.)

Using 17 energies (and 7 angles):



# BREAKDOWN-SCALE INFERENCE

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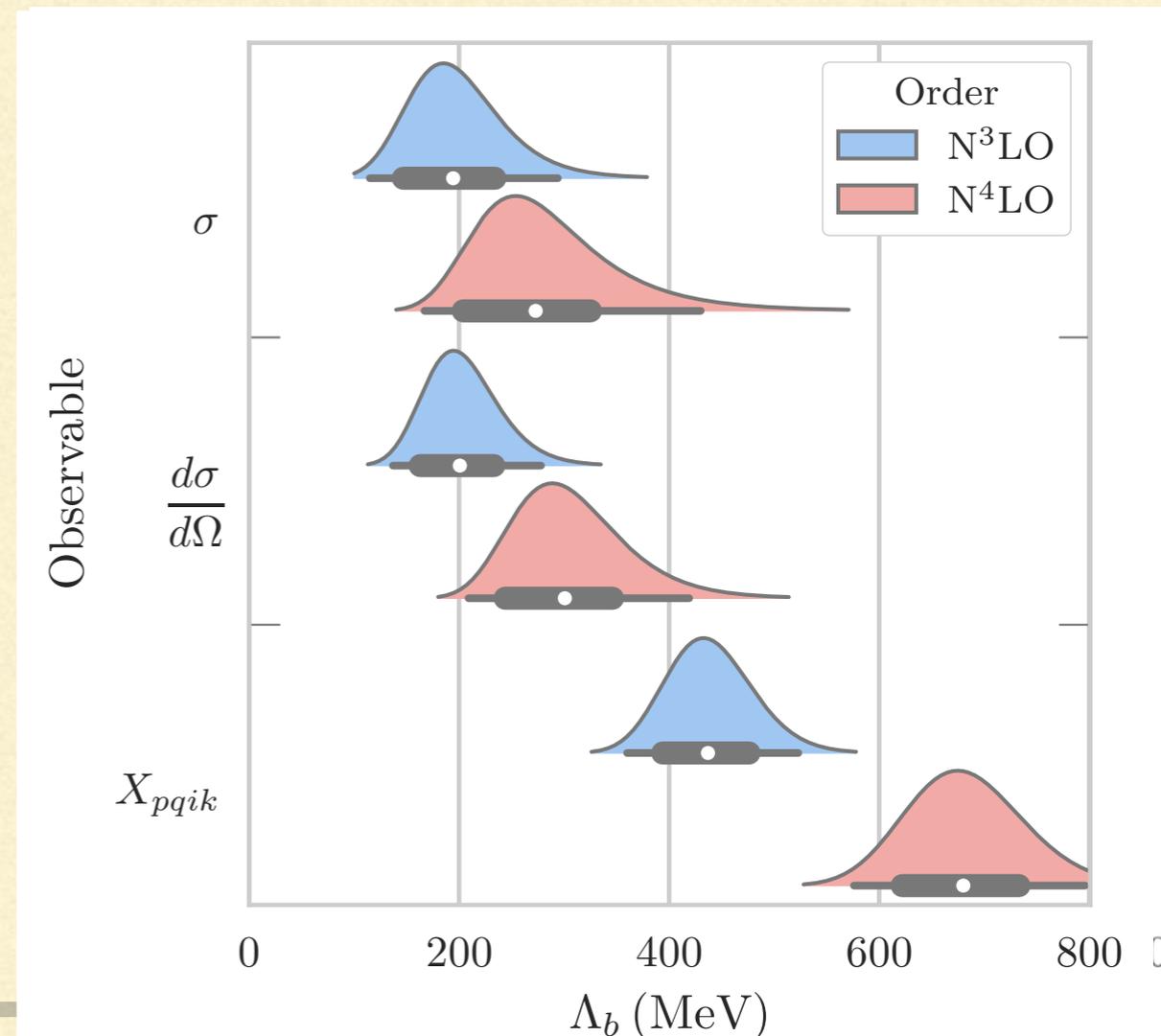
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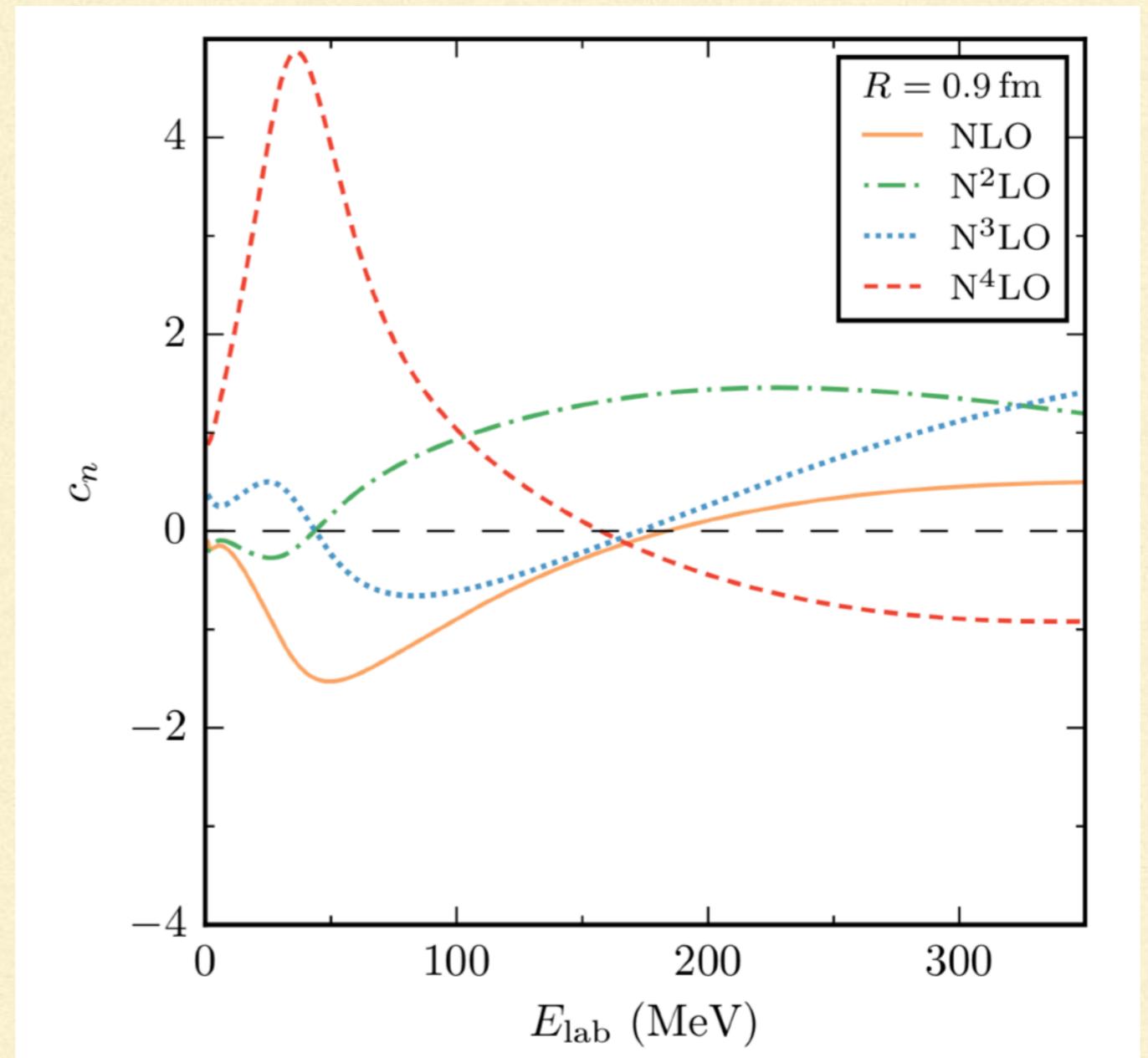
Using 17 energies (and 7 angles):

$R=1.2$  fm



# FUNCTIONAL DATA

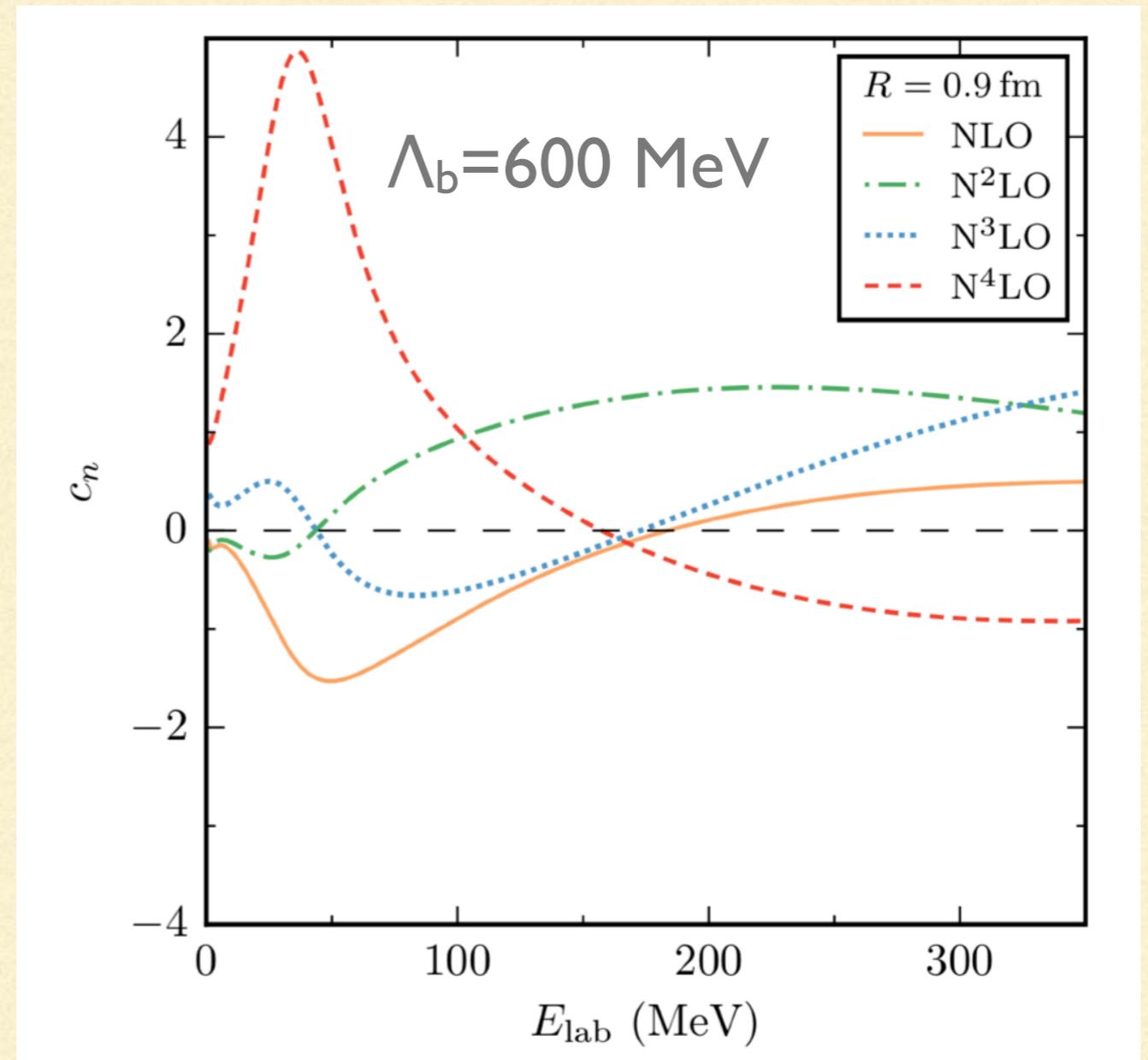
- But we don't have 119 independent data points
- We have a function for each observable at each order
- Can we understand the properties of these functions, so we can do  $\Lambda_b$  inference and compute success ratios rigorously?



$$\sigma(E) = \sigma_0(E) \left[ 1 + c_2(E)x^2 + c_3(E)x^3 + c_4(E)x^4 + c_5(E)x^5 \right]$$

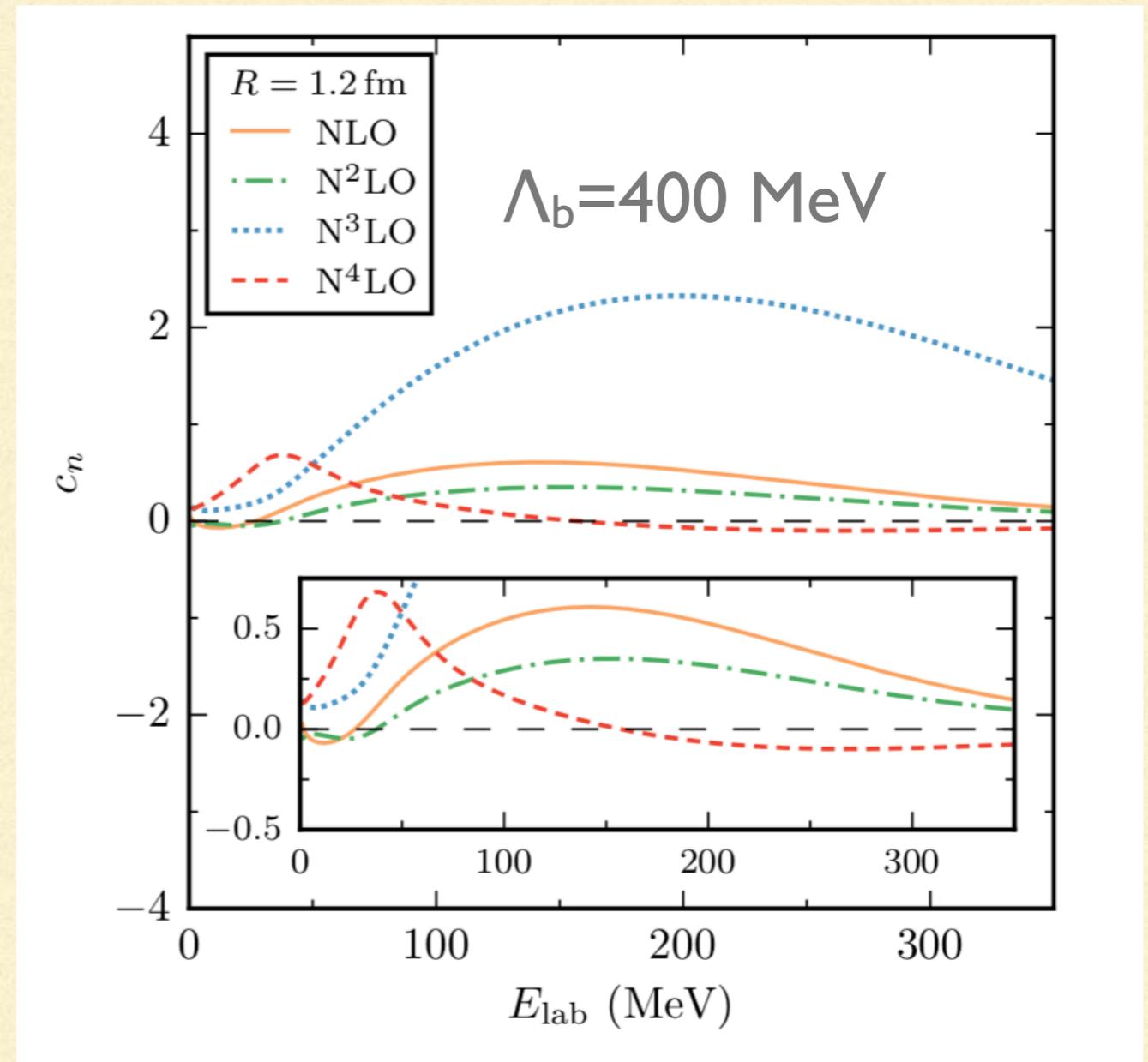
# OBSERVATIONS AND QUESTIONS

- $c_n$ 's do not grow or shrink with  $n$ : good  $\Lambda_b$  choice
- Bounded functions, mostly between -2 and 2
- Each “takes a turn” at being largest
- Not oscillating quickly in this energy range



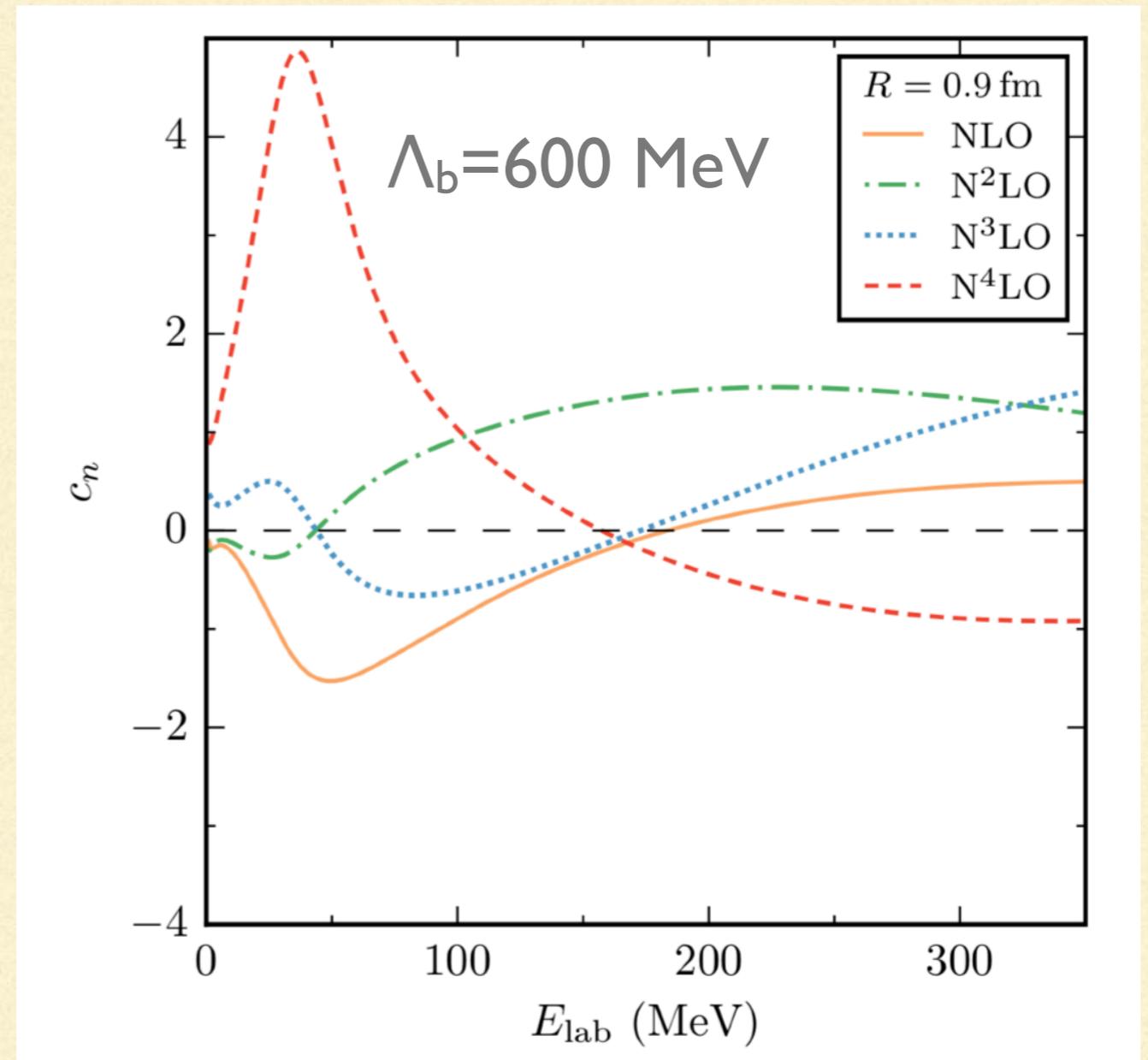
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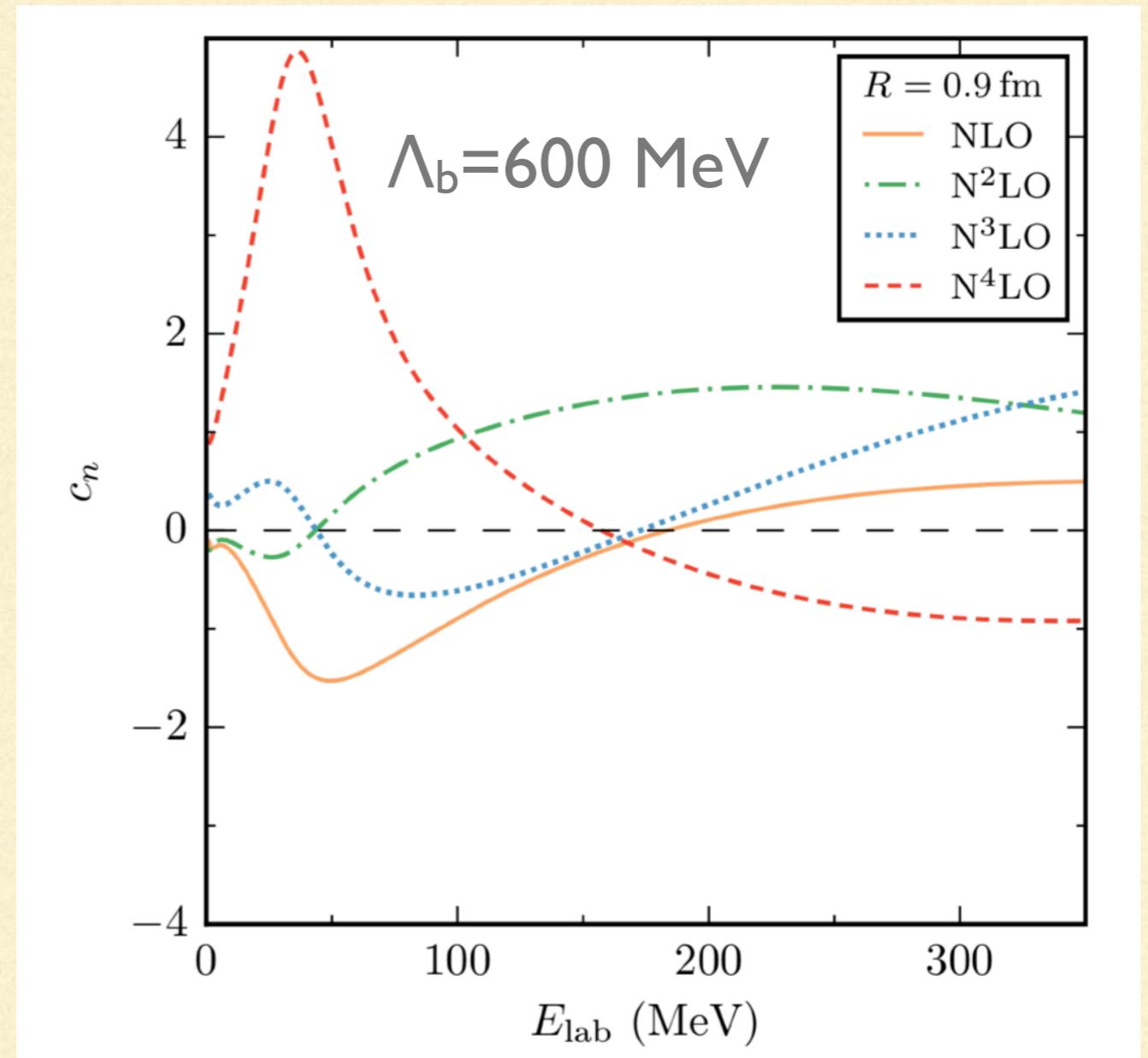


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## Physics questions:

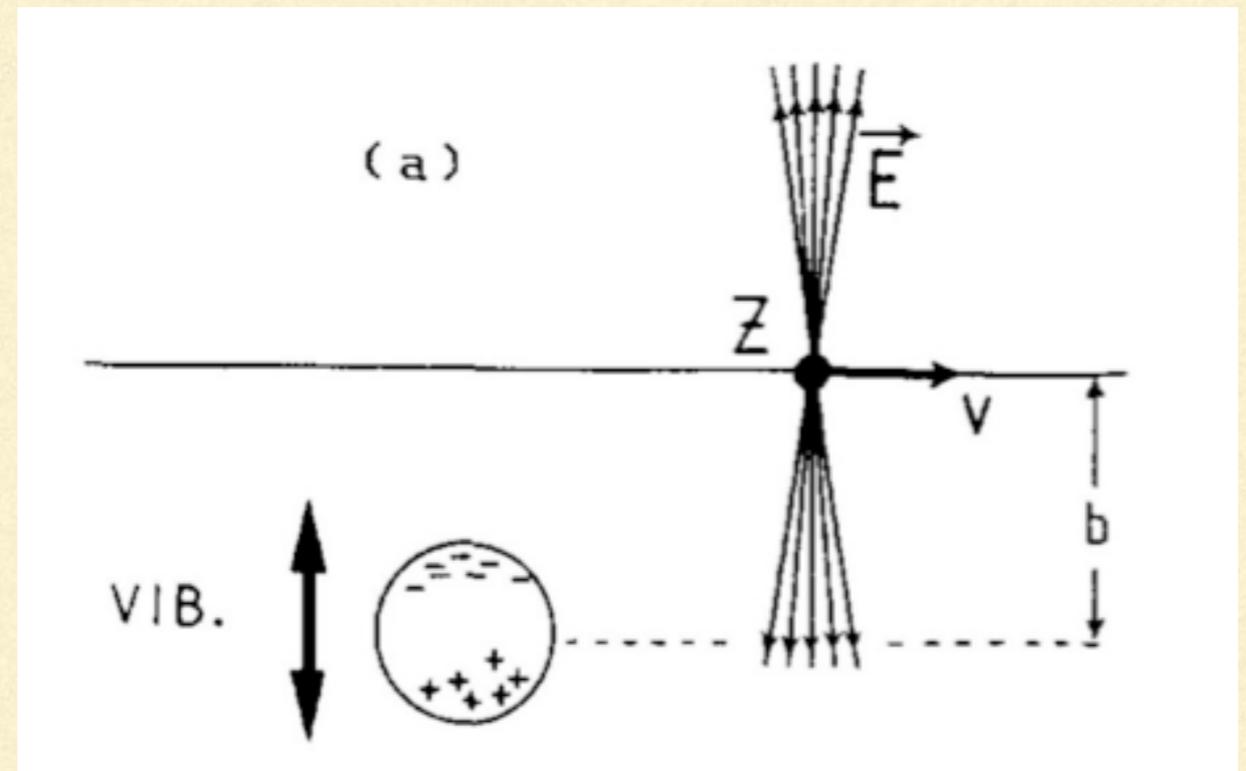
- Do curves all fluctuate around zero with some common variance?
- What is the correlation length? Is it different at each order?



# Coulomb dissociation of halo nuclei

Bertulani, arXiv:0908.4307

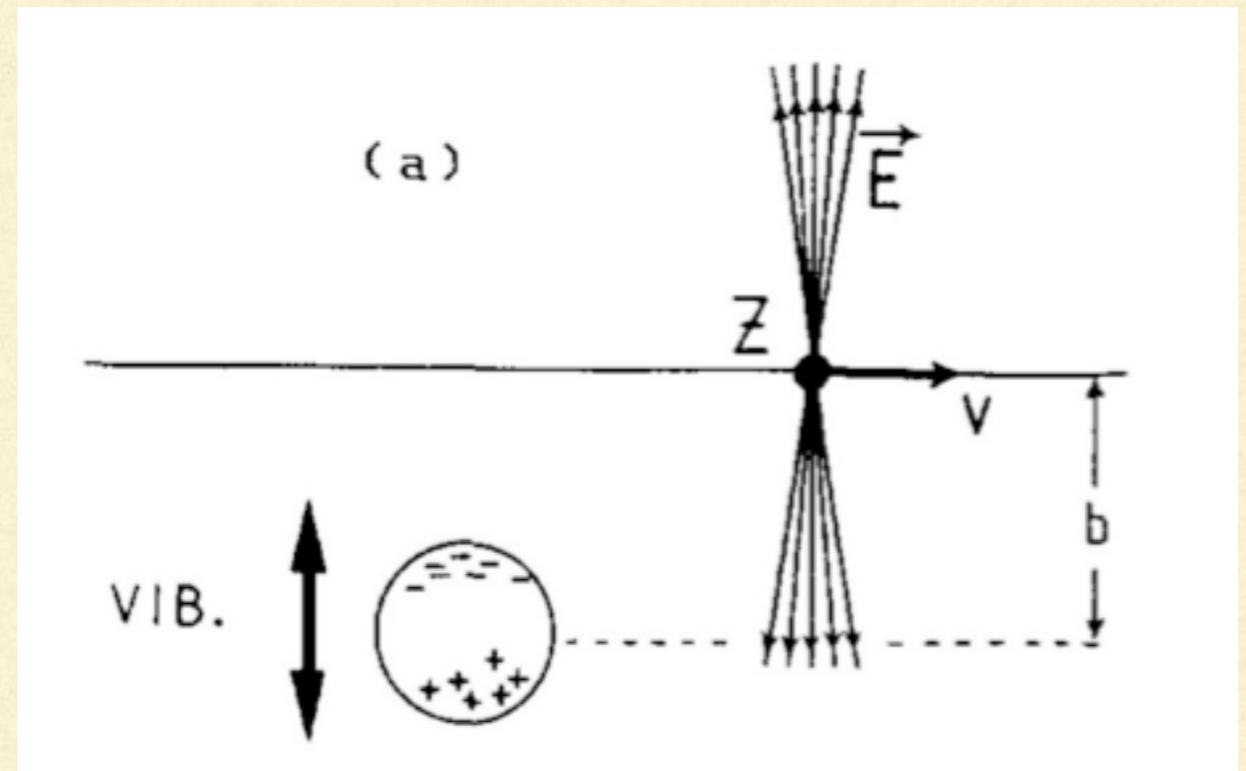
- Coulomb dissociation: collide halo nucleus (we hope peripherally) with a high- $Z$  nucleus
- Do with different  $Z$ , different nuclear sizes, different energies to test systematics



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- Coulomb excitation dissociation cross section (p.v.  $b \gg R_{\text{target}}$ )

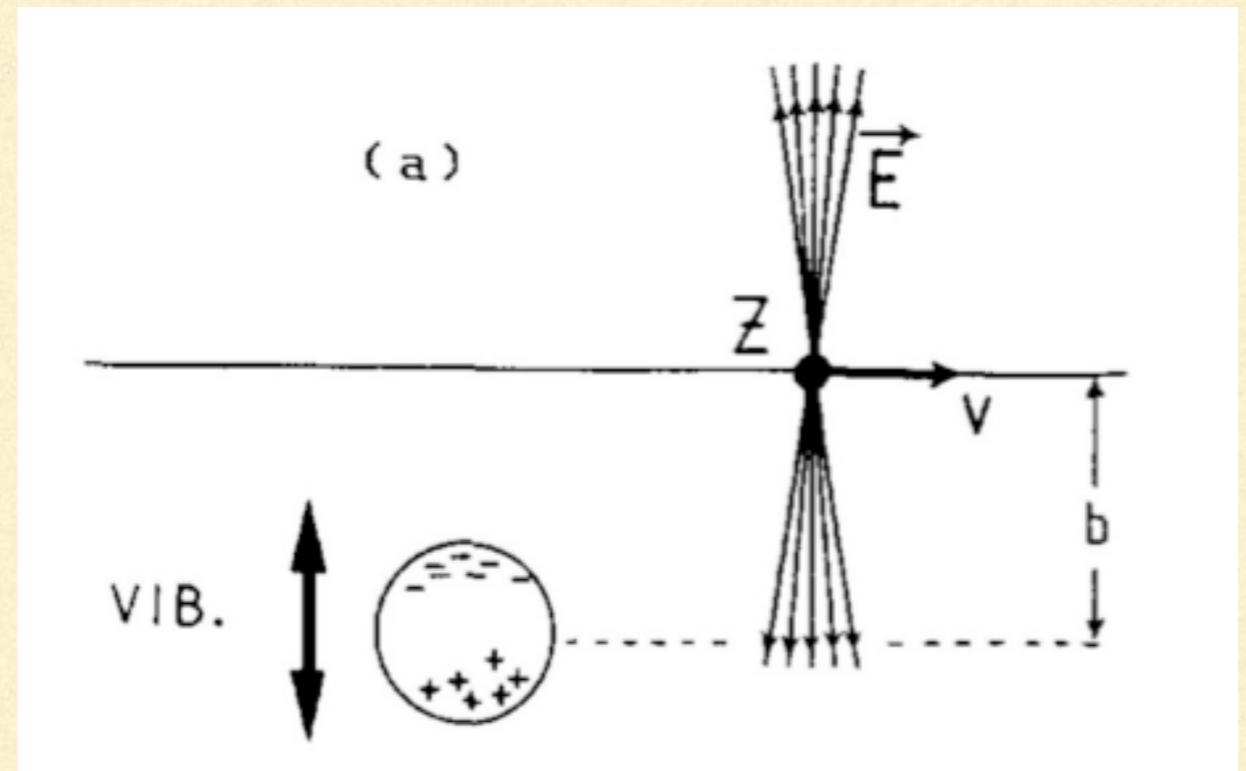
$$\frac{d\sigma_C}{2\pi b db} = \sum_{\pi L} \int \frac{dE_\gamma}{E_\gamma} n_{\pi L}(E_\gamma, b) \sigma_\gamma^{\pi L}(E_\gamma)$$

- $n_{\pi L}(E_\gamma, b)$  virtual photon numbers, dependent only on kinematic factors. Number of equivalent (virtual) photons that strike the halo nucleus.

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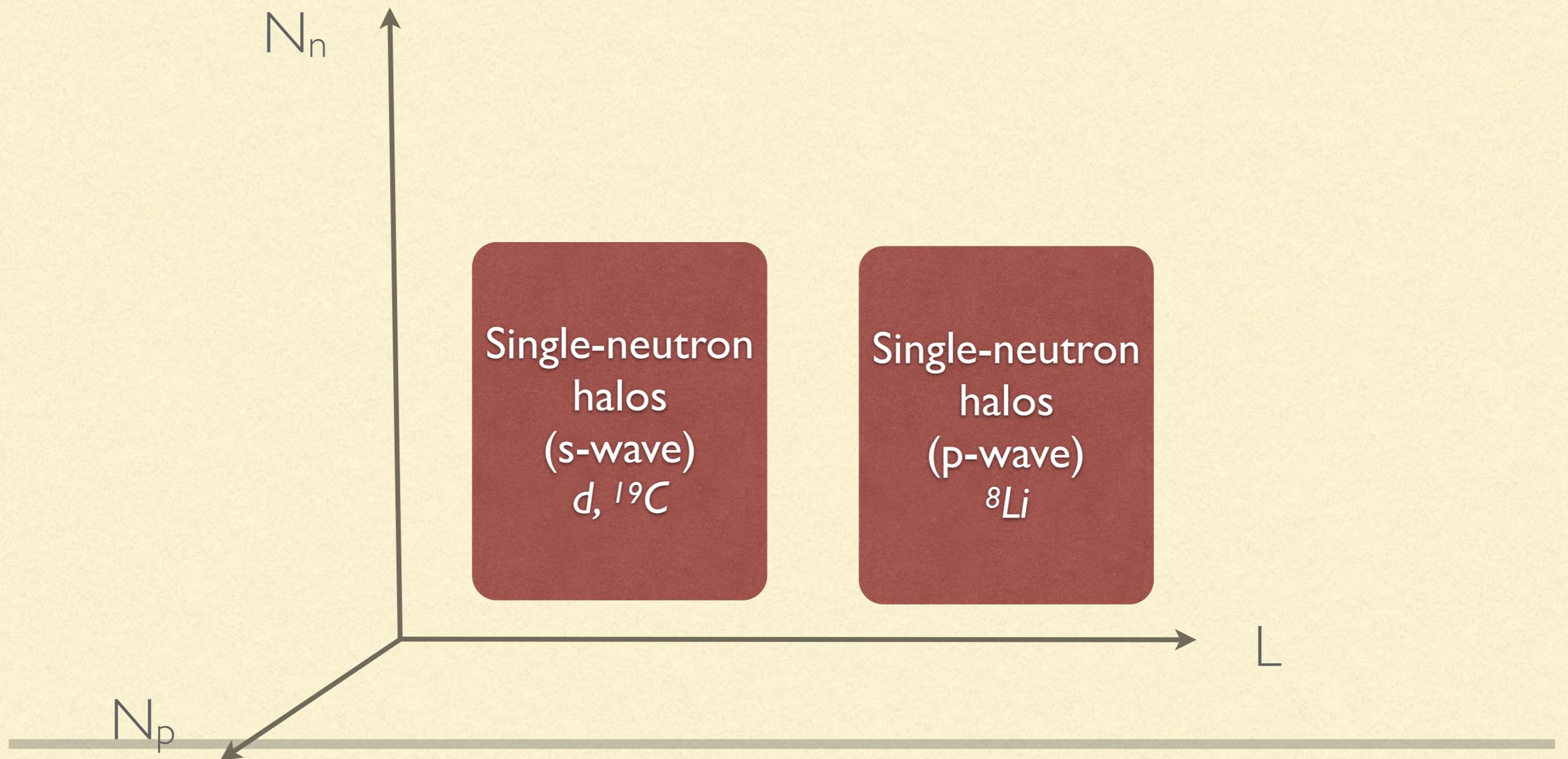


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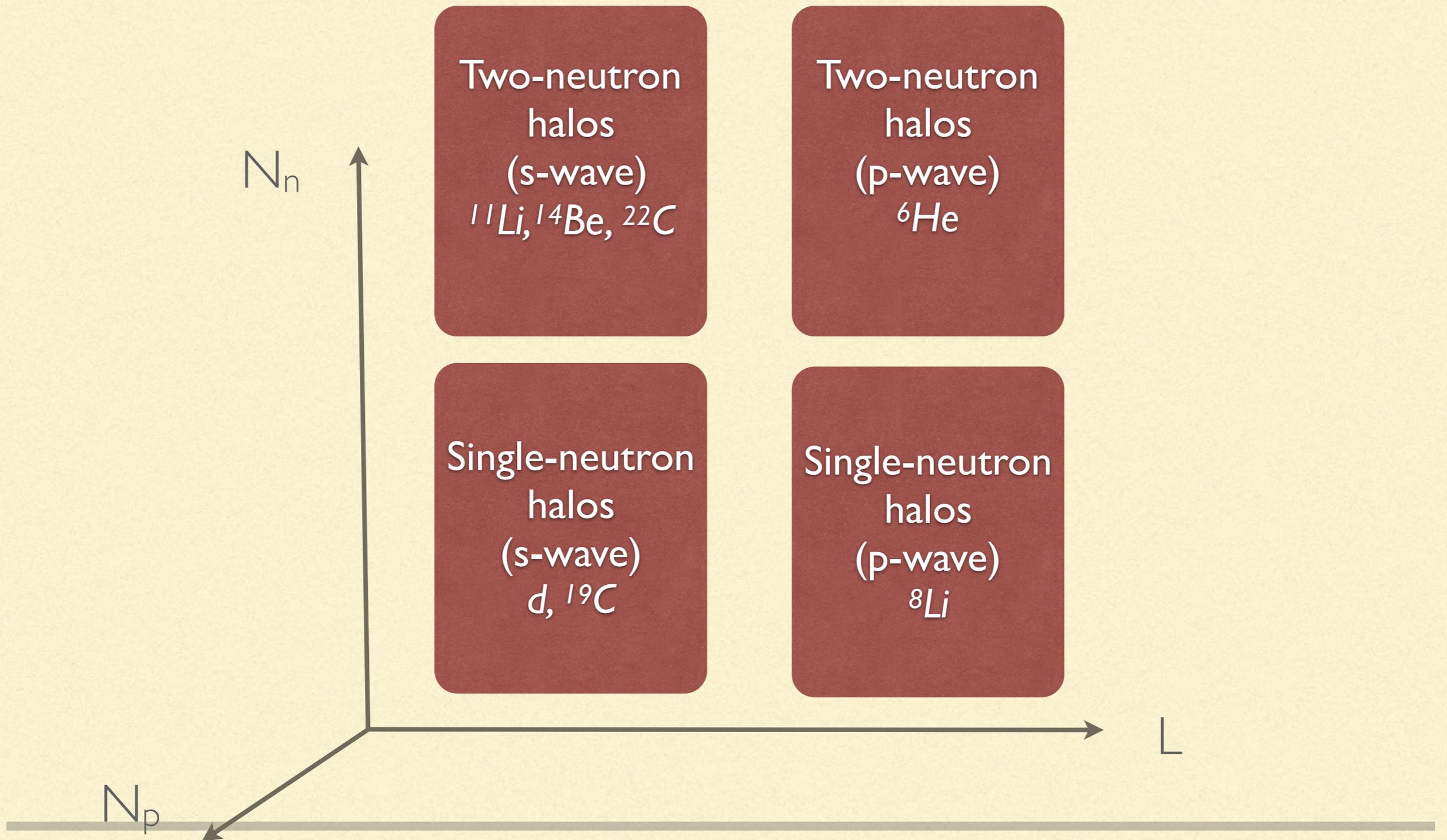
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- $n_{\pi L}(E_\gamma, b)$  virtual photon numbers, dependent only on kinematic factors. Number of equivalent (virtual) photons that strike the halo nucleus.
- $\sigma_\gamma^{\pi L}(E_\gamma)$  can then be extracted: it's the (total) cross section for dissociation of the nucleus due to the impact of photons of multipolarity  $\pi L$ .

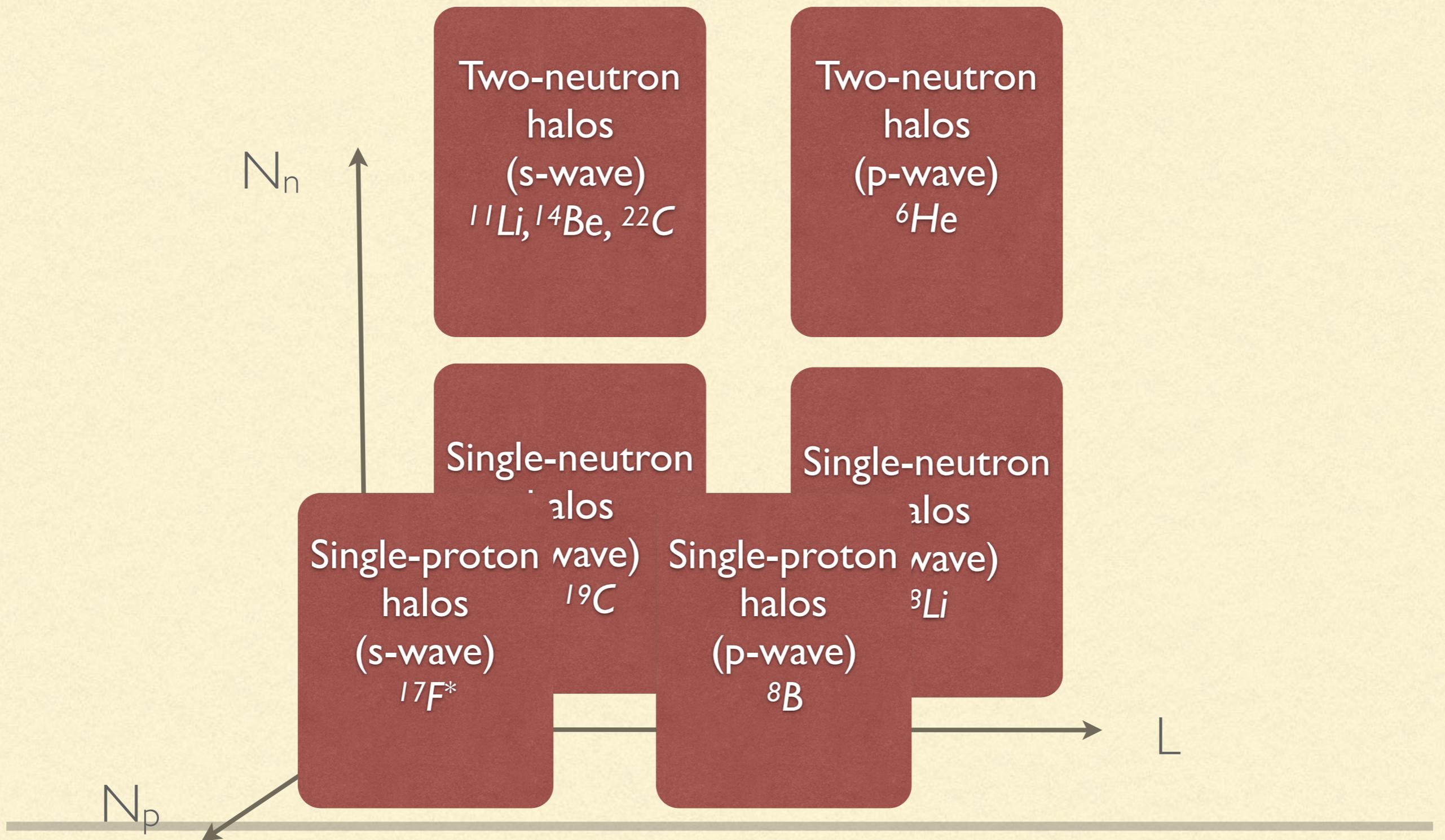
# The multi-dimensional Halo EFT space



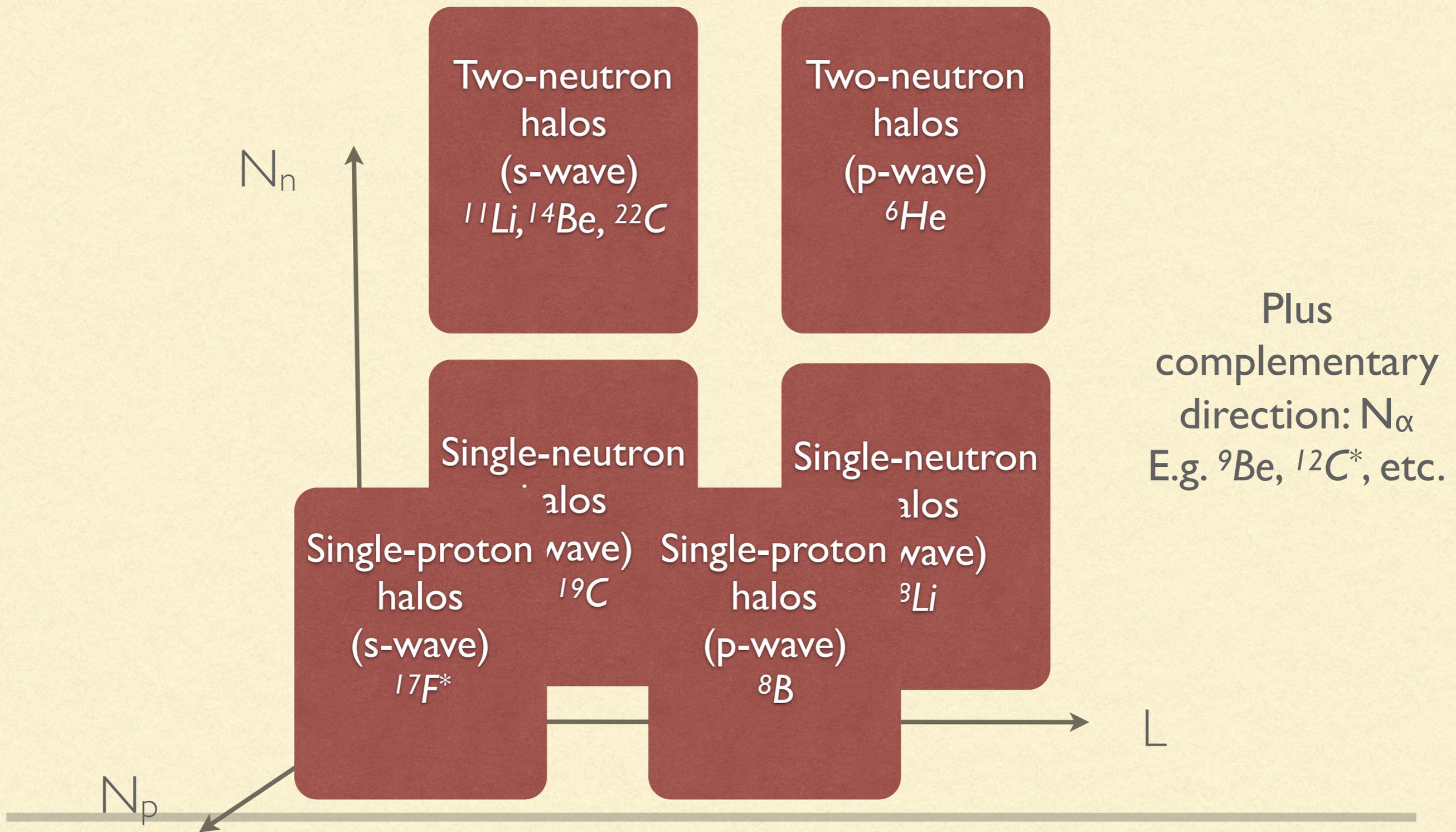
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# Lagrangian for s- and p-wave states

s-wave: Kaplan, Savage, Wise (1998); van Kolck (1999); Birse, Richardosn, McGovern 1999)

p-wave: Bertulani, Hammer, van Kolck (2002); Bedaque, Hammer, van Kolck (2003)

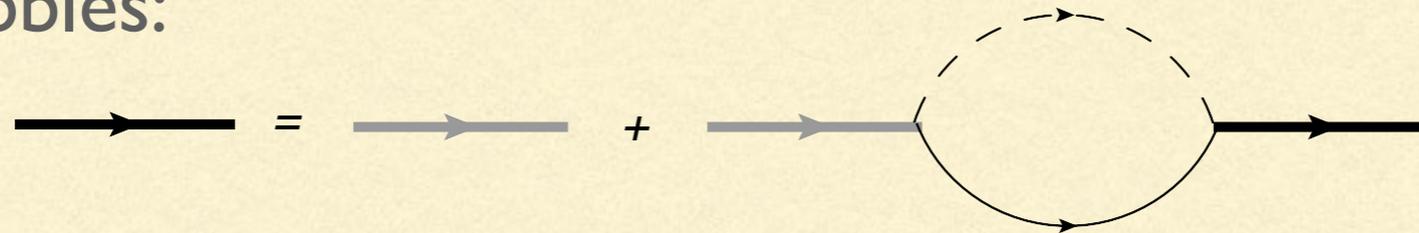
$$\begin{aligned}
 \mathcal{L} = & c^\dagger \left( i\partial_t + \frac{\nabla^2}{2M} \right) c + n^\dagger \left( i\partial_t + \frac{\nabla^2}{2m} \right) n \\
 & + \sigma^\dagger \left[ \eta_0 \left( i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_0 \right] \sigma + \pi_j^\dagger \left[ \eta_1 \left( i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_1 \right] \pi_j \\
 & - g_0 [\sigma n^\dagger c^\dagger + \sigma^\dagger n c] - \frac{g_1}{2} \left[ \pi_j^\dagger (n i\overleftrightarrow{\nabla}_j c) + (c^\dagger i\overleftrightarrow{\nabla}_j n^\dagger) \pi_j \right] \\
 & - \frac{g_1}{2} \frac{M - m}{M_{nc}} \left[ \pi_j^\dagger i\overrightarrow{\nabla}_j (nc) - i\overleftarrow{\nabla}_j (n^\dagger c^\dagger) \pi_j \right] + \dots,
 \end{aligned}$$

- $c, n$ : “core”, “neutron” fields.  $c$ : boson,  $n$ : fermion
- $\sigma, \pi_j$ : S-wave and P-wave fields
- Minimal substitution generates leading EM couplings

# Dressing the s-wave state

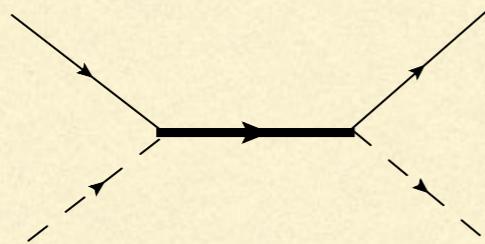
Kaplan, Savage, Wise; van Kolck; Gegelia;  
Birse, Richardson, McGovern

- $\sigma_{nc}$  coupling  $g_0$  of order  $R_{\text{halo}}$ ,  $nc$  loop of order  $1/R_{\text{halo}}$ . Therefore need to sum all bubbles:



$$D_\sigma(p) = \frac{1}{\Delta_0 + \eta_0 [p_0 - \mathbf{p}^2 / (2M_{nc})] - \Sigma_\sigma(p)}$$

$$\Sigma_\sigma(p) = -\frac{g_0^2 m_R}{2\pi} \left[ \mu + i \sqrt{2m_R \left( p_0 - \frac{\mathbf{p}^2}{2M_{nc}} + i\eta \right)} \right] \quad (\text{PDS})$$



$$t = \frac{2\pi}{m_R} \frac{1}{\frac{1}{a_0} - \frac{1}{2} r_0 k^2 + ik}$$

$$D_\sigma(p) = \frac{2\pi\gamma_0}{m_R^2 g_0^2} \frac{1}{1 - r_0\gamma_0} \frac{1}{p_0 - \frac{\mathbf{p}^2}{2M_{nc}} + B_0} + \text{regular}$$

Counting in S waves:  
 $a_0 \sim R_{\text{halo}} \sim 1/\gamma_0$ ;  $r_0 \sim R_{\text{core}}$ .  
 $r_0 = 0$  at LO.

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# One-slide p-wave review

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$$\langle \mathbf{k} | t_1 | \mathbf{k}' \rangle = -\frac{6\pi}{m_R} \frac{\mathbf{k} \cdot \mathbf{k}'}{-\frac{1}{a_1} + \frac{1}{2}r_1 k^2 - ik^3}$$

Bethe (1949)

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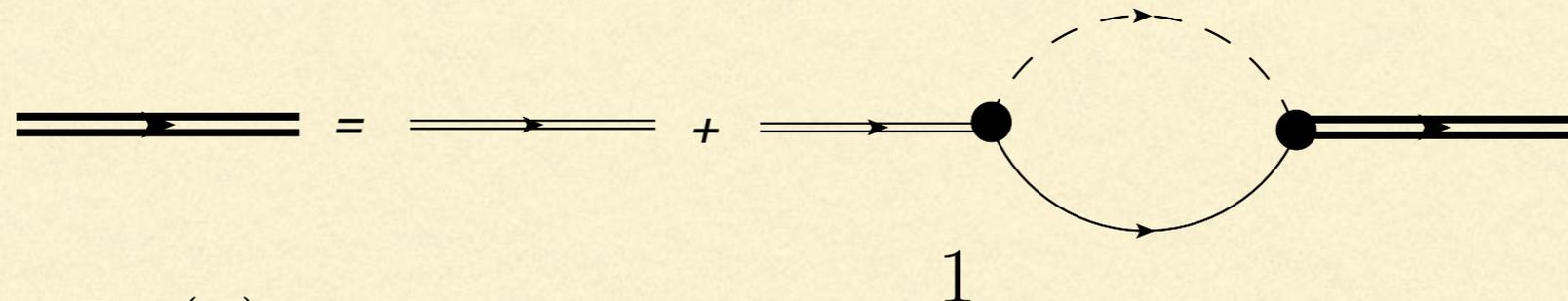
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- But what if there is a low-energy p-wave resonance?
- Causality says  $r_1 \lesssim -1/R$  Wigner (1955); Hammer & Lee (2009); Nishida (2012)
- So low-energy resonance/bound state would seem to have to arise due to cancellation between  $-1/a_1$  and  $1/2 r_1 k^2$  terms.
- $a_1 \sim R/M_{\text{lo}}^2$  gives  $k_R \sim M_{\text{lo}}$  Bedaque, Hammer, van Kolck (2003)

# Dressing the p-wave state

Bertulani, Hammer, van Kolck (2002); Bedaque, Hammer, van Kolck (2003)

- Proceed similarly for p-wave state as for s-wave state


$$D_{\pi}(p) = \frac{1}{\Delta_1 + \eta_1 [p_0 - \mathbf{p}^2 / (2M_{nc})] - \Sigma_{\pi}(p)}$$

- Here both  $\Delta_1$  and  $g_1$  are mandatory for renormalization at LO

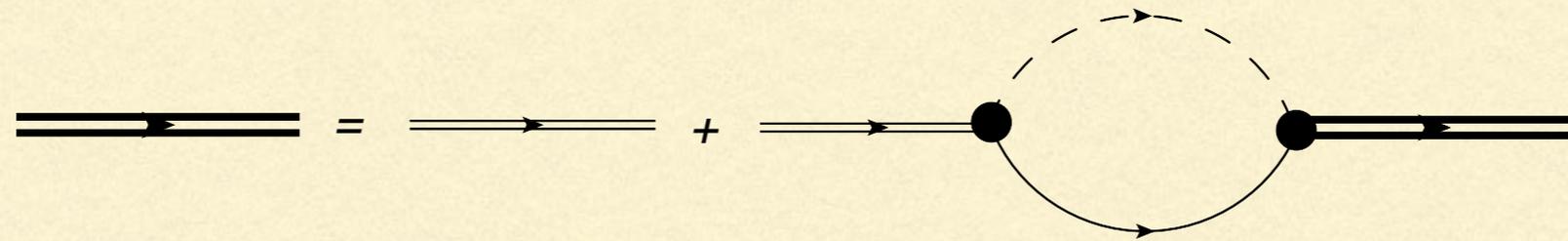
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- If  $a_1 > 0$  then pole is at  $k=i\gamma_1$  with  $B_1=\gamma_1^2/(2m_R)$ :

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# A narrow p-wave resonance/bound state

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- Resonance width is  $\sim E_R k_R/r_1$ , so it is parametrically narrow. Need to resum width if  $k^2 - k_R^2$  gets small
-

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# P-wave FSI in $\gamma_{E1} + {}^{11}\text{Be} \rightarrow {}^{10}\text{Be} + n$

Typel & Baur, Phys. Rev. Lett. 93, 142502 (2004); Nucl. Phys. A759, 247 (2005); Eur. Phys. J. A 38, 355 (2008)

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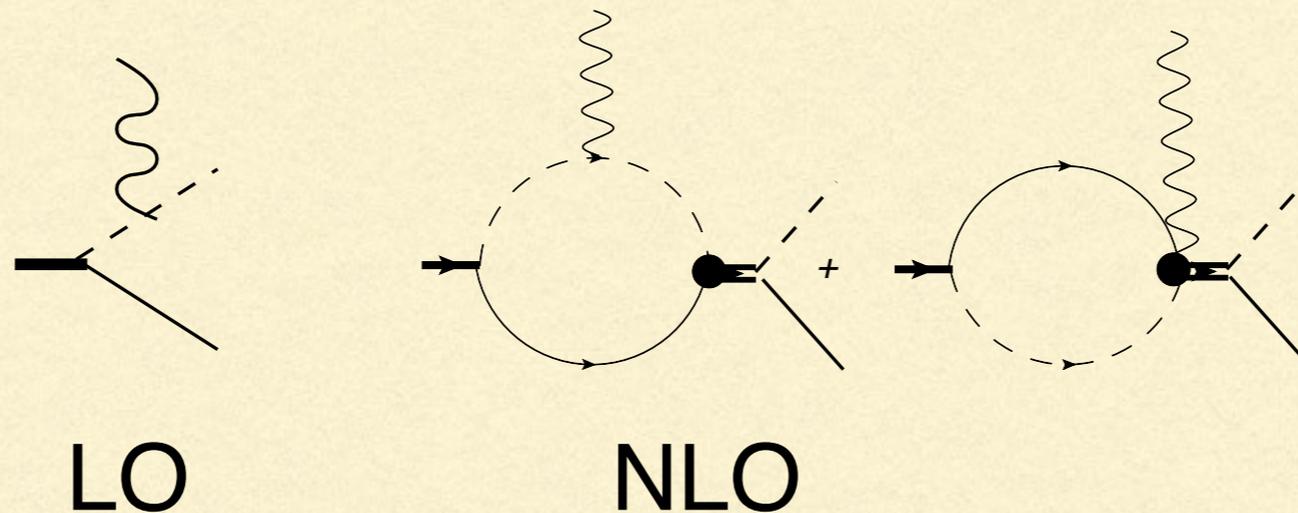
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$$k^3 \cot \delta_1 = -1/2 r_1 (k^2 + \gamma_1^2) \Rightarrow \delta_1 \sim R_{\text{core}}/R_{\text{halo}} \text{ if } k \sim 1/R_{\text{halo}} \sim \gamma_1.$$

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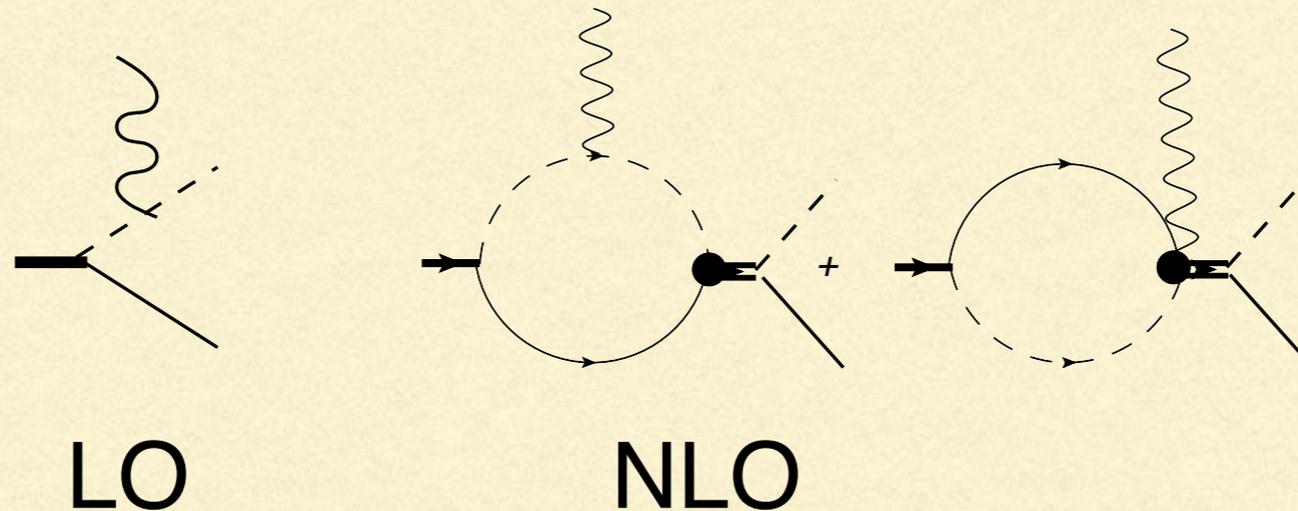
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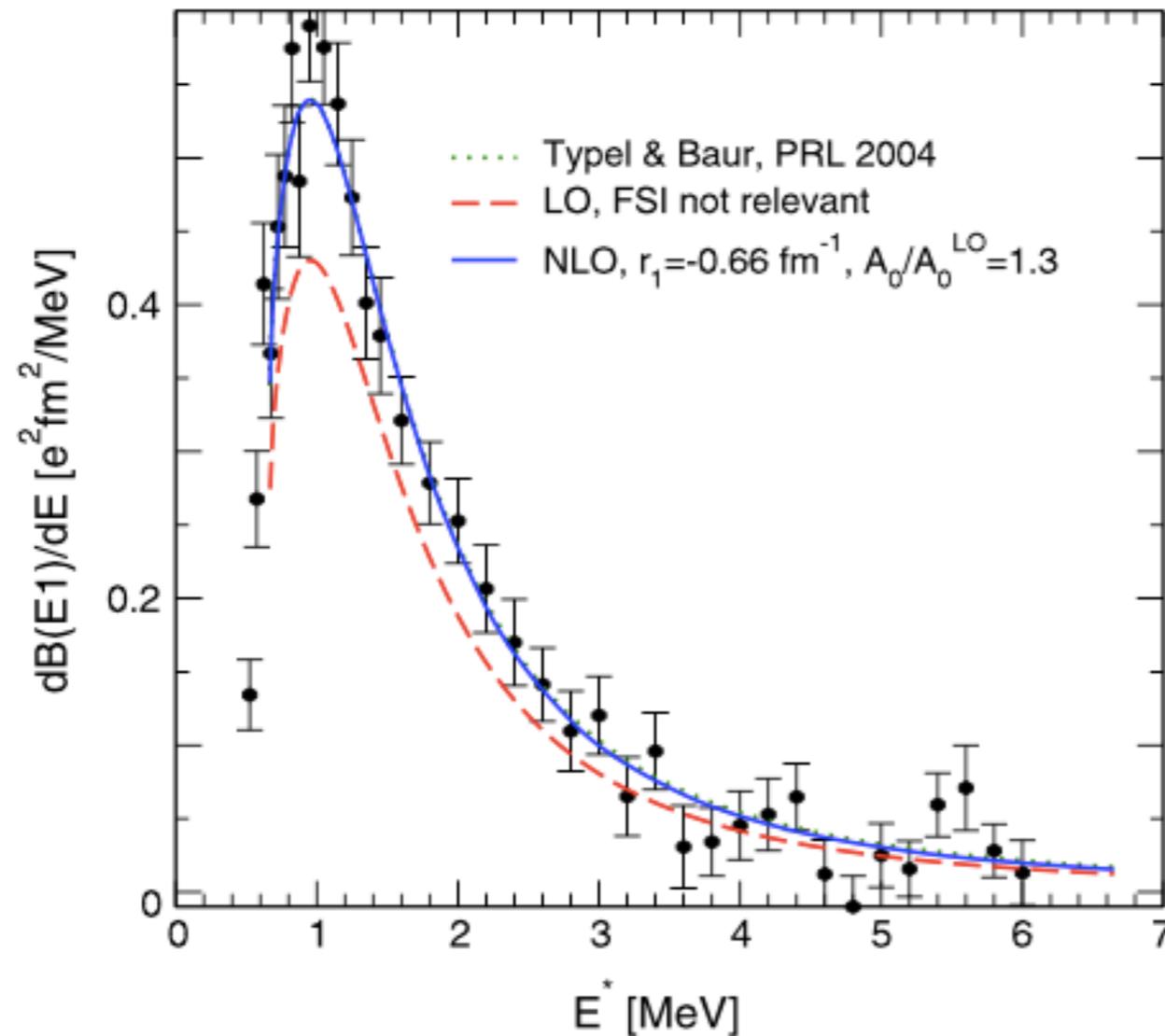


- Need both  $\gamma_1$  and  $r_1 \equiv A_1$  at NLO in this observable.  $A_0$  also becomes a free parameter at NLO: fit it to Coulomb dissociation data

# Coulomb dissociation of $^{11}\text{Be}$ : result

Data: Palit et al., 2003

Analysis: Hammer, Phillips. NPA, 2011

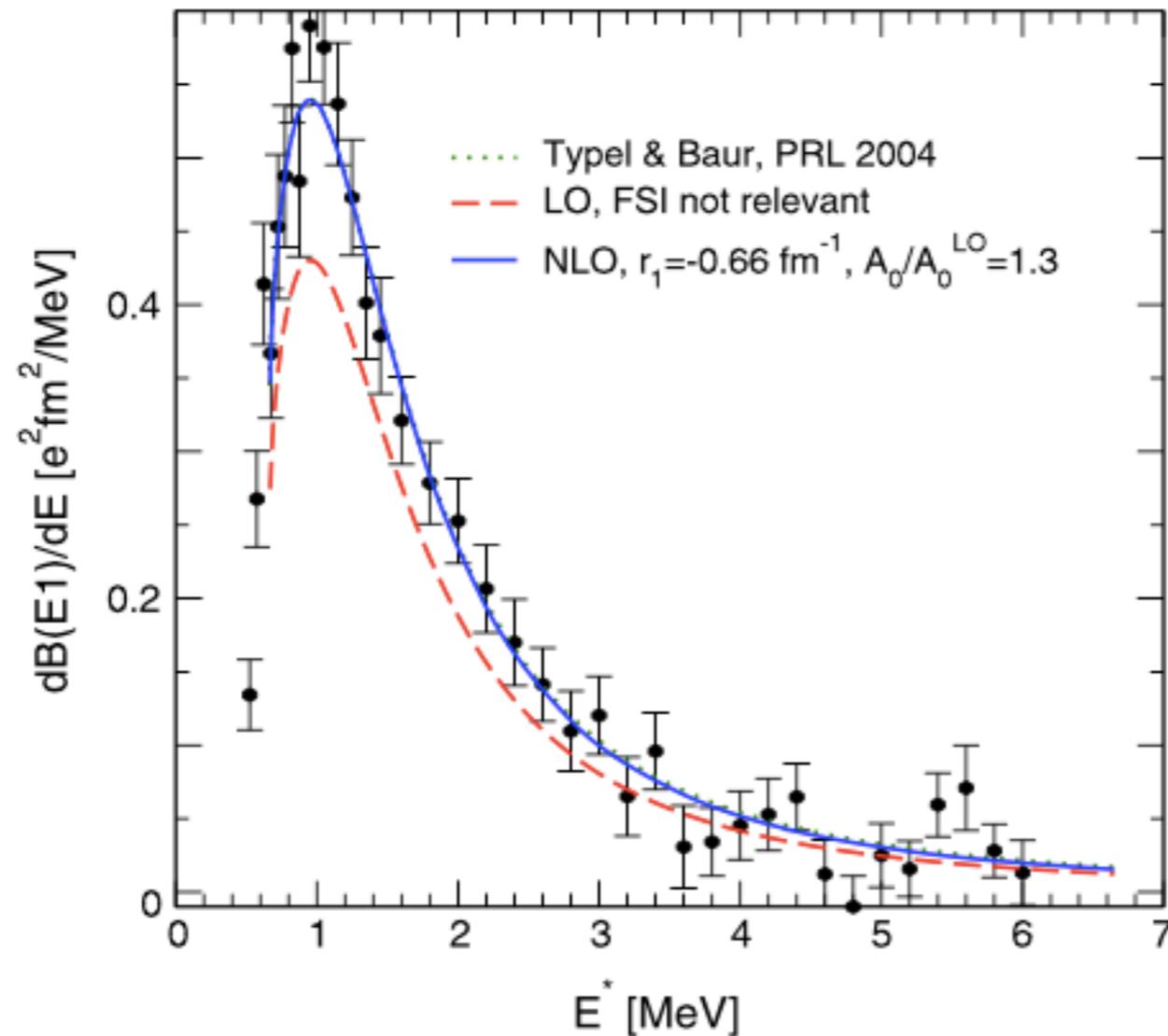


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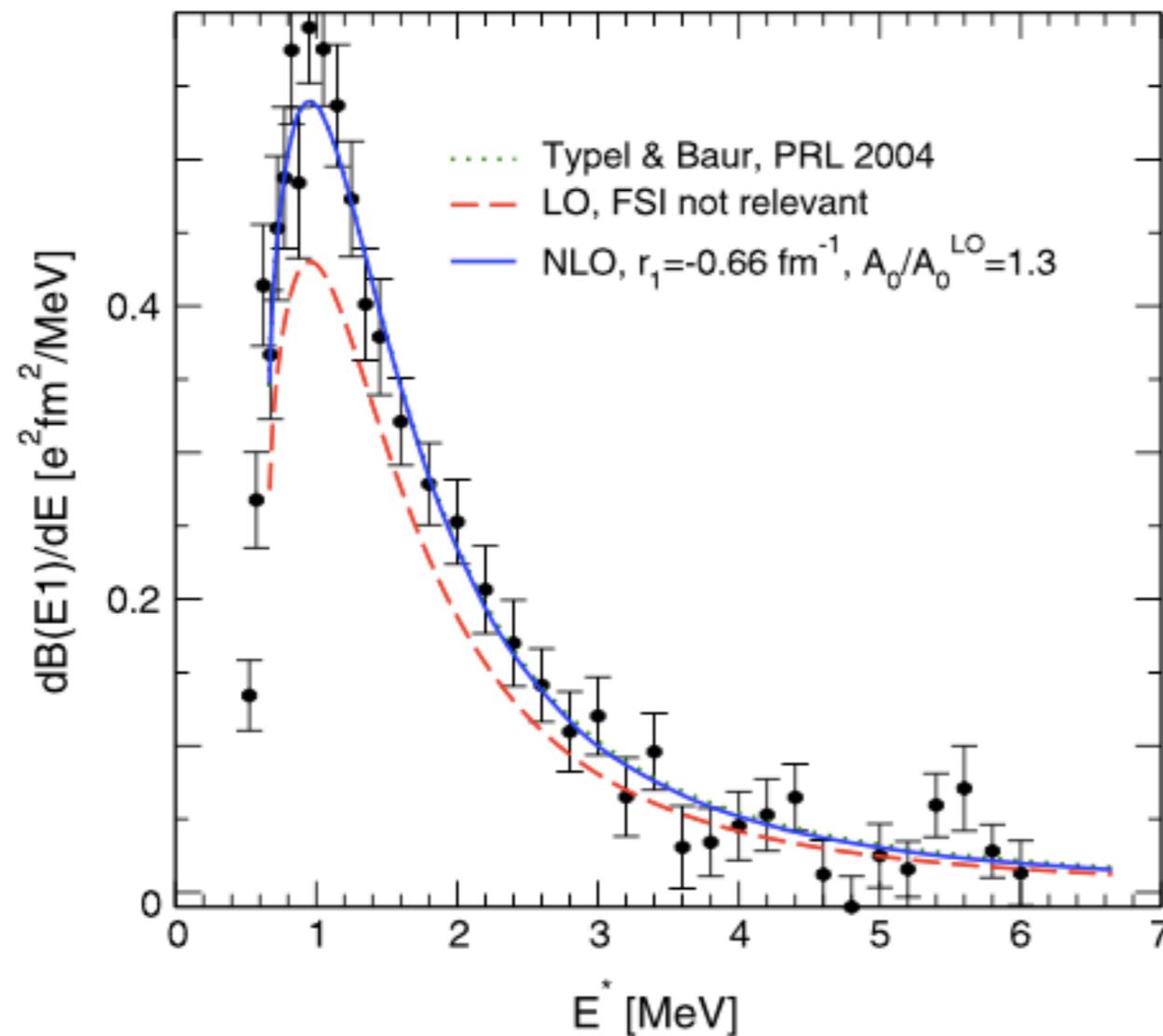
- Here value of  $r_1$  used to fit  $B(E1: 1/2^+ \rightarrow 1/2^-)$  works.

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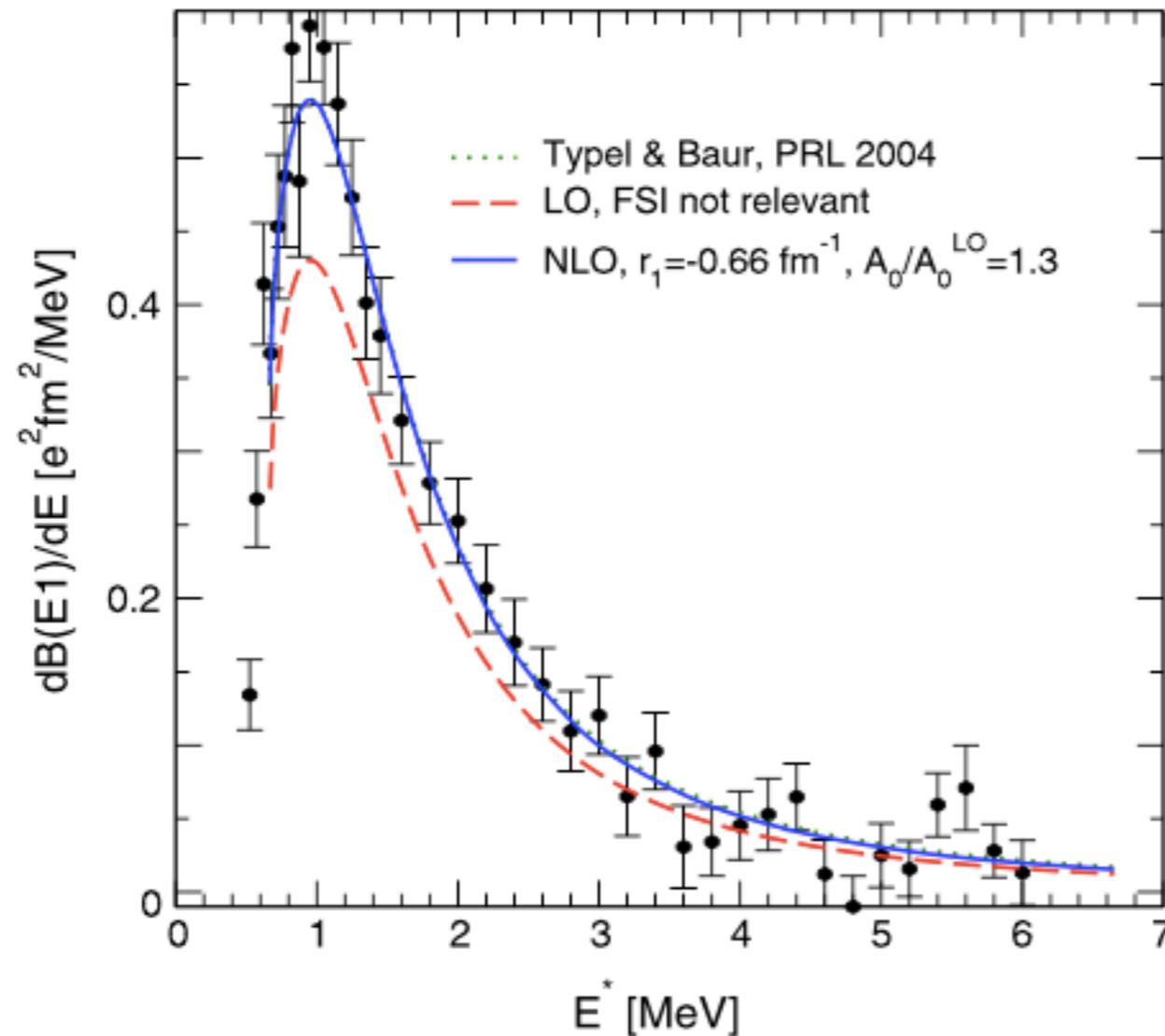
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$$\text{NLO: } (\langle r_c^2 \rangle + \langle r_{\text{Be}}^2 \rangle)^{1/2} = 2.44 \text{ fm}$$

# Coulomb dissociation of ${}^7\text{Be}$ : result

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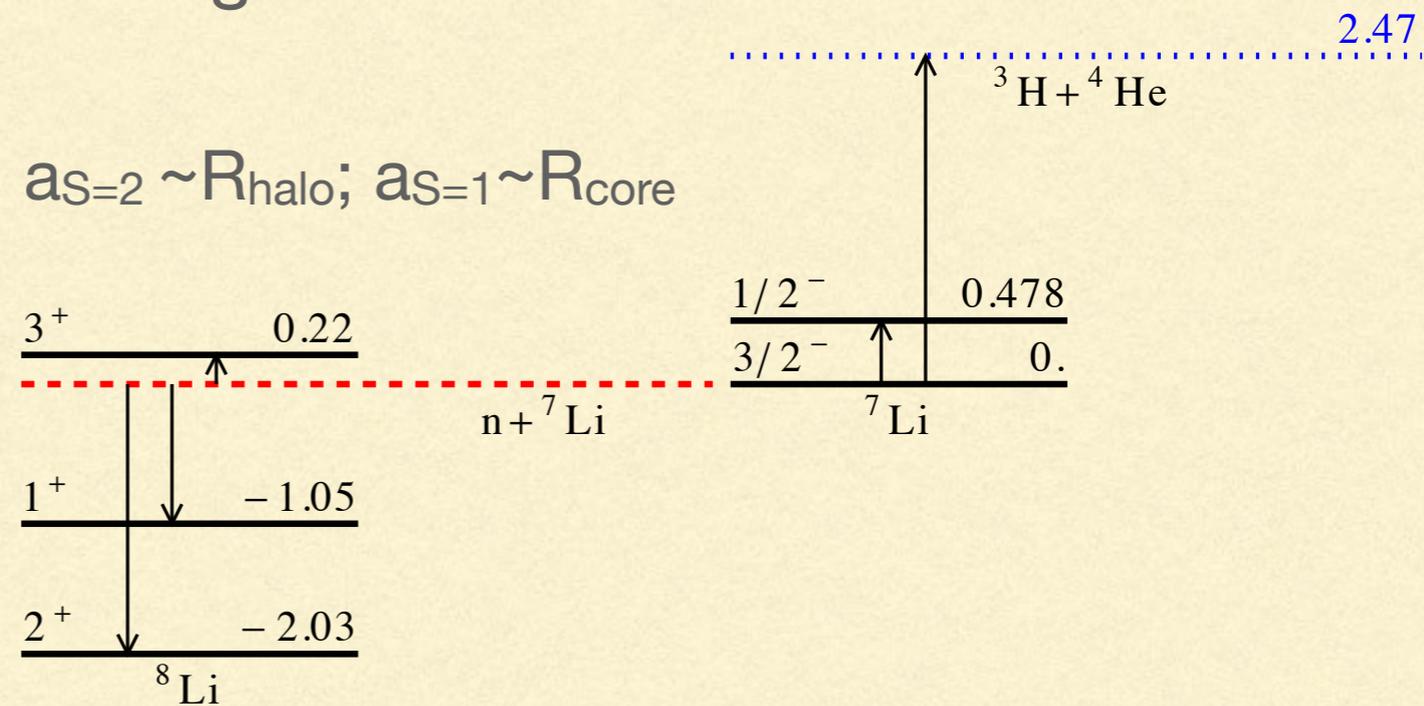
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Use of ab initio input, e.g. for ANC?



- ${}^7\text{Li}$  ground state is  $3/2^-$ : S-wave n scattering in  ${}^5S_2$  and  ${}^3S_1$

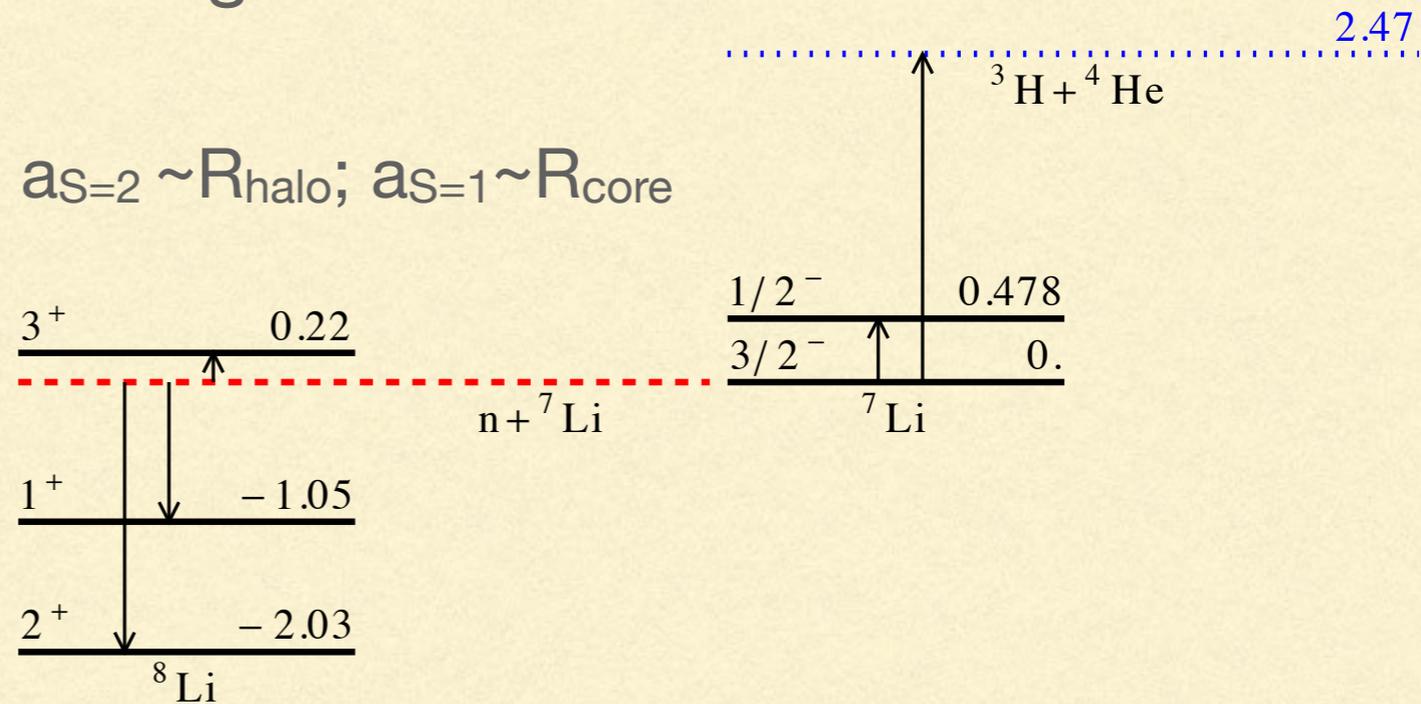




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- $a_{S=2} = -3.63(5)$  fm,  $a_{S=1} = 0.87(7)$  fm

$a_{S=2} \sim R_{\text{halo}}$ ;  $a_{S=1} \sim R_{\text{core}}$

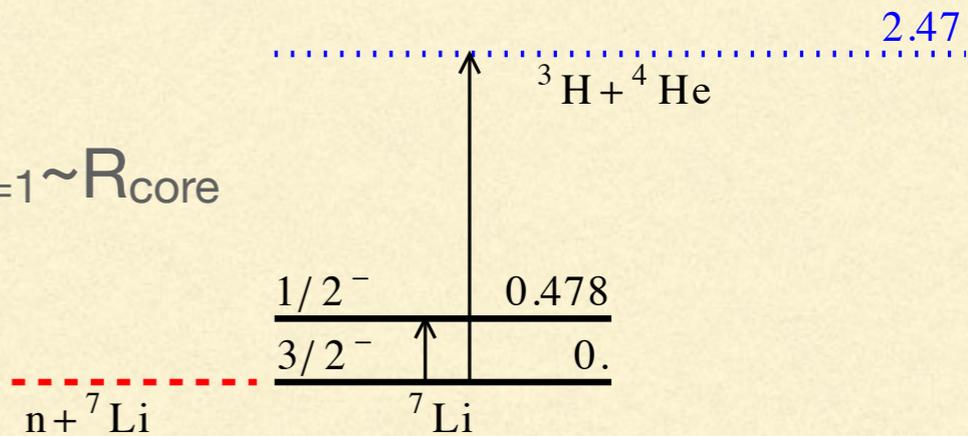
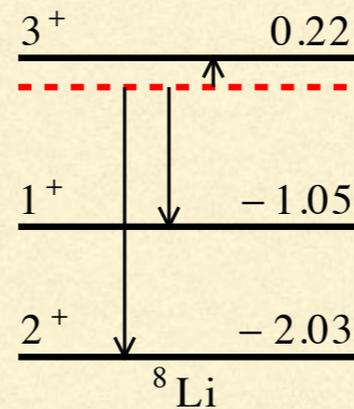
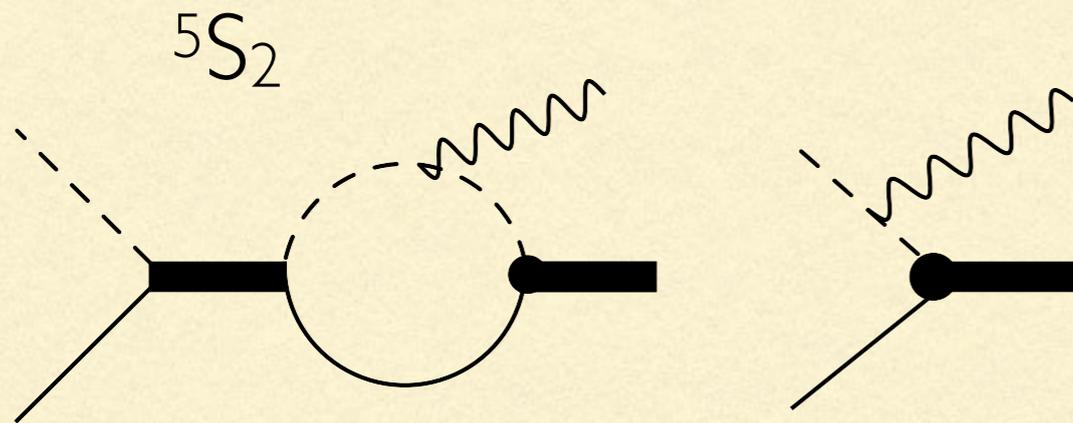




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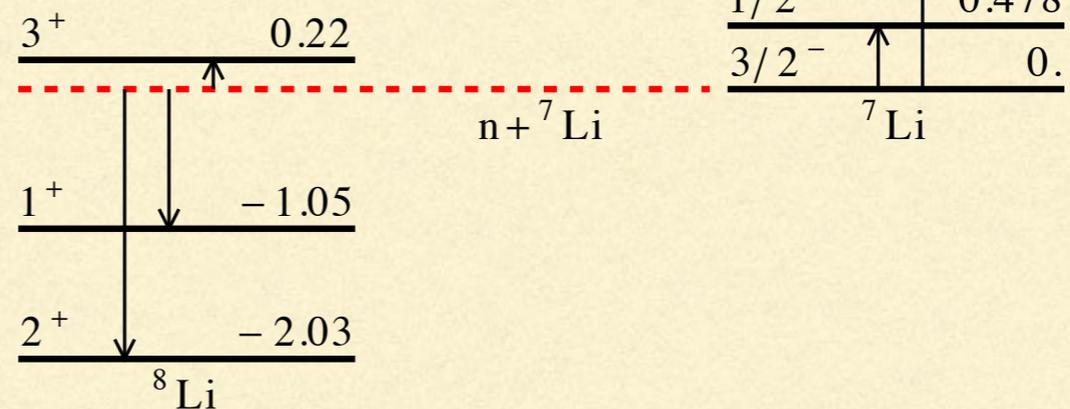
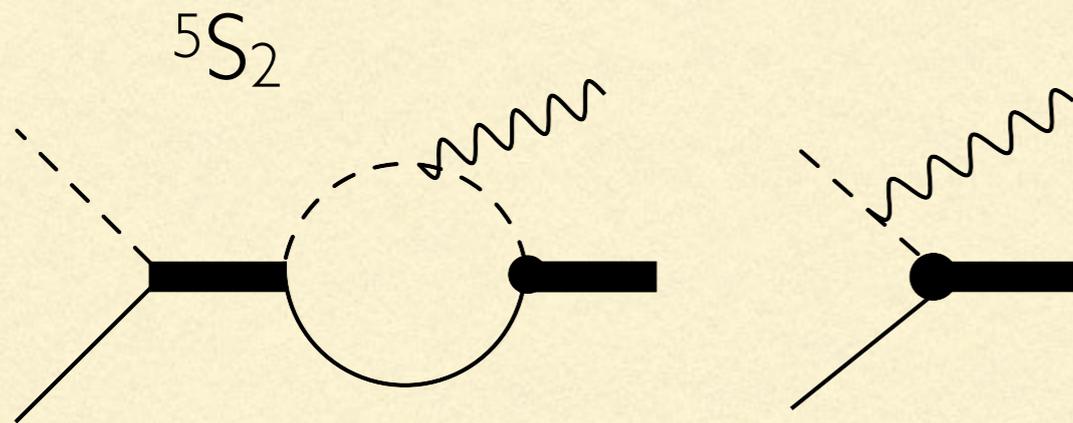




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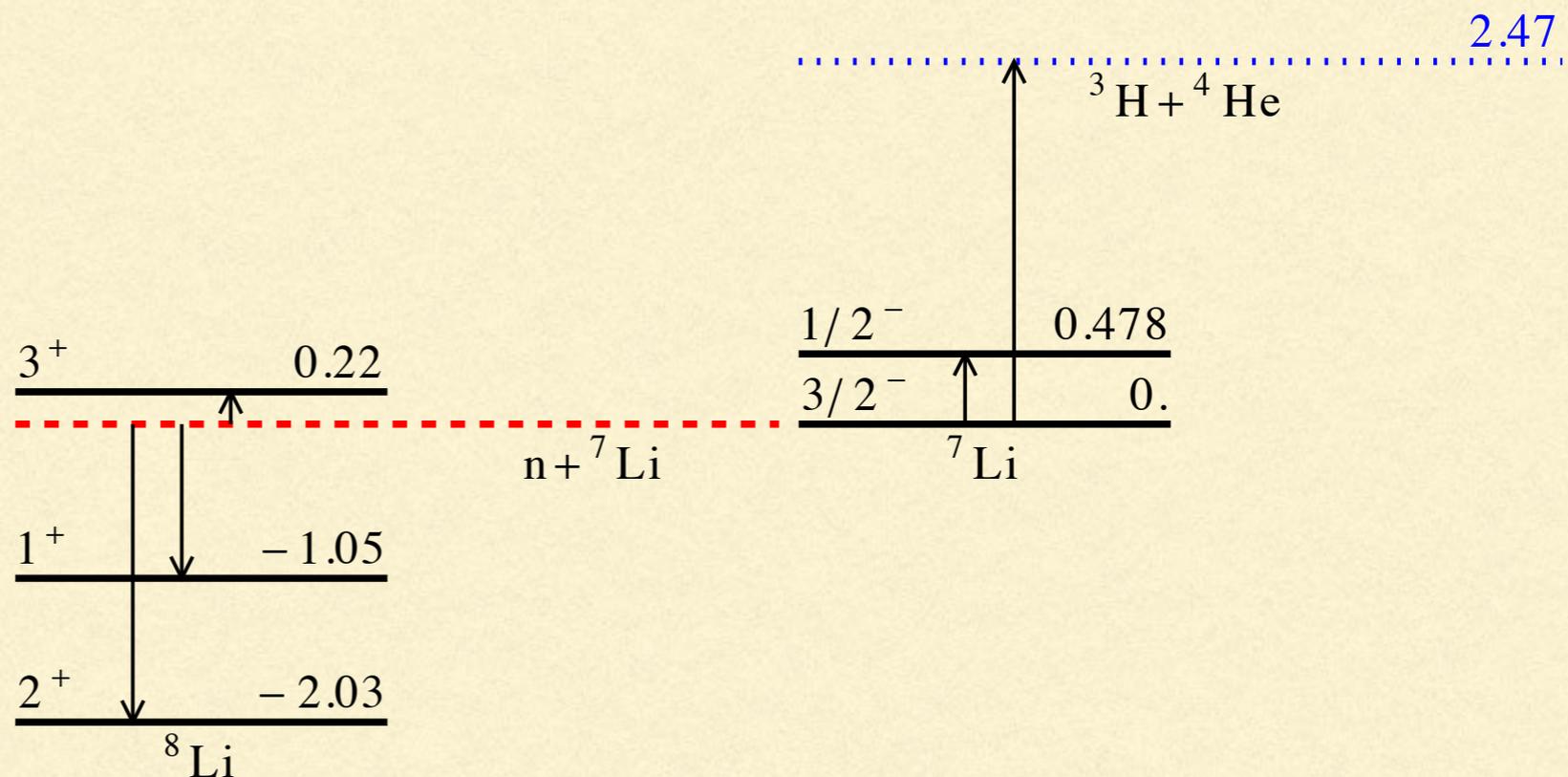
- LO calculation:  $S=2$  (with ISI) and  $S=1$  into P-wave bound state

$$E1 \propto \int_0^\infty dr u_0(r) r u_1(r);$$

$$u_0(r) = 1 - \frac{r}{a}; u_1(r) = A_1 \exp(-\gamma_1 r) \left( 1 + \frac{1}{\gamma_1 r} \right)$$

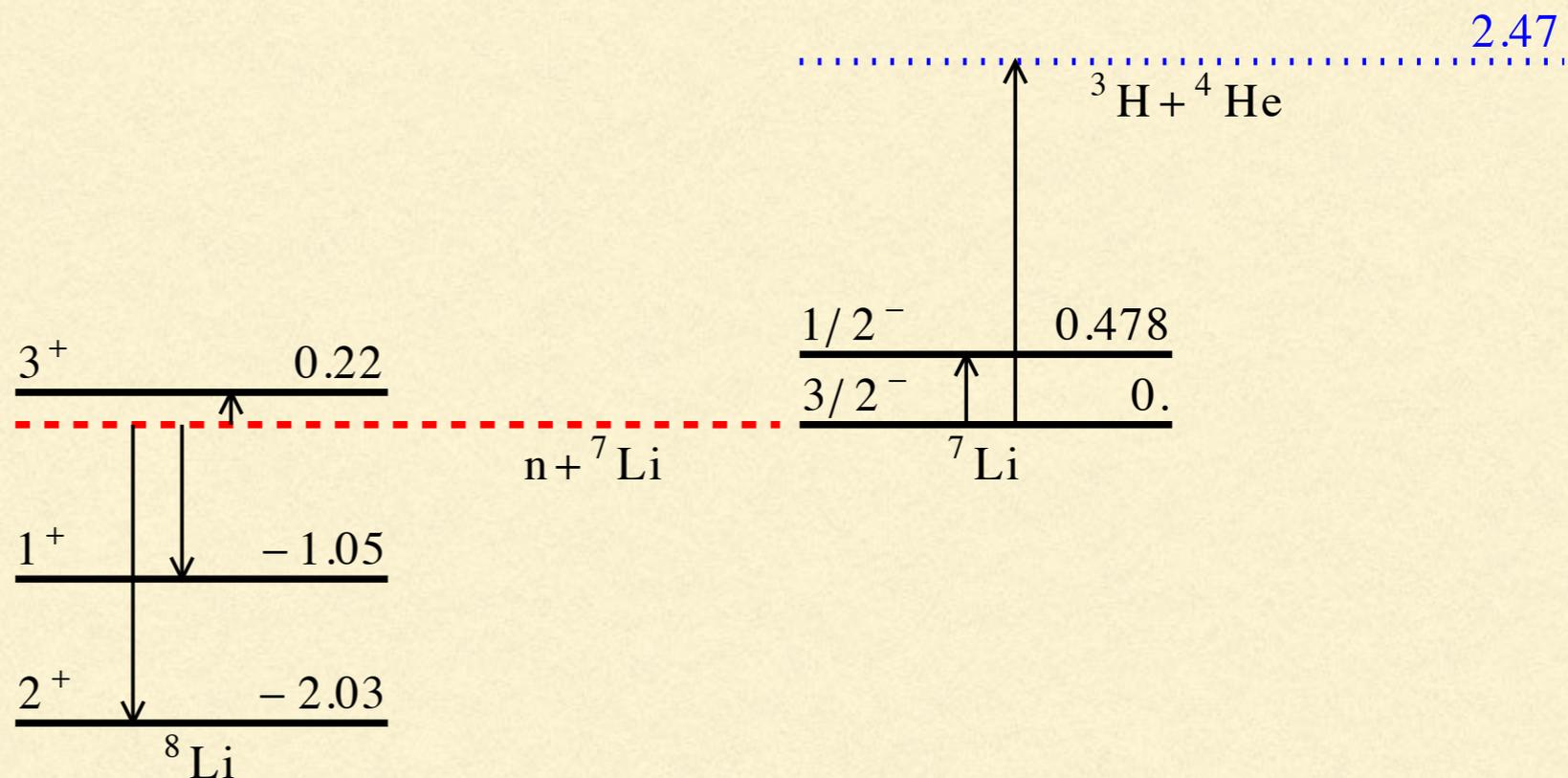
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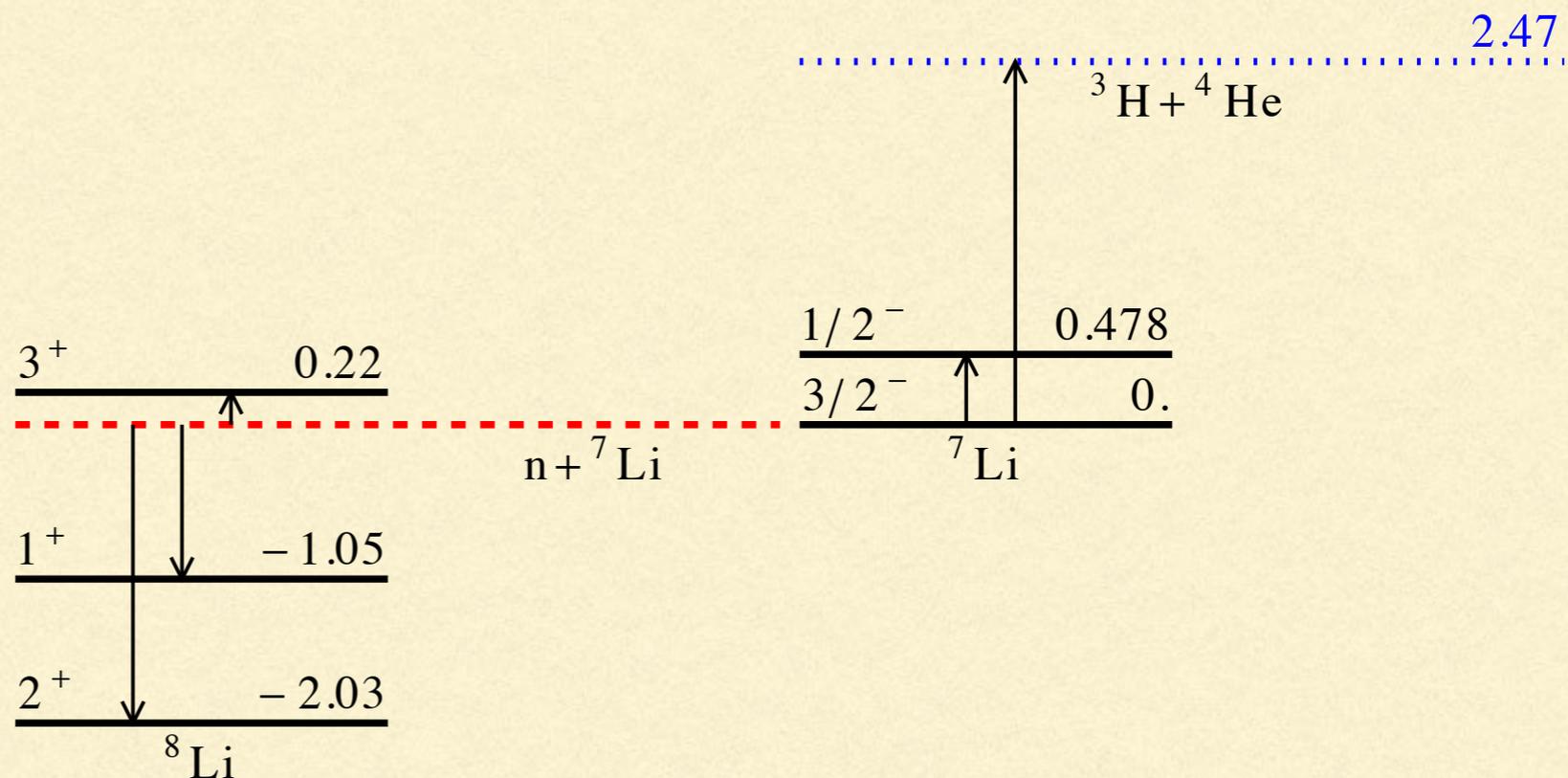
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-

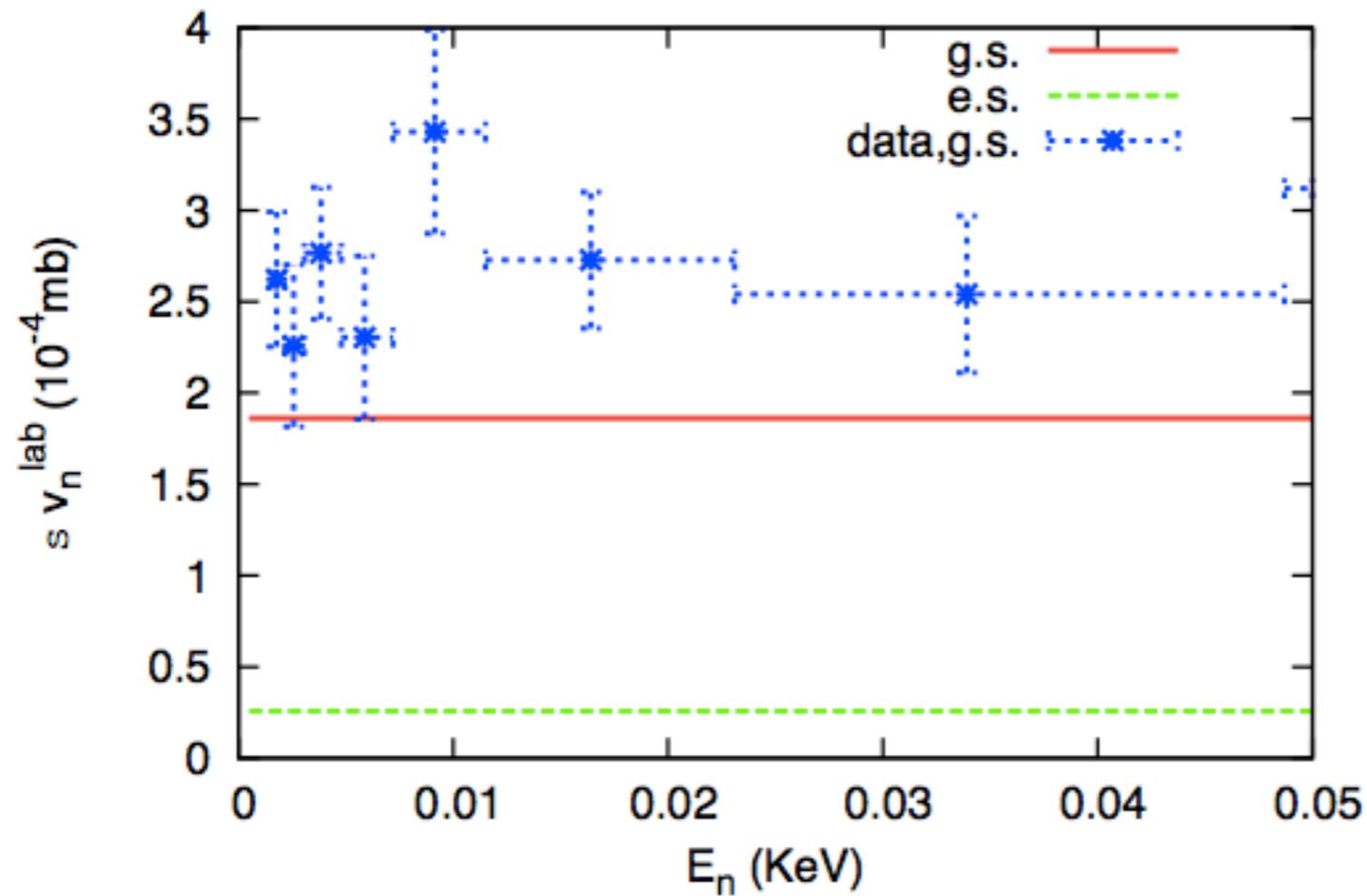
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	$A_{(3P2)}$	$A_{(5P2)}$	$A_{(3P2^*)}$	$A_{(3P1)^*}$	$A_{(5P1)^*}$
Nollett	-0.283(12)	-0.591(12)	-0.384(6)	0.220(6)	0.197(5)
Trache	-0.284(23)	-0.593(23)		0.187(16)	0.217(13)

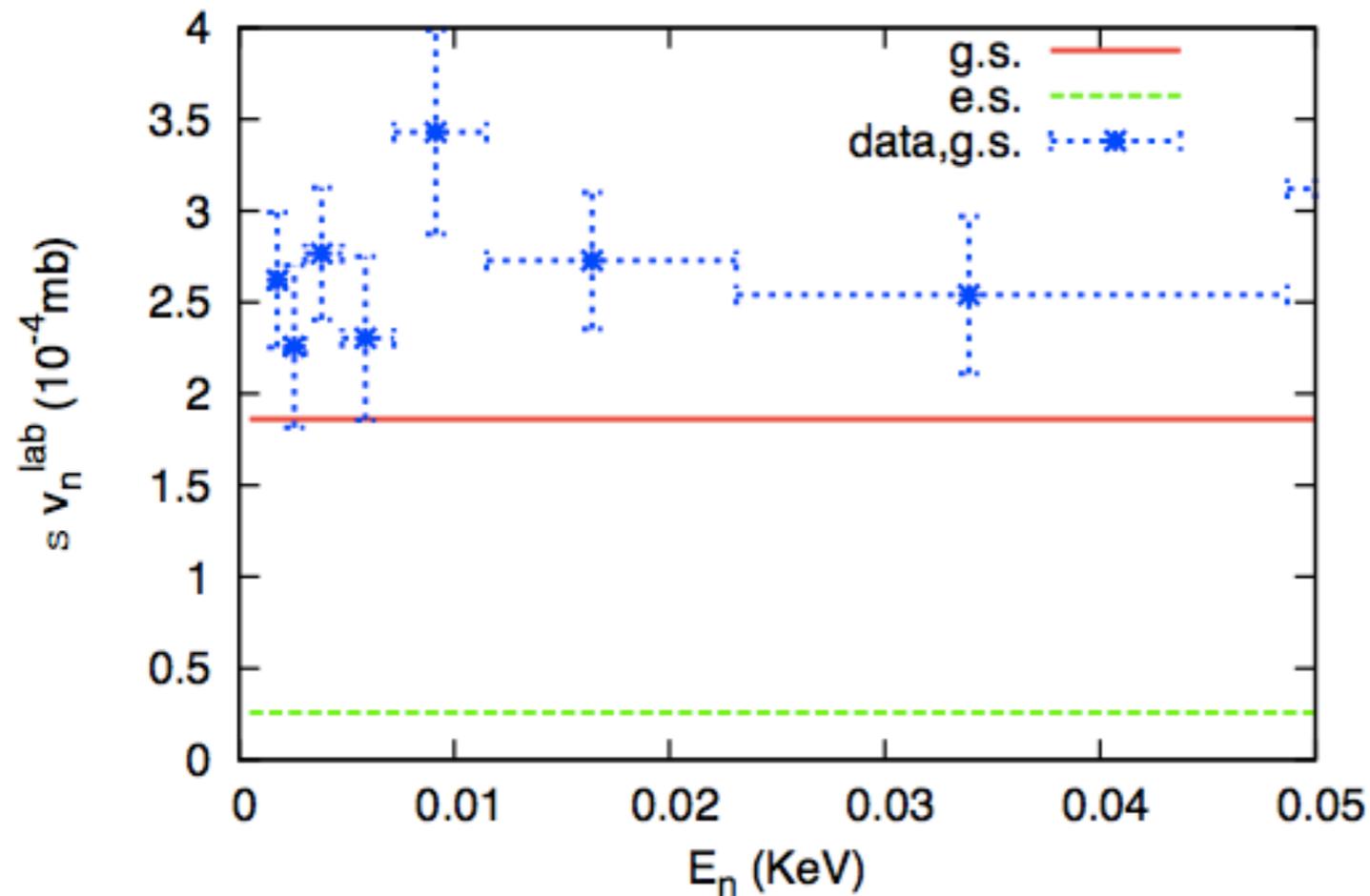
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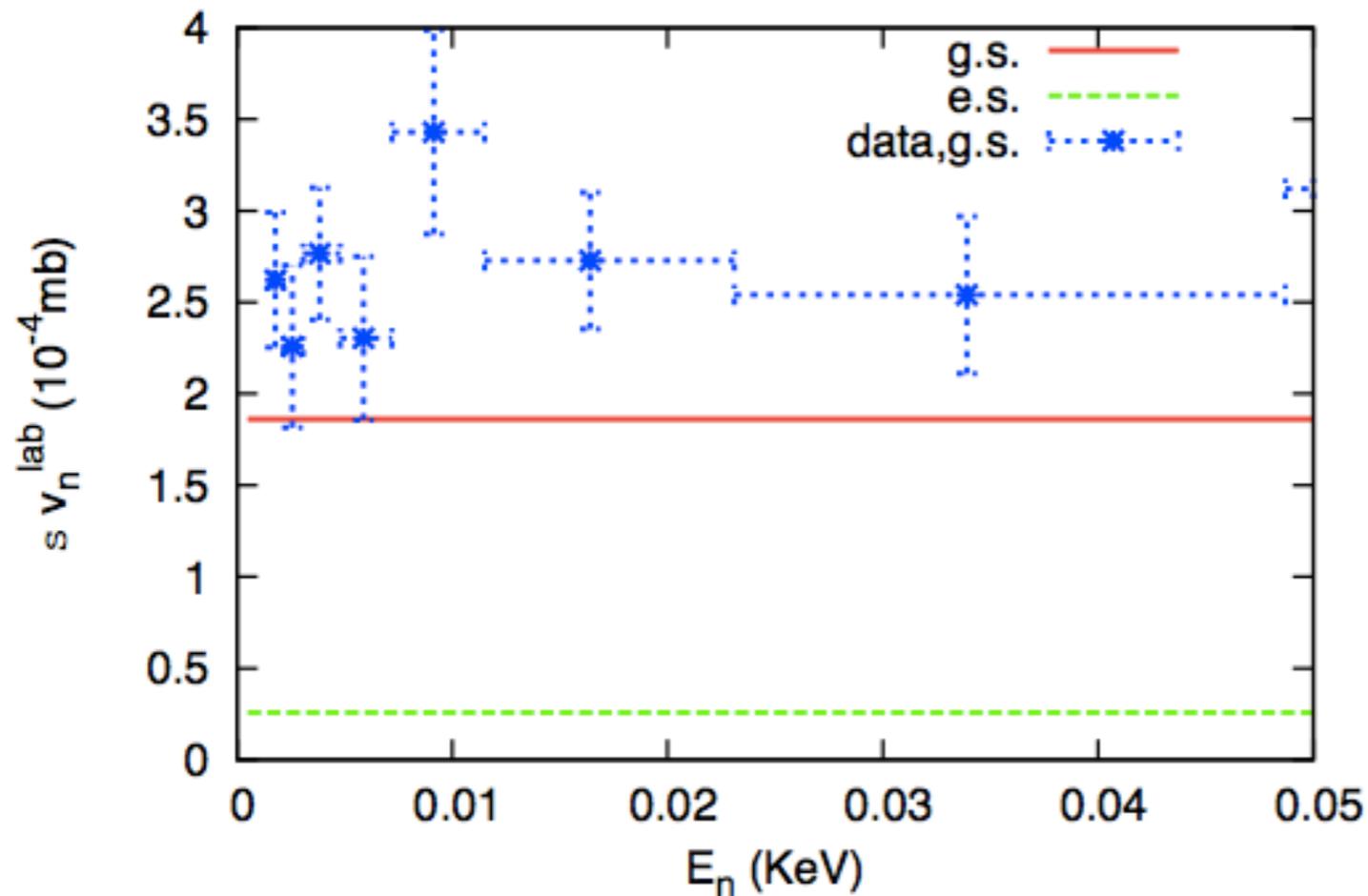
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Experiment  $> 0.86$

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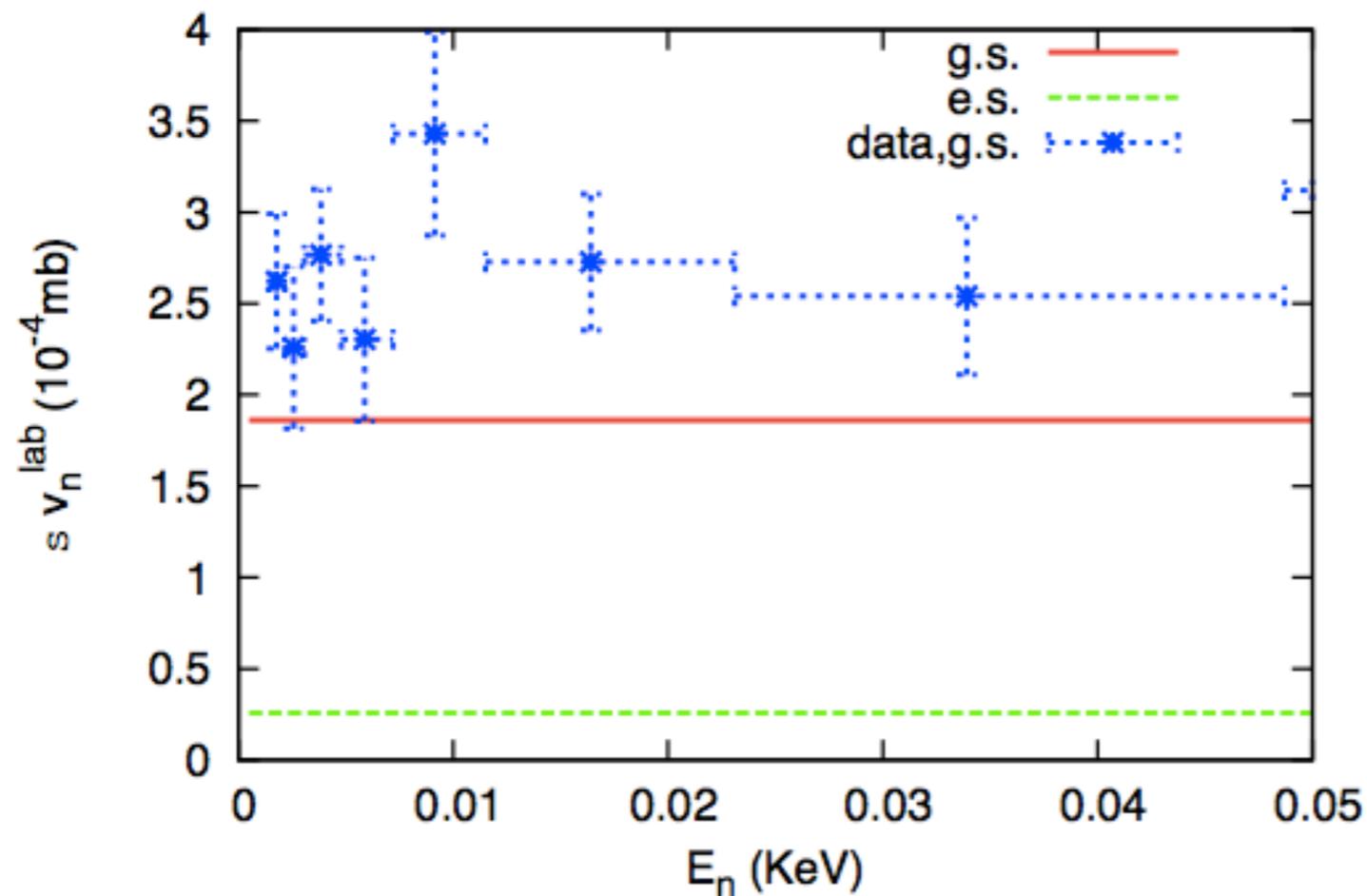
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Dynamics **predicted** through *ab initio* input

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# Data situation

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- 42 data points for  $100 \text{ keV} < E_{c.m.} < 500 \text{ keV}$ 
    - Junghans (BE1 and BE3)
    - Fillipone
    - Baby
    - Hammache (1998 and 2001)
  - CMEs
    - 2.7% and 2.3%
    - 11.25%
    - 5%
    - 2.2% (1998)
  - Subtract MI resonance: negligible impact at 500 keV and below
  - Deal with CMEs by introducing five additional parameters,  $\xi_j$
-

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# Building the pdf

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- Bayes:

$$\text{pr}(\vec{g}, \{\xi_i\} | D; T; I) = \text{pr}(D | \vec{g}, \{\xi_i\}; T; I) \text{pr}(\vec{g}, \{\xi_i\} | I),$$

- First factor: likelihood

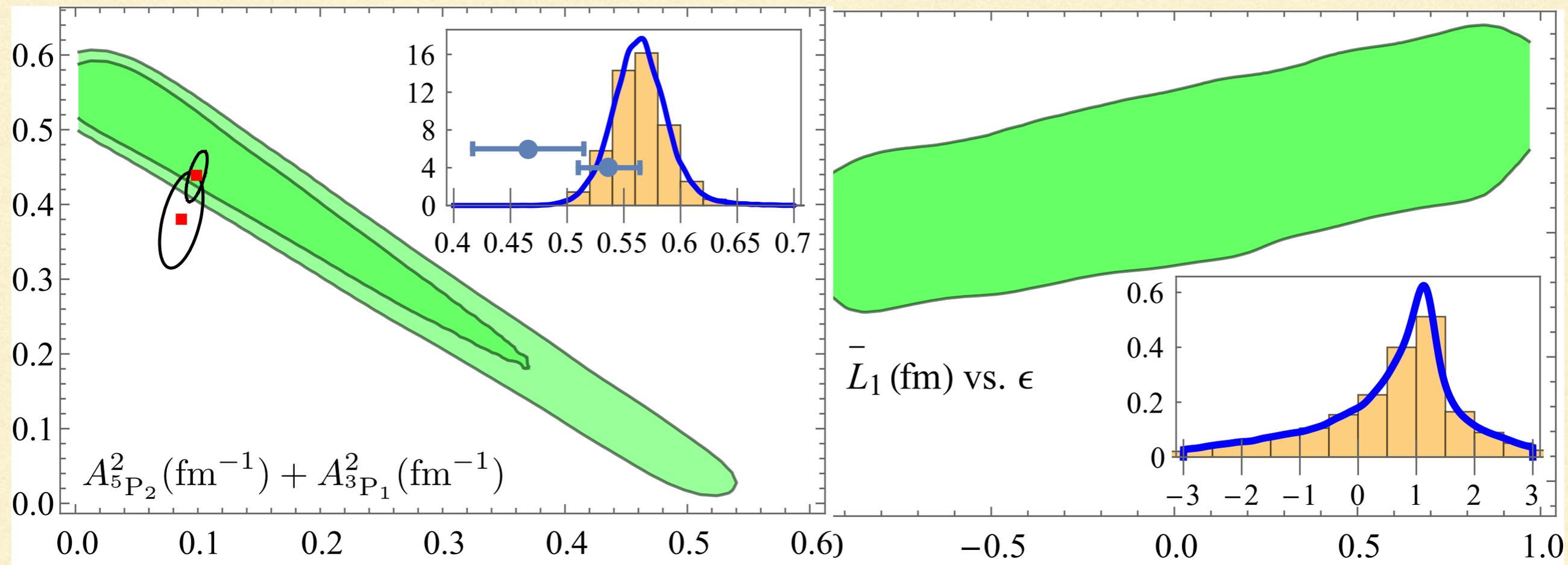
$$\ln \text{pr}(D | \vec{g}, \{\xi_i\}; T; I) = c - \sum_{j=1}^N \frac{[(1 - \xi_j)S(\vec{g}; E_j) - D_j]^2}{2\sigma_j^2},$$

- Second factor: priors

- Independent gaussian priors for  $\xi_j$ , centered at zero and with width=CME
  - Gaussian priors for  $a_{s=1}$  and  $a_{s=2}$ , based on Angulo et al. measurement
  - All other EFT parameters assigned flat priors, corresponding to natural ranges
  - No s-wave resonance below 600 keV
-

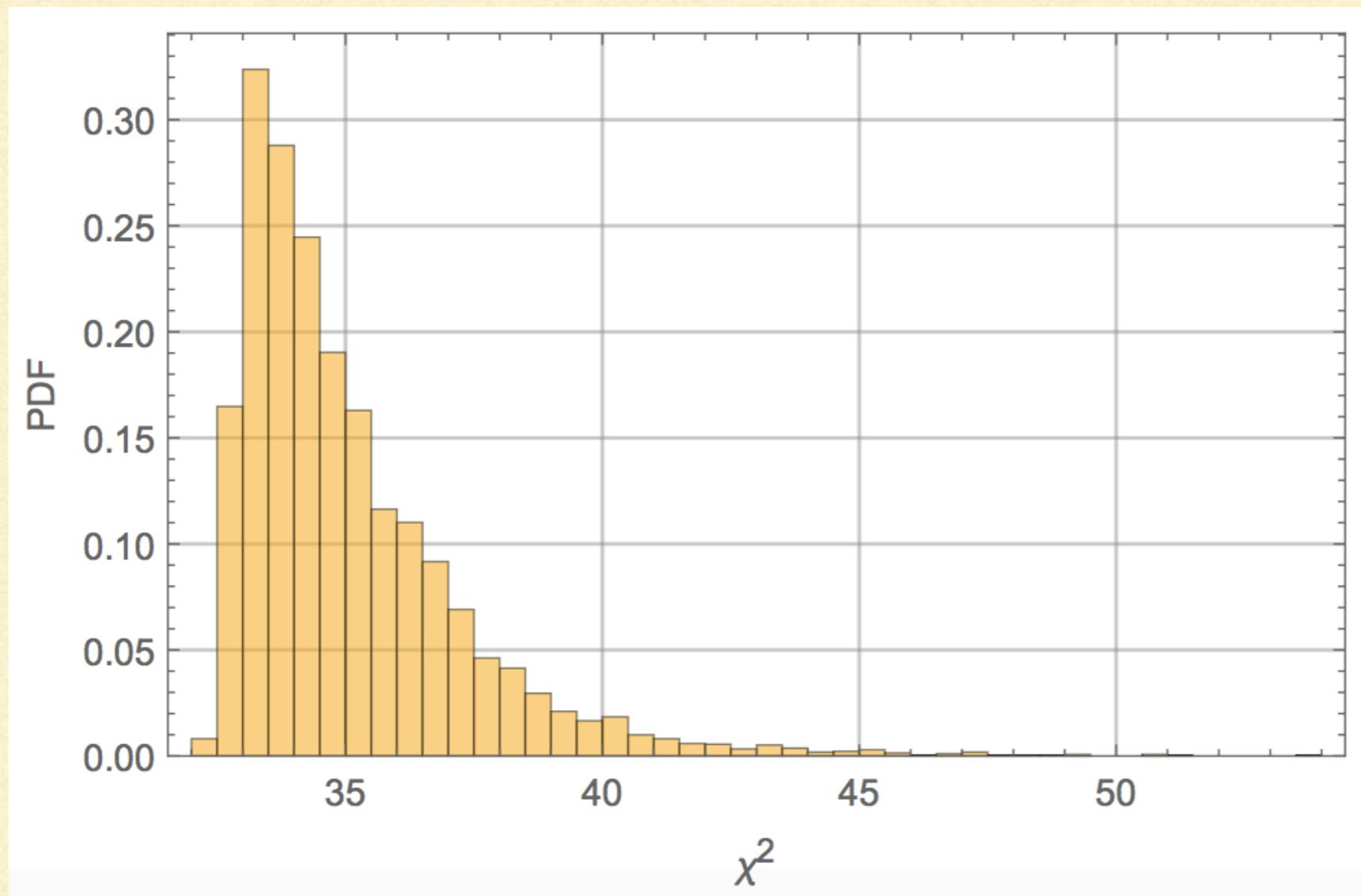
# Marginalizing $\rightarrow$ pdfs

$$\text{pr}(g_1, g_2 | D; T; I) = \int \text{pr}(\vec{g}, \{\xi_i\} | D; T; I) d\xi_1 \dots d\xi_5 dg_3 \dots dg_9$$



- ANCs are highly correlated but sum of squares strongly constrained
- One spin-1 short-distance parameter:  $0.33 \bar{L}_1 / (\text{fm}^{-1}) - \epsilon_1$

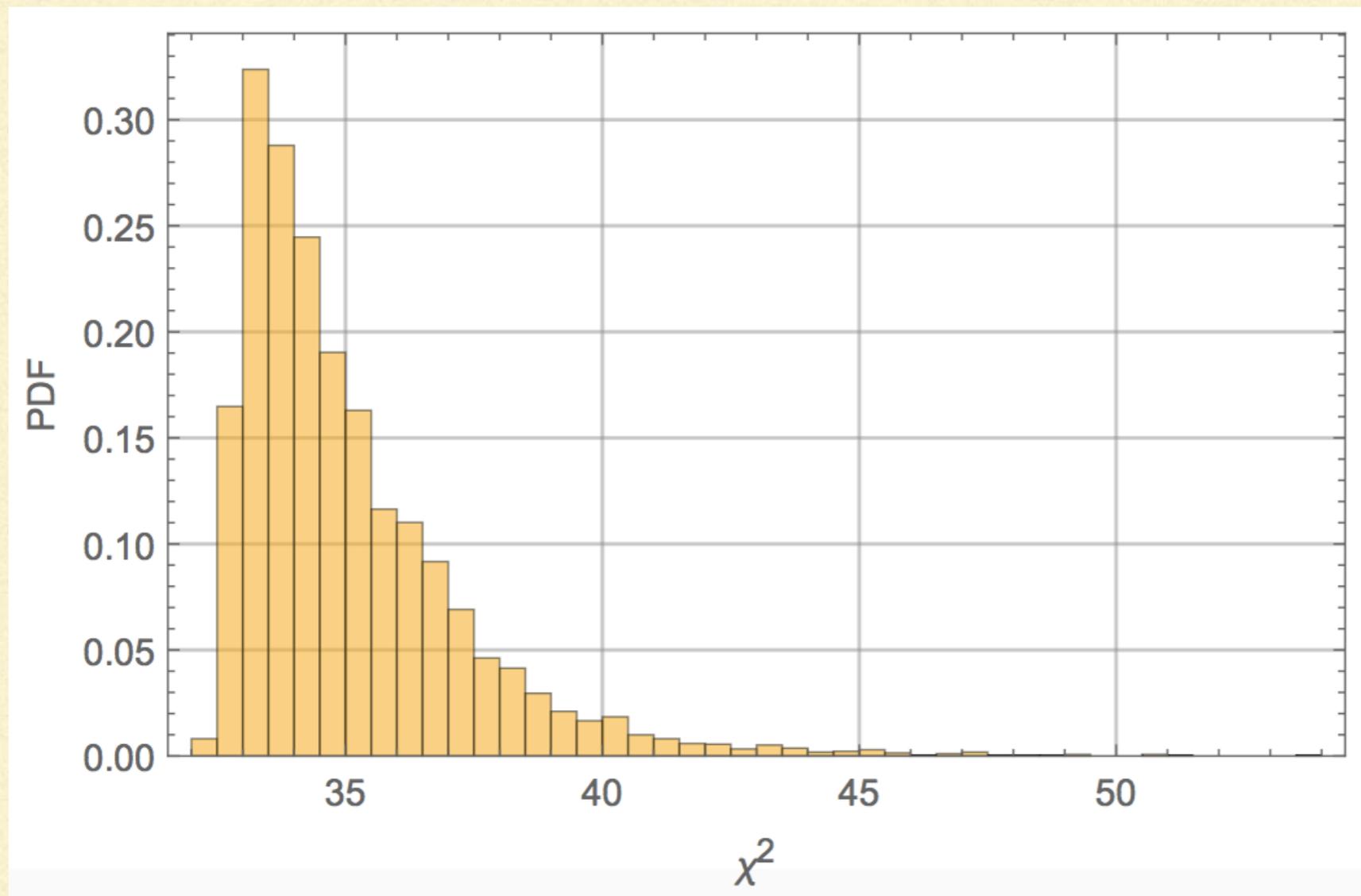
# More questions we can answer



42 data points,  
7 parameters “fit”  
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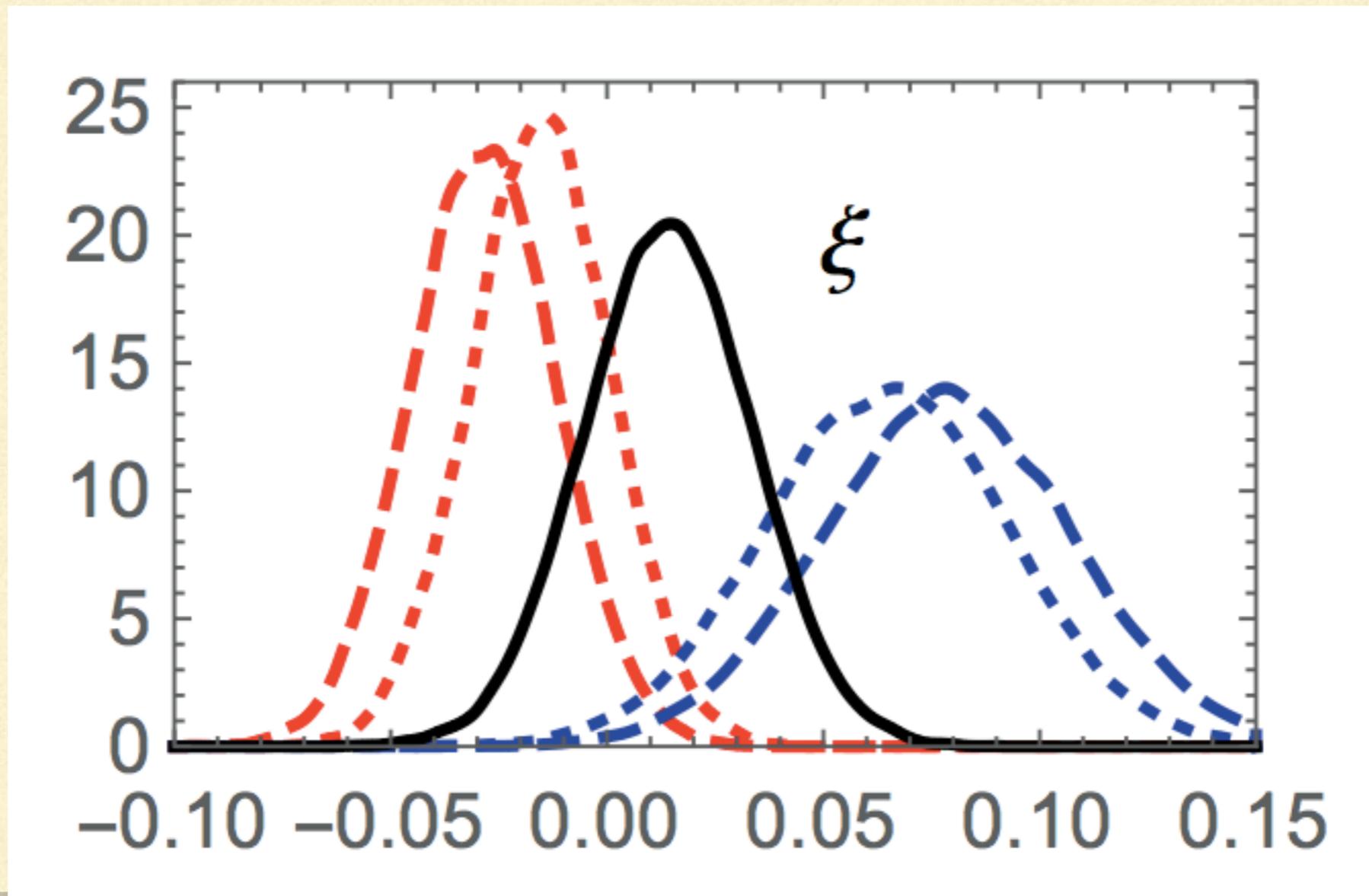
- Is it a “good fit”?



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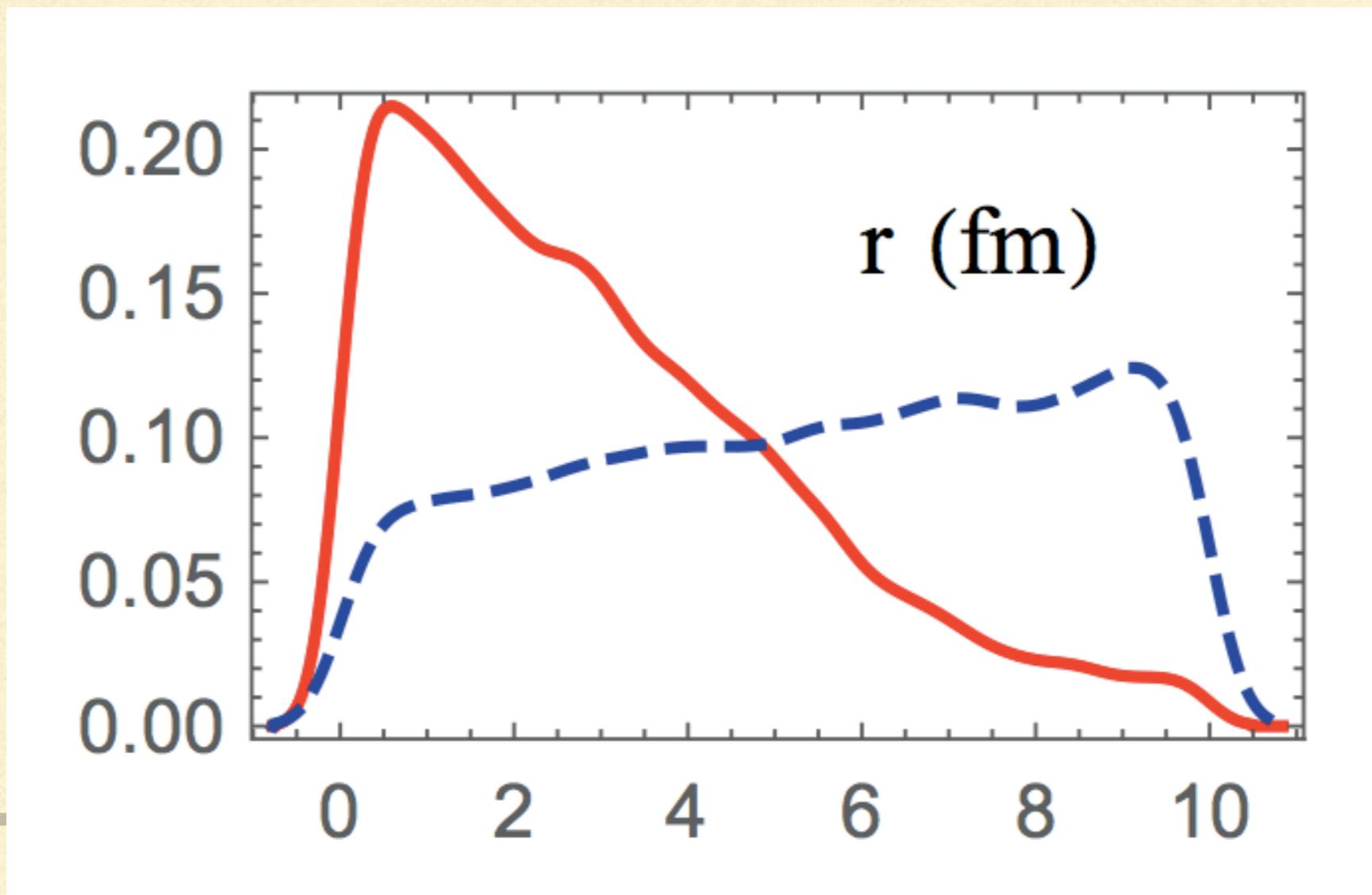
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- Is it a “good fit”?
- Did the experimentalists understand their systematic errors?



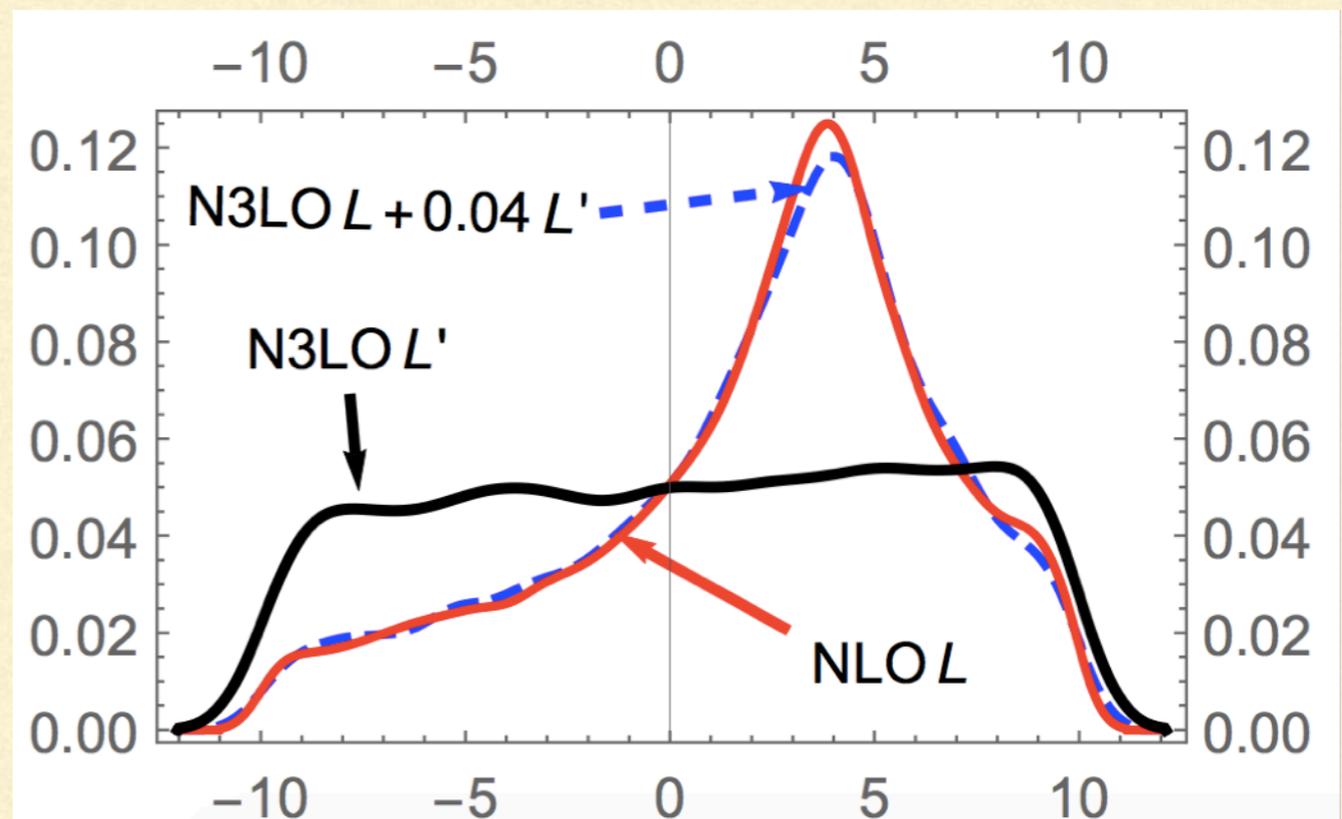
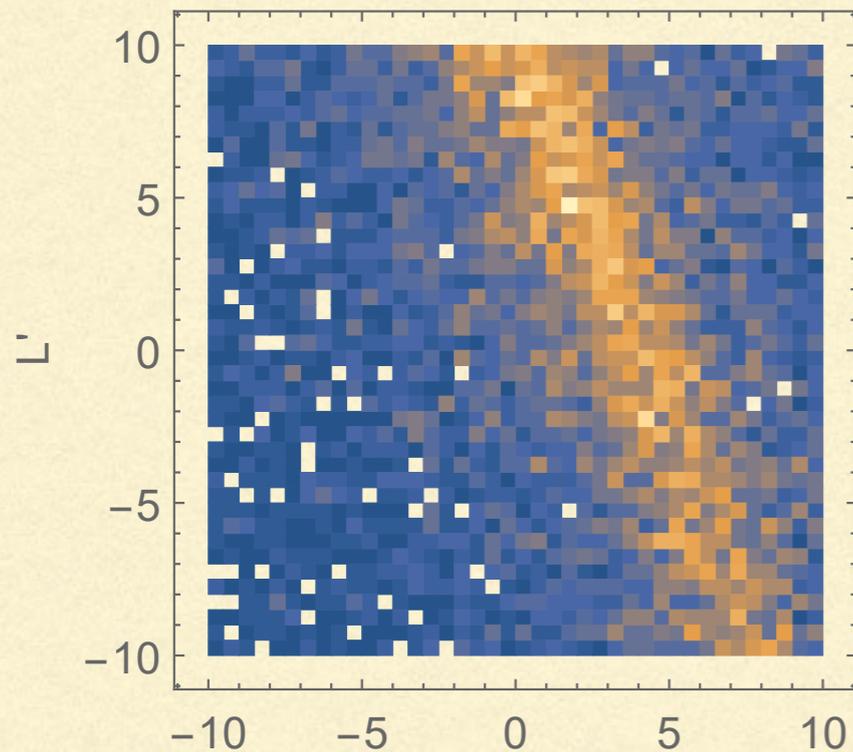
# More questions we can answer

- Is it a “good fit”?
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- Are there parameters that are not well constrained by these data?



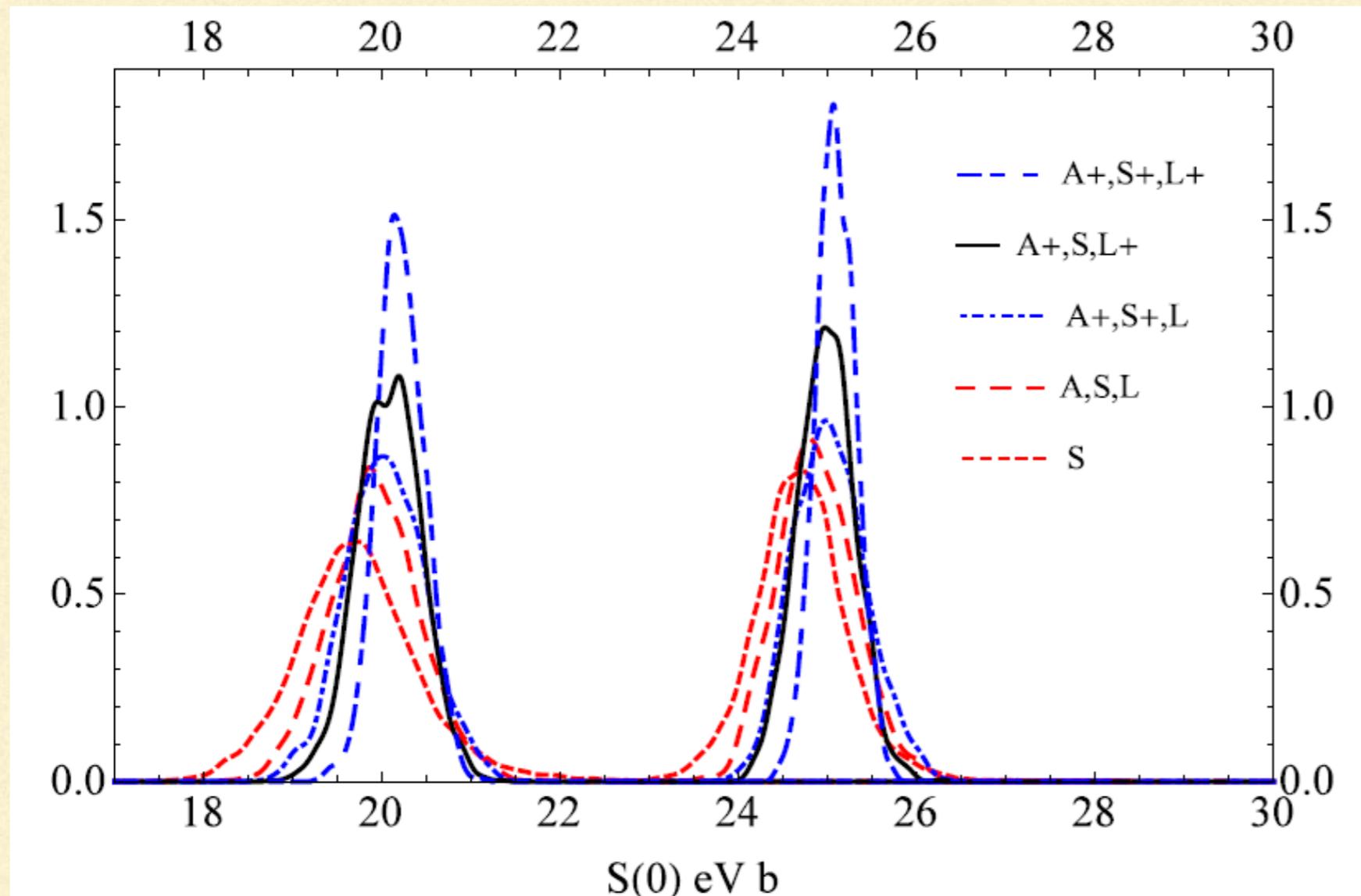
# Truncation error

- N2LO correction=0 (technically only in absence of excited state)
- EFT s-wave scattering corrections (shape parameter)~0.8%
- E2, M1 contributions < 0.01%, Radiative corrections: ~0.1%
- So first correction is at N3LO, i.e.,  $\bar{L}_i \rightarrow \bar{L}_i + k^2 \bar{L}'_i$



# Planning improvements

Use extrapolant to simulate impact of hypothetical future data that could inform posterior pdf for  $S(0)$



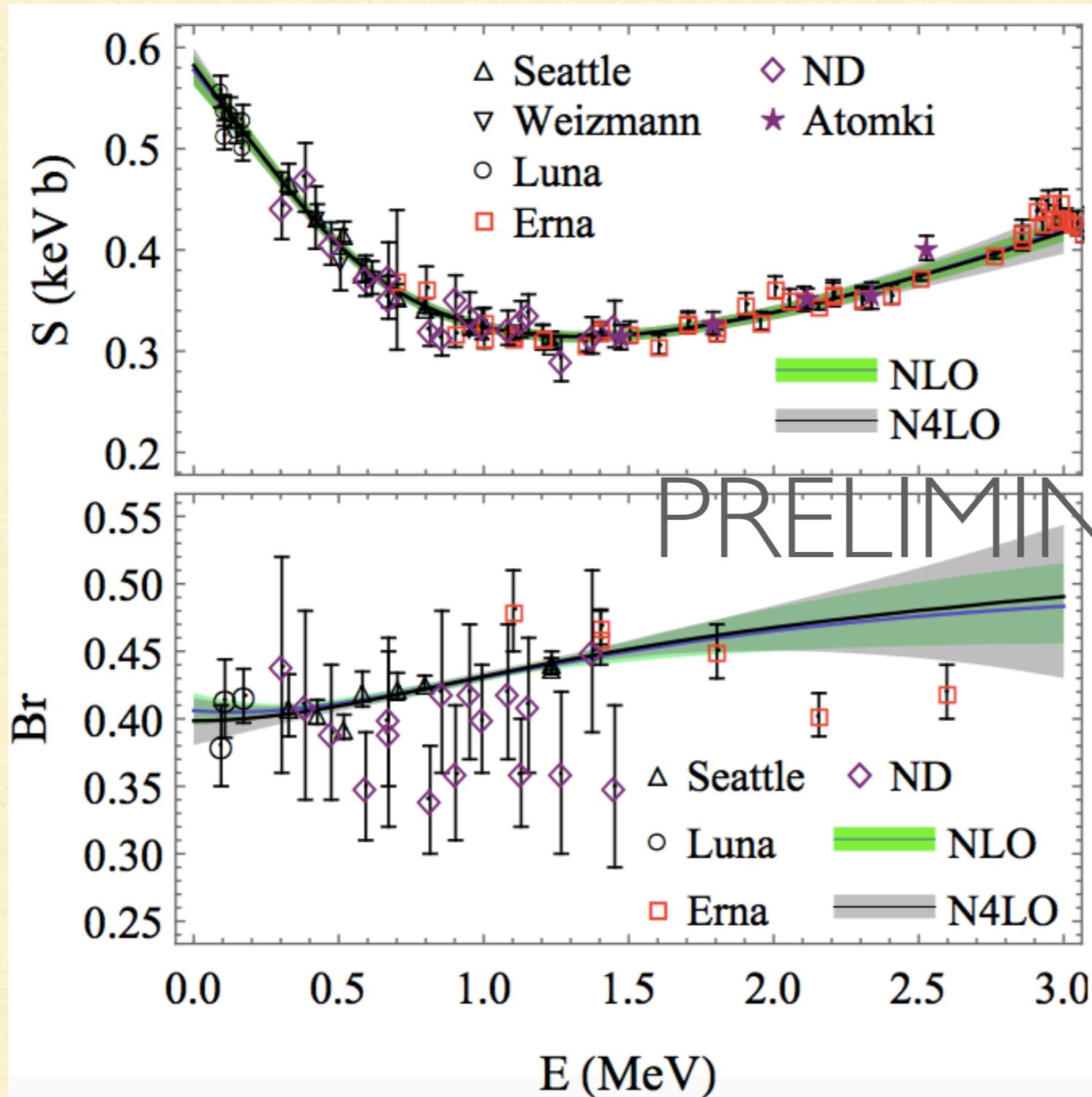
Left-to-right:  
42 data points all of  
similar quality  
to Junghans et al.

A: ANC  
S:  $a_{S=1}$  and  $a_{S=2}$   
L: short-distance

Note that 1 keV uncertainty in  $S_{Ip}$  of  ${}^8\text{B}$  may not be negligible effect

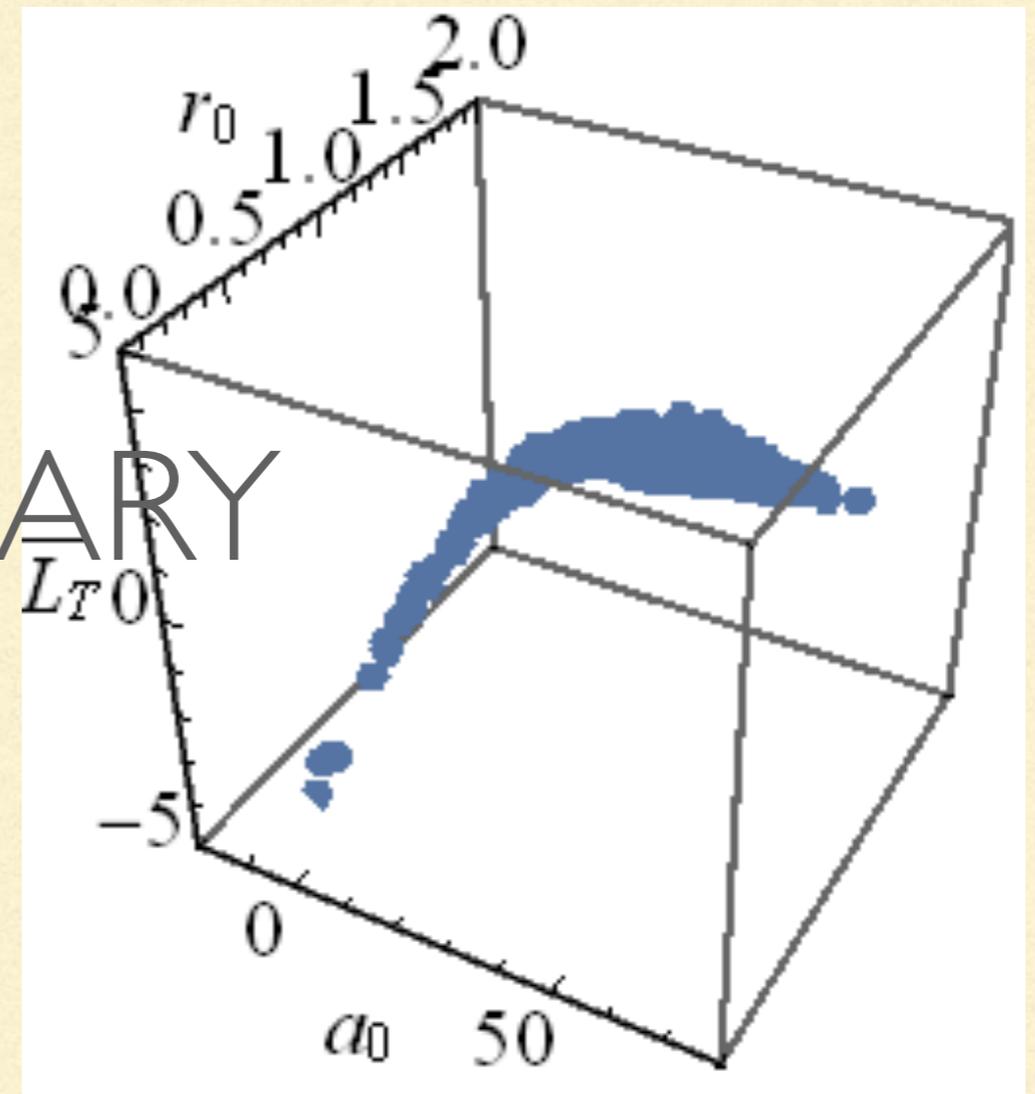
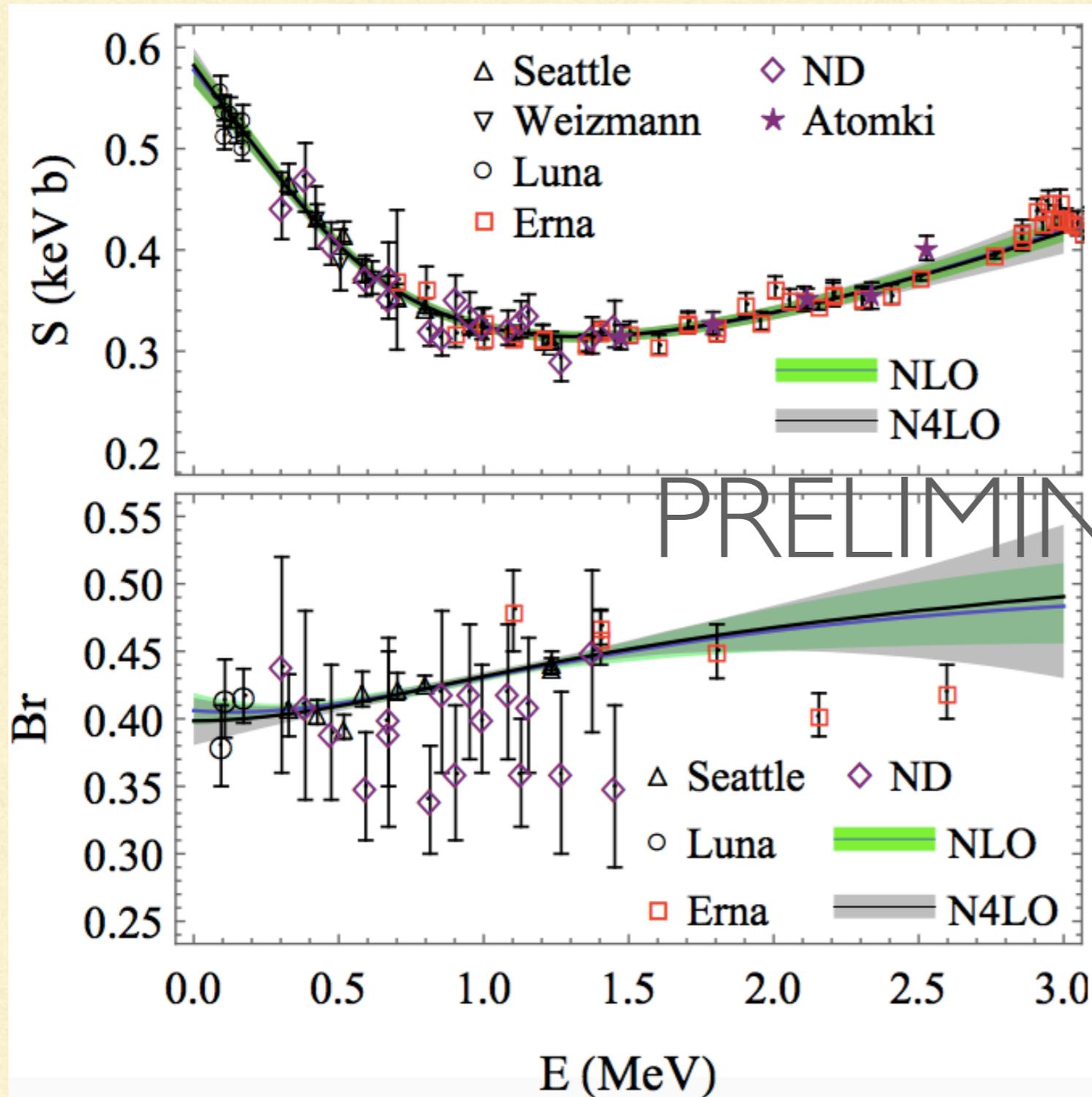
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Zhang, Nollett, DP, in preparation



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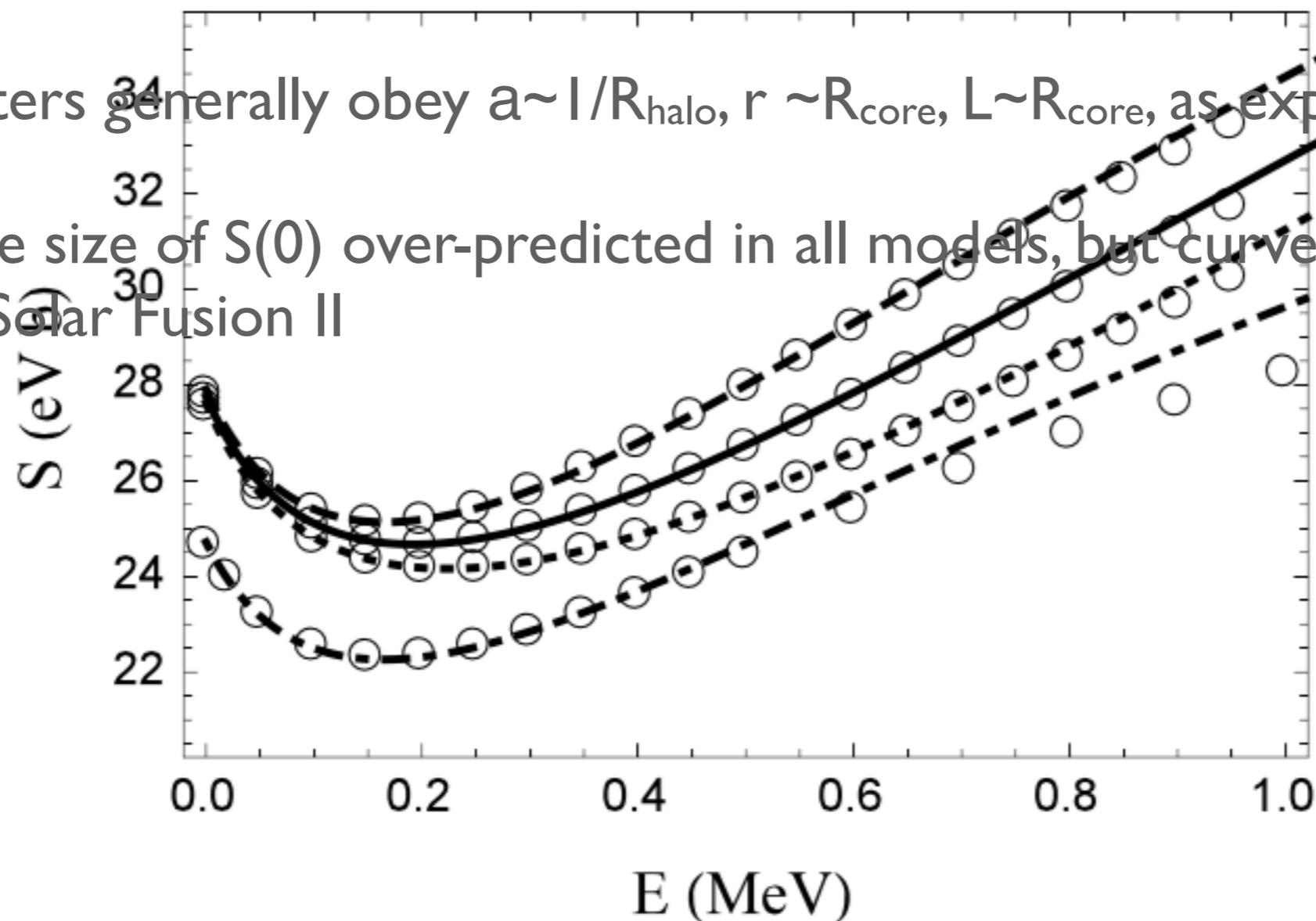


# Halo EFT as a “super model”

- Halo EFT is also the EFT of all the models used to extrapolate the cross section in “Solar Fusion II”
- Differences are sub-% level between 0 and 0.5 MeV

- Parameters generally obey  $a \sim 1/R_{\text{halo}}$ ,  $r \sim R_{\text{core}}$ ,  $L \sim R_{\text{core}}$ , as expected

- Absolute size of  $S(0)$  over-predicted in all models, but curves rescaled in fits for Solar Fusion II



$C_{2,2}^{(3P_2)}$   
0.20068  
0.20066  
0.20065  
0.10900

$L_2$   
2.68987  
3.10464  
4.18777  
3.73317

TABLE  
for scat  
assumed

is  $\text{fm}^{-1}$ ,  
explicitly